

# Magnetic loops: A comparison of extrapolations from the photosphere with chromospheric measurements

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- Coronal magnetic field models
- Potential and linear force-free fields
- Non-linear force-free fields
- Observational tests
- Conclusions

# Coronal magnetic field models

| <b>Model</b>     | <b>Mathematics</b>                            | <b>Observations needed</b> | <b>Validity</b>                          |
|------------------|---|----------------------------|--|
| Potential Fields | $\nabla \times B = 0$<br>$\nabla \cdot B = 0$ | Line of sight magnetogram  | (Global) current free regions, quiet sun |

# Potential and Linear Force-Free Fields

We have to solve the equations:

$$(\nabla \times \mathbf{B}) = \alpha \mathbf{B}$$

$$\nabla \cdot \mathbf{B} = 0$$

# Non-linear force-free fields

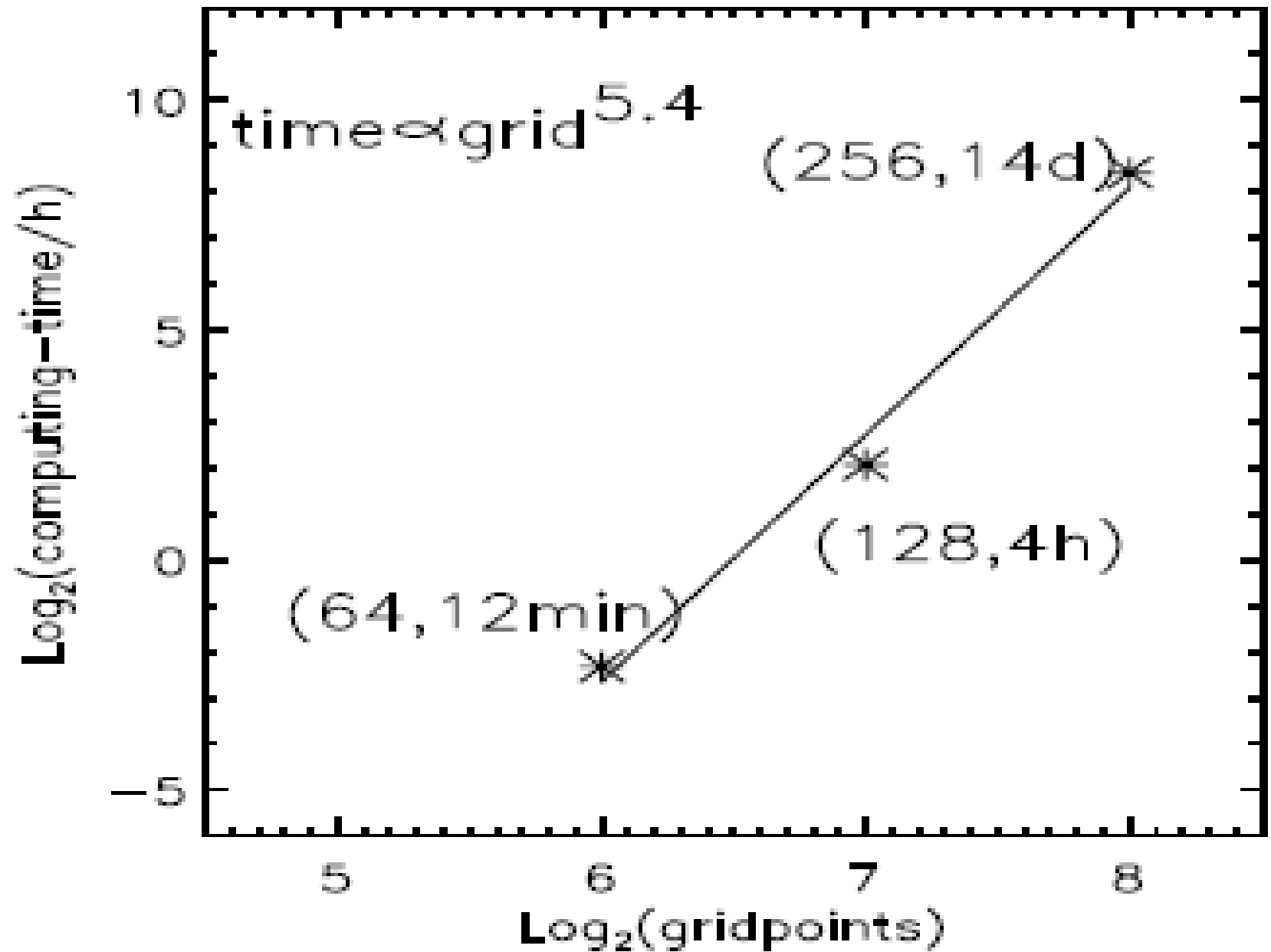
- Why do we need non-linear force-free fields?
  - In general  $\alpha$  changes in space.
  - Potential and linear force-free fields have no free energy to be released during an eruption.

# Non-linear Force-Free Fields

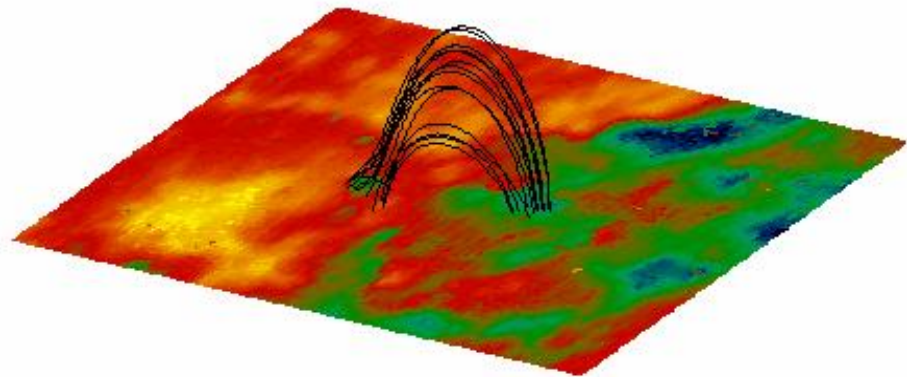
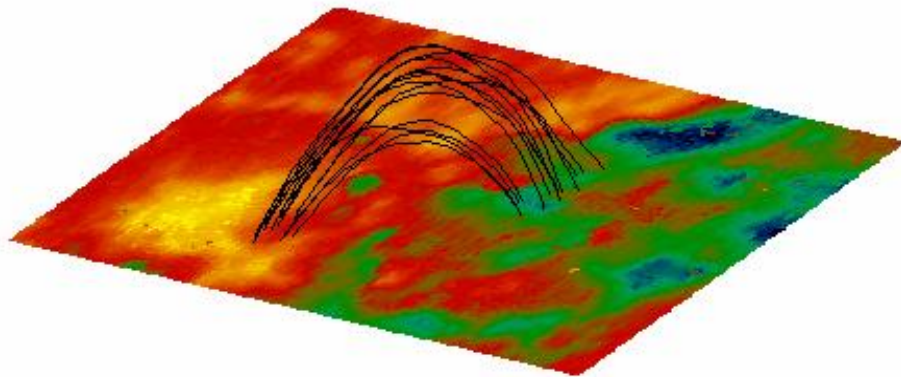
Force-free magnetic fields have to obey

$$(\nabla \times \mathbf{B}) \times \mathbf{B} = \mathbf{0}, \quad \nabla \cdot \mathbf{B} = 0$$

# Cubic-box, 9 Procs

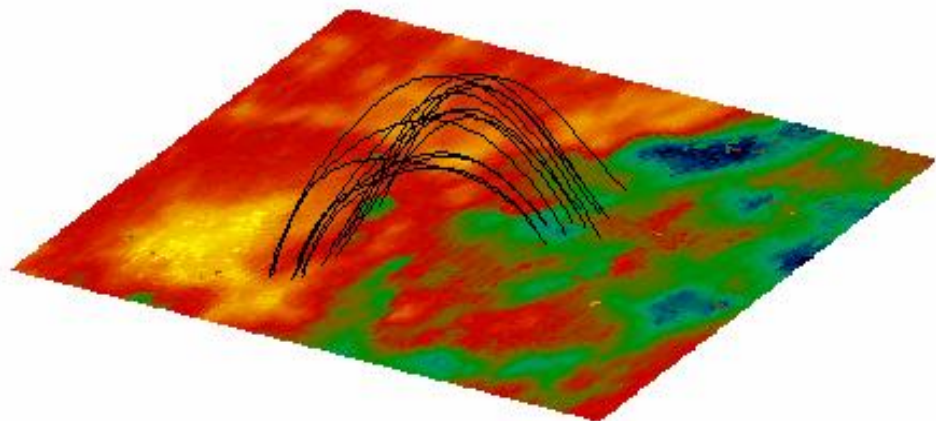
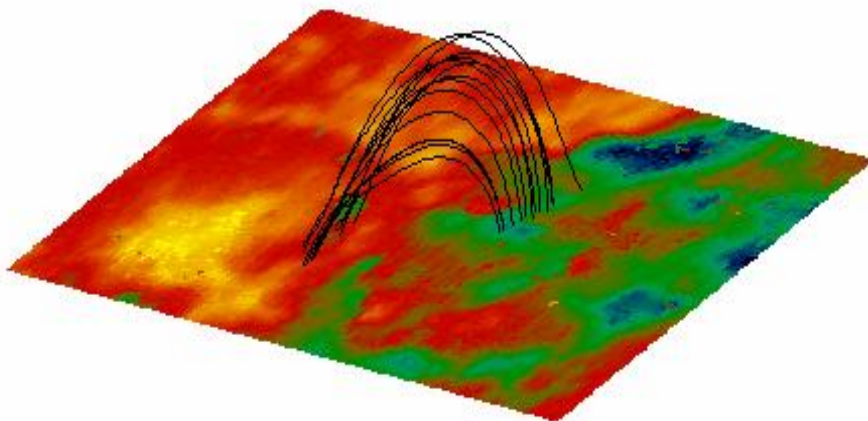


# Comparison of observed magnetic loops and extrapolations from photospheric measurements



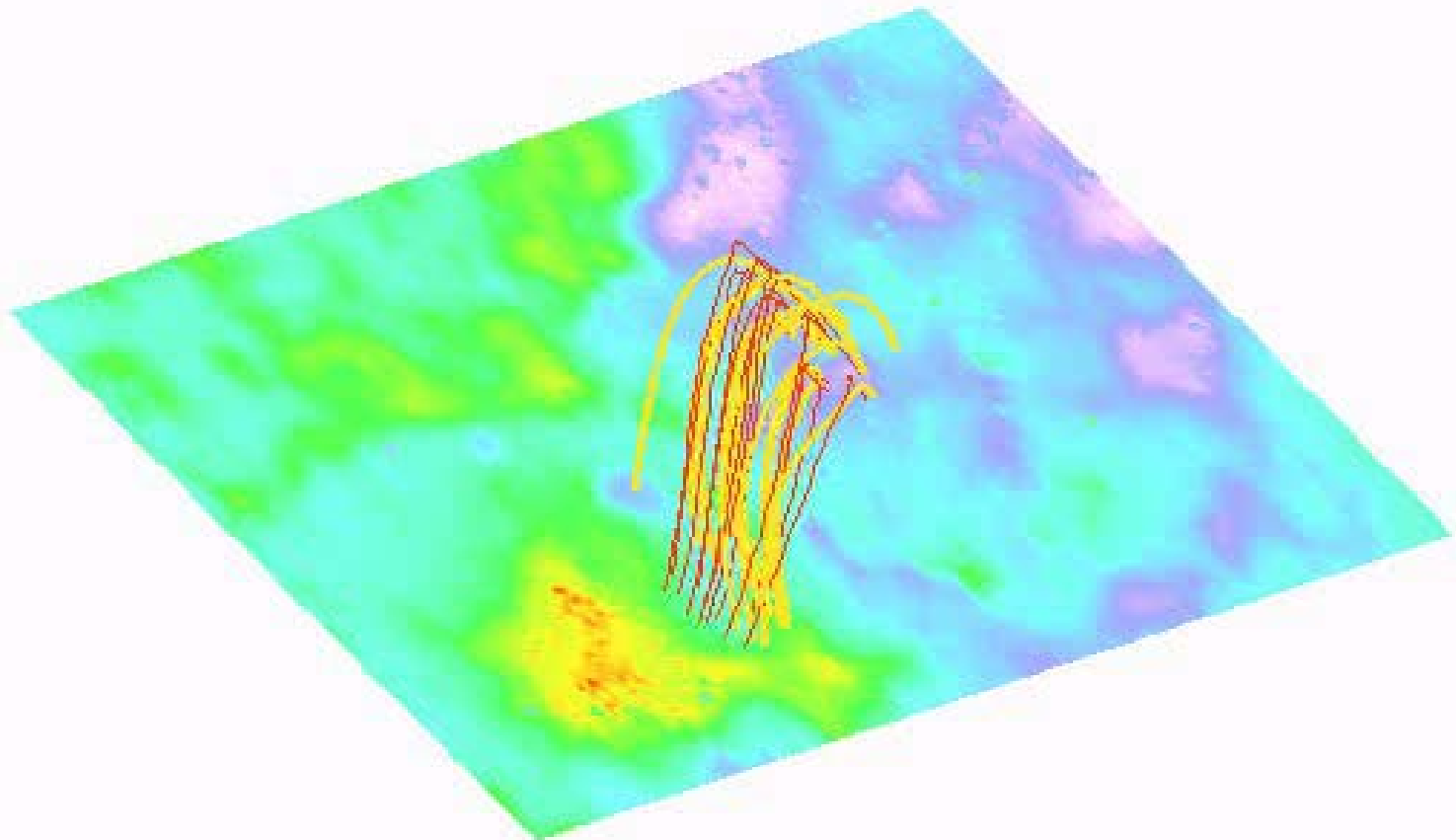
Measured loops in a newly developed AR (Solanki, Lagg, Woch, Krupp, Collados, Nature 2003)

Potential field reconstruction

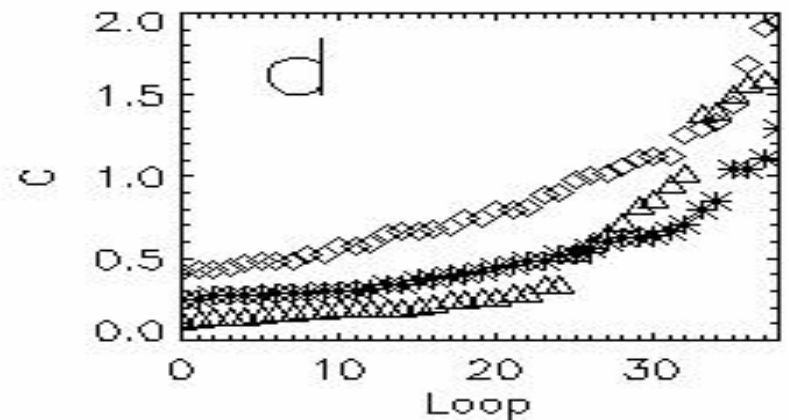
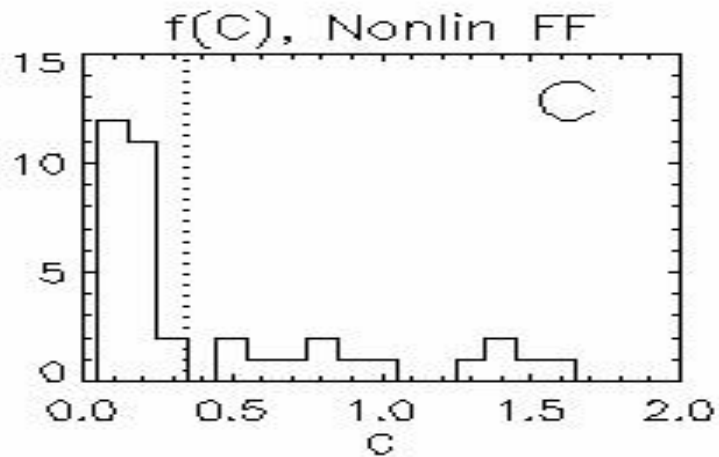
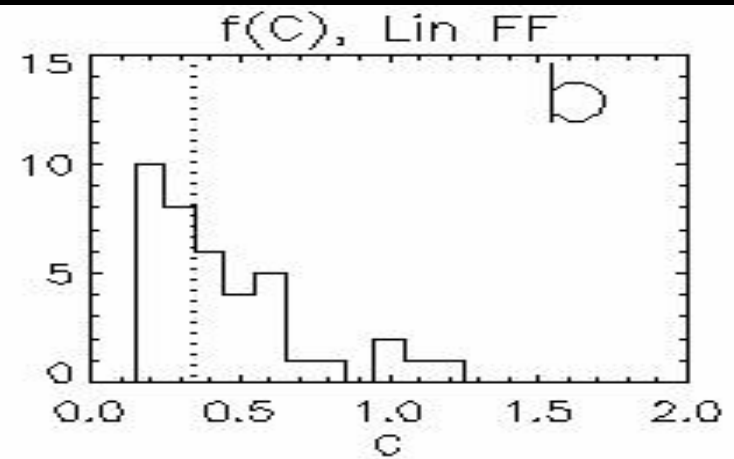
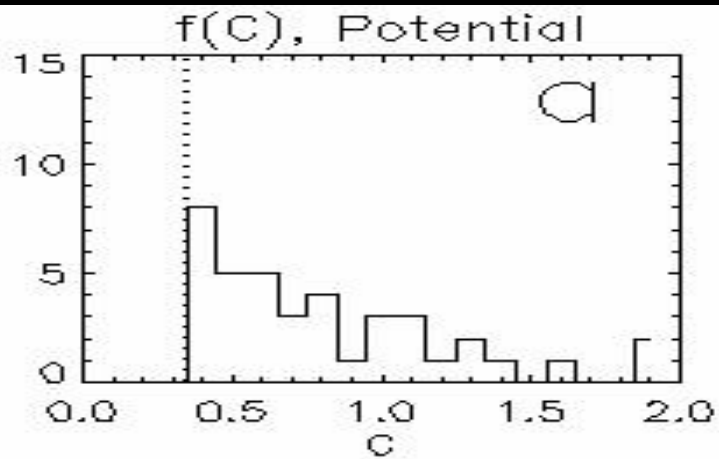


Linear force-free reconstruction

Non-linear force-free reconstruction







We compared measurements of magnetic loops in a newly developed active region with extrapolations from the photosphere. We got the best agreement of measured and extrapolated loops for a non-linear force-free magnetic field model.

# Conclusions

- Potential magnetic fields and linear force-free fields are popular due to their mathematic simplicity and available data. (e.g. from MDI on SOHO, Kitt Peak)
- Non-linear force-free fields are necessary to describe active regions exactly. More challenging both observational (Vector magnetograms) and mathematical.
- Current vectormagnetograms have limited field of view and only occasional available. Future: Solis and Solar B.