

# ESA's exoplanetary program

## The Darwin Mission

Status & recent progress



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# Outline:

- 1. ESA Cosmic vision program**
- 2. Scientific rationale**
- 3. Status of the Darwin study**
- 4. Current stepwise approach**
- 5. Possible role for small missions**
- 6. Where do we go from here?**

# Cosmic Vision 2015-2025

## Current status

- Implementation of current H2000+ program in preparation:
  - GAIA: 2011/12
  - Bepi-Colombo: 2014
  - LISA/Pathfinder ?
  - Solar orbiter under review
- CV1525 addresses four grand themes
- **Call for proposals 2006**
- **Call for one mission within small financial constraint for 2015-16**
- + one larger mission without formal constraint flight in ~2020 (Darwin?)
- 3 candidates for either will be selected for (further) study and will receive funding for technology development

# Cosmic Vision is centered around four Grand Themes:

## Cosmic Vision

Space Science for Europe 2015-2025



Российская Космическая Агенция  
Агентство космических исследований

1. What are the conditions for planet formation and the emergence of life?
  - From gas and dust to stars and planets
  - **From exo-planets to biomarkers**
  - Life and habitability in the Solar System
2. How does the Solar system work?
3. What are the Fundamental Physical Laws of the Universe?
4. How did the Universe originate and what is it made of?

### Proposed strategy:

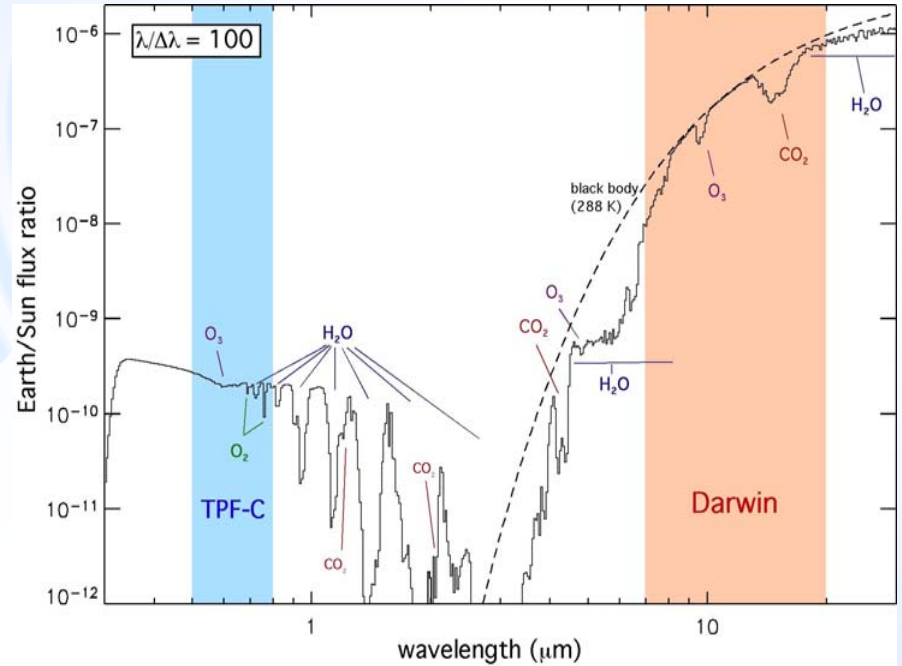
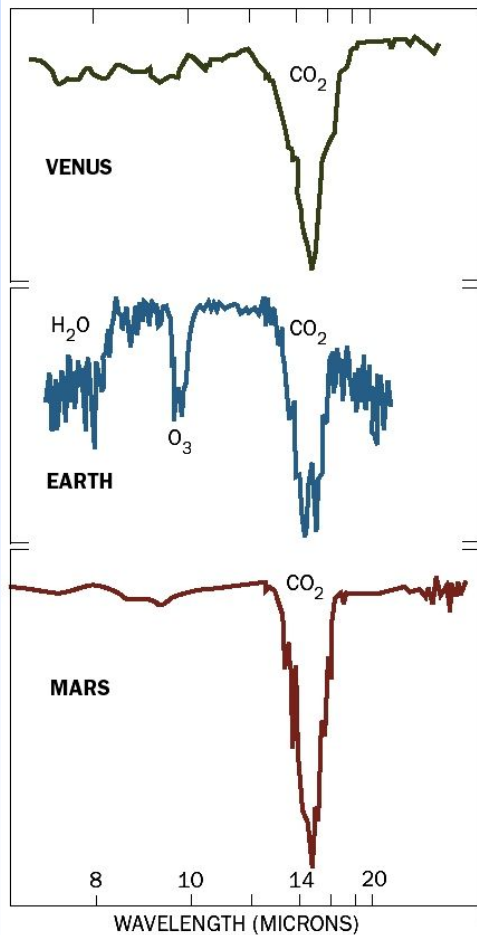
- First: In-depth analysis of terrestrial planets (Darwin / NIRI)
- Next: Understand the conditions for star, planet and life formation (Far IR observatory / Solar Polar Orbiter)
- Later: Census of Earth-sized planets & explore Europa (Terrestrial Planet Astrometric Surveyor / Europa orbiter / lander)
- Finally: Image terrestrial exo-planet (beyond 2025) (Large Optical Interferometer)



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# Bio markers



The goal is to find rocky planets with an atmosphere out of photo-chemical balance.

(cf. Lederberg, 1965 and Lovelock, 1965)

**Working definition:**

**Life  $\equiv$  O<sub>3</sub> + H<sub>2</sub>O + CO<sub>2</sub>**

Ideally, detect reduced molecules (CH<sub>4</sub>, N<sub>2</sub>O) as well

# The quest for Terrestrial exoplanets

Significantly smaller than currently detected planets but we are getting there

- No doubt that microlensing will start to pick up them
- RV requires 0.1 m/s (possible) over 3 years (1AU/G2V) with large telescope and no disturbing P-modes (6 month time scale 0.5-1 m/s) This is for the Habitable Zone. Real problem is stellar activity

Direct observation with next generation ELT's require:

- 100m class telescopes! Can only work on G-stars to 18pc! (35 targets)

**If telescope surface Perfect!!**

- Bright star ==> contrast ratios of  $10^{10}$ - $10^{11}$  at subarcsec separation!
  - Diffraction control, wavefront quality (Strehl ratio), wavefront stability → 2-3 orders of magnitude above today

**If we need to observe directly: Go To Space!!**

# Two extreme cases are possible and need to be considered:

- Earth-like planets are common
- Earth-like planets are extremely scarce

This means any search need to be designed such that a negative result is also meaningful

Therefore, the DARWIN mission is designed to search and characterize Earth-like planets around at least 300 – 500 stars



# In-depth analysis of Earth-like planets

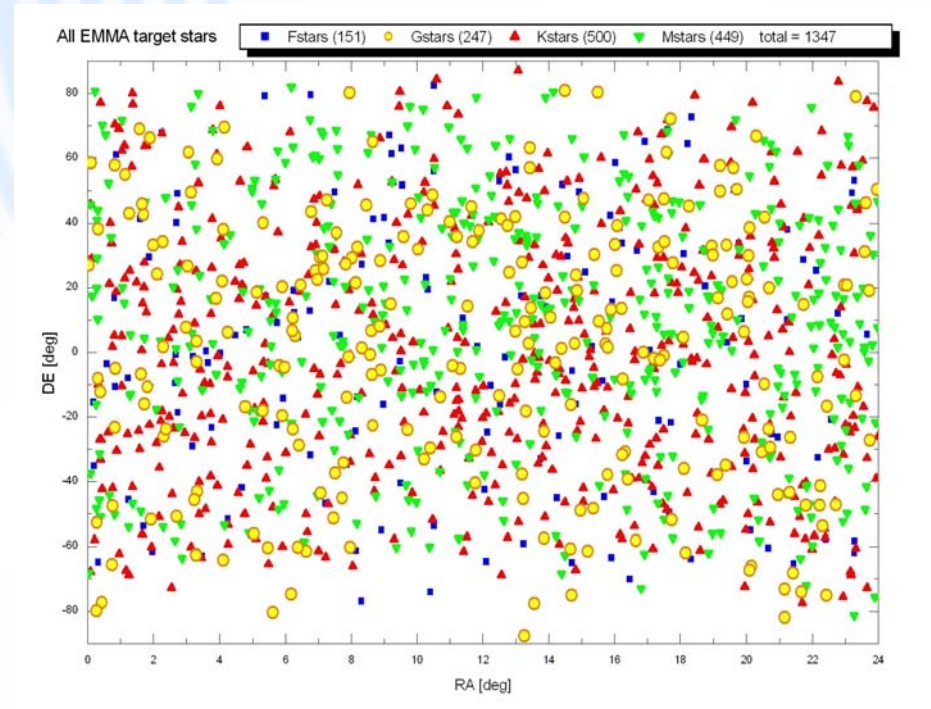
## What strategy?

- Search for rocky planets and determine if they are common or rare
- Determine the physical conditions of and on these planets
- Determine whether these planets are in principle habitable
- Find out if these planets have life on them
- To place ourselves in context we need the science of Comparative Planetology

# DARWIN Target catalogue

Single stars out to 25 pc:

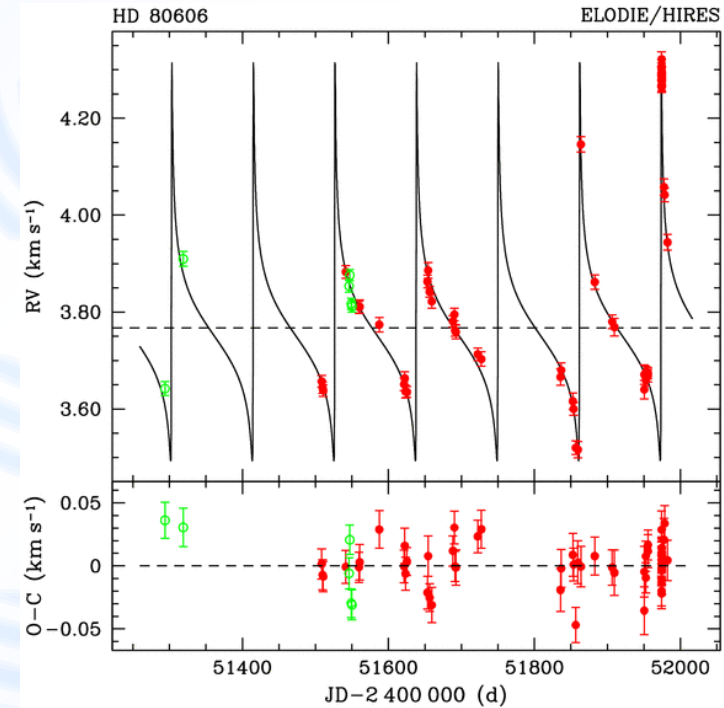
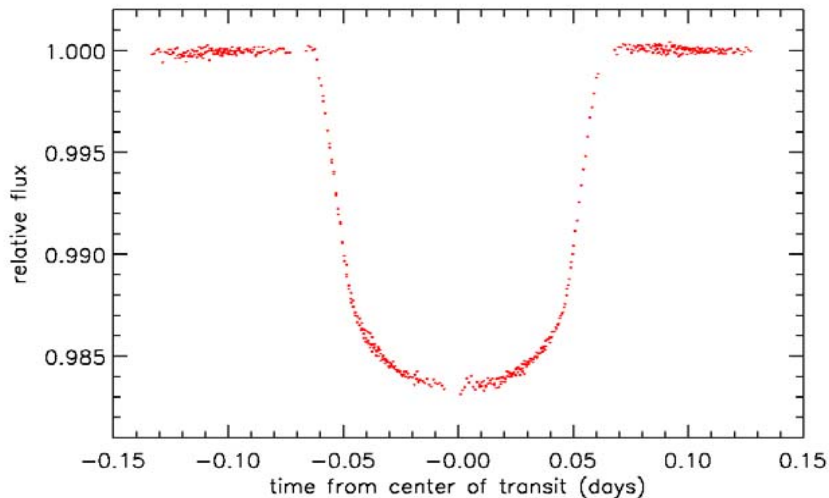
Spec. Type	Number
F	43
G	100
K	244 (incomplete)
M	241 (incomplete)
<b>Total:</b>	<b>608</b>



All missing late K and M stars within 25 pc will be found by GAIA in ~ 1.5 year of mission time because of large parallax and sensitivity (GAIA can observe down to  $R = 25$  magnitude)

# Exo-planets do exist in abundance...

- >200 planets in > 140 systems since 1995
- But most detected so far are Jupiter-size planets in tight, eccentric orbits
- Some detected with  $6 < M < 10 M_{\oplus}$  but either in short period orbits or around M-stars!



Naef, Latham et al., 2001  
 $e = 0.927$

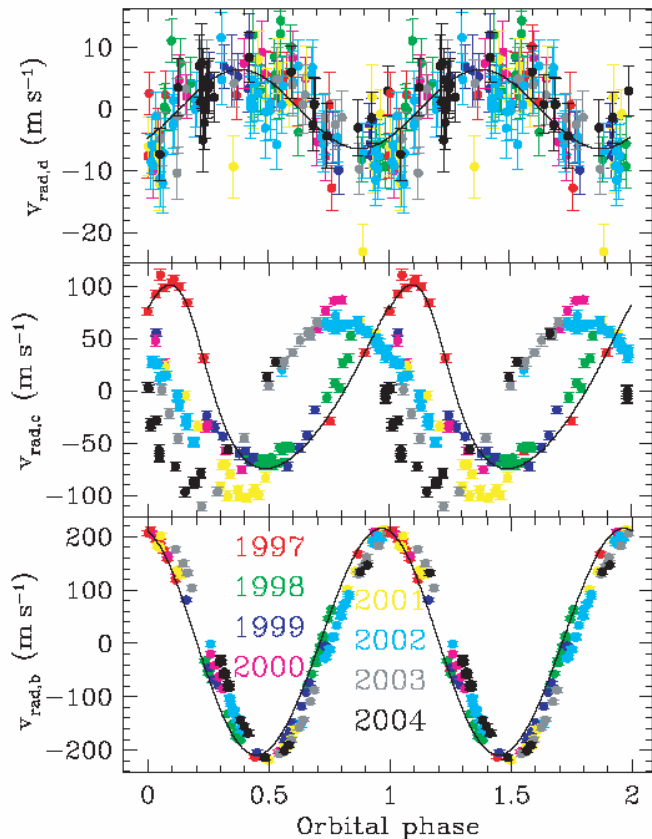
Charbonneau et al., 2000

# probably down to Earth-like masses...

## 7.5 $M_{\oplus}$ ....

at 0.021 AU around an dM4 GJ 876

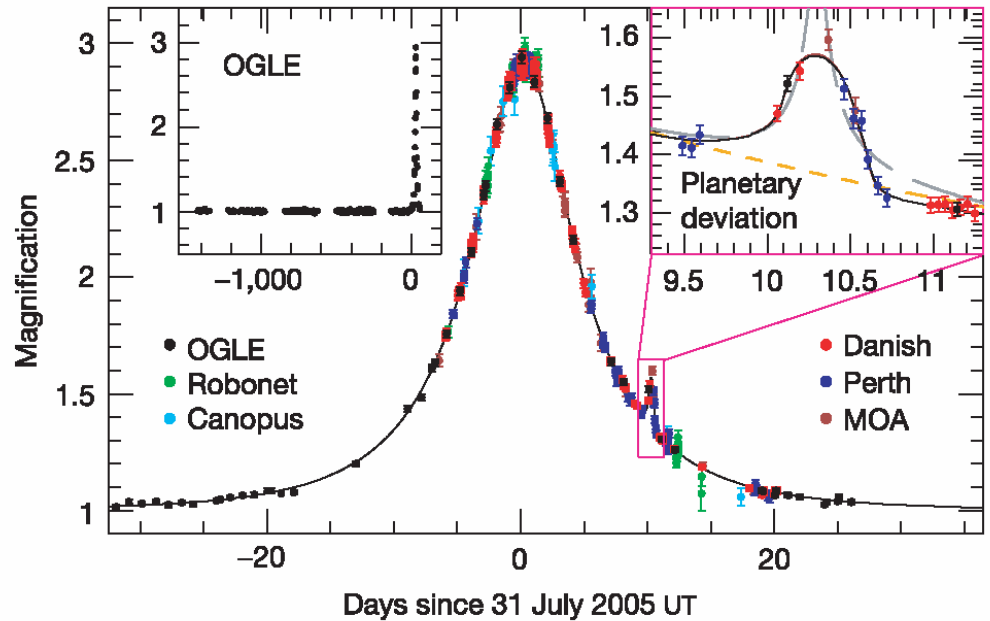
Rivera et al., ApJ 634, Nov. 2005



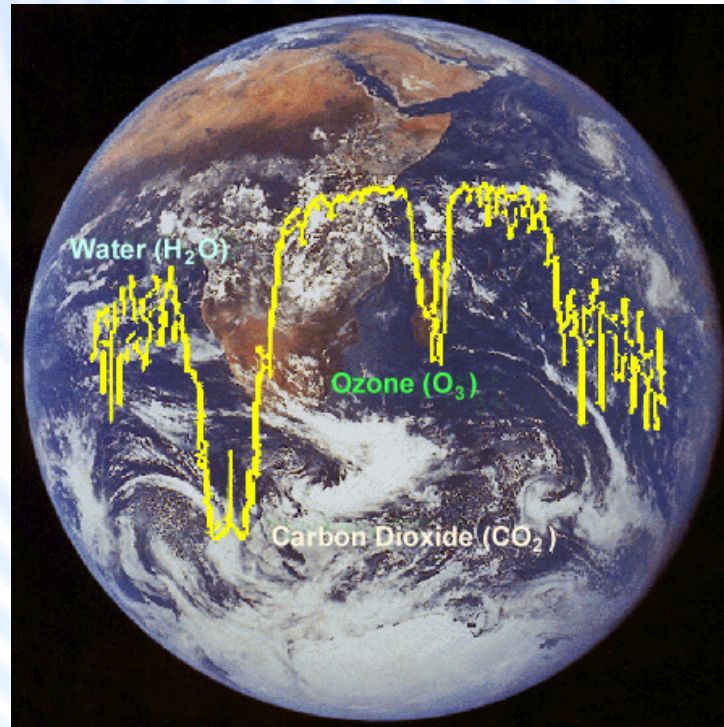
## 5.5 $M_{\oplus}$ ....

at 2.6 AU around an 0.22  $M_{\text{sun}}$  M-dwarf

Beaulieu et al., Nature, Jan. 2006

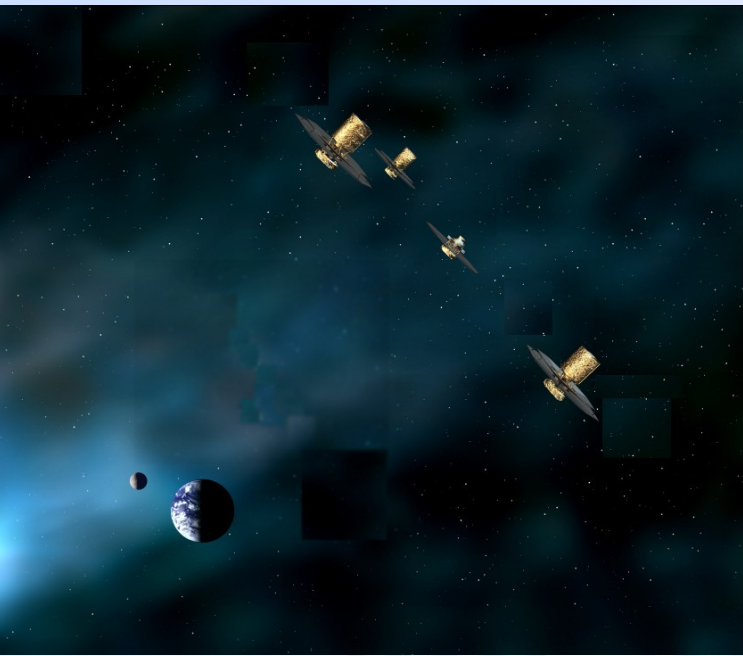


# ... but the 'Holy Grail' is still eluding us



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Darwin is a nulling interferometer connecting 3 - 4 telescopes, each on its own s/c with phase delays extinguishing the light on the optical axis, while enhancing the light a small distance (in the Habitable Zone for the star in question) away. Interferometry takes place in the beam combiner s/c for a total of 4-5 s/c. Mirrors in 4m class are required to reach all stars out to 25-30pc. Modulation (to alleviate noise factors) is introduced by using sub-interferometers (and 'chop' the signal) and by rotating the array, essentially producing a transmission pattern on the sky **'searching'** the HZ.

# Status of Darwin

- Data processing study by Alcatel Alenia (F) to finish October 2005
- End-to-end simulator development by EADS Astrium (D) + 18 months
- Two parallel system studies by Alcatel Alenia (F) and EADS Astrium (D) kicked-off in October 2005, finish in April 2007
  - Study Darwin in triangular or X-array formation (i.e. 3 or 4 tel.)
  - 3 – 4 m class telescopes
  - Spectral range: 6 – 20  $\mu\text{m}$
  - Spectral resolution 20 – 50 (120)
  - A5 launch to L2
  - 5 years nominal mission time + 5 years extended
- To be proposed by community for the Cosmic Vision as a major mission this Fall



# Status of the field

<b>GENIE:</b>	<b>(ESA) ground-based nulling at VLTI:</b>	<b>on hold (polarisation, CV approval)</b>
<b>COROT:</b>	<b>(F) transition search, large rocky planets:</b>	<b>launch 2006</b>
<b>PRISMA:</b>	<b>(S, D, F) 2 Free Flying techno precursor s/c for 35 M€:</b>	<b>launch 2008</b>
<b>KEPLER:</b>	<b>(NASA) transition search, also for Earths:</b>	<b>launch 2008</b>
<b>HERSCHEL:</b>	<b>(ESA) dust levels in target systems:</b>	<b>launch 2008</b>
<b>GAIA:</b>	<b>(ESA) completion of star catalogue:</b>	<b>launch 2011 / 12</b>
<b>SIM:</b>	<b>(NASA) space interferometer:</b>	<b>launch, 201?</b>
<b>TPF-C:</b>	<b>(NASA) space coronagraph:</b>	<b>on hold</b>
<b>TPF-I:</b>	<b>(NASA) nulling interferometer: Darwin partner/counterpart</b>	<b>on hold</b>

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# Darwin preparatory work

- **Technology Research Program**
  - 5 Meuro in 2006, 4.5 Meuro in 2007
  - Expect similar funding during Cosmic Vision
  - Major goal: 'realistic' nulling breadboard
  - Bring enabling technology in place
  
- **HERSCHEL program**
  - Building on Spitzer legacy
  - Unbiased search for Kuiper belts / exo-zodi in GT program of closest 80 FGK stars.
  - Sensitivity is 1 Kuiper-belt around G2V at 10pc with S/N of 5 in ~30 min.

# COROT:

- Launch in fall of 2006 (Soyuz from Bajkonur) into LEO
- Capacity to detect 2 Earth-radii planets in short period orbits
- Field of view 2 deg<sup>2</sup>; 1 field every 150 days for 2.5 years
- Two tasks: astroseismology and planet finding
- ~50 astroseismology targets and 60000 planet-finding targets



# KEPLER:

- Launch 2009, results in 1-3 years, total mission time 4 years
- Scans 100,000 stars in a 150 sq. deg. field in Cygnus-Lyra
- Expected to find ~50 'real Earths'  
= 1  $R_{\oplus}$  in HZ = 1 AU around G2V)



# Astrometry in space & on ground

## Goals:

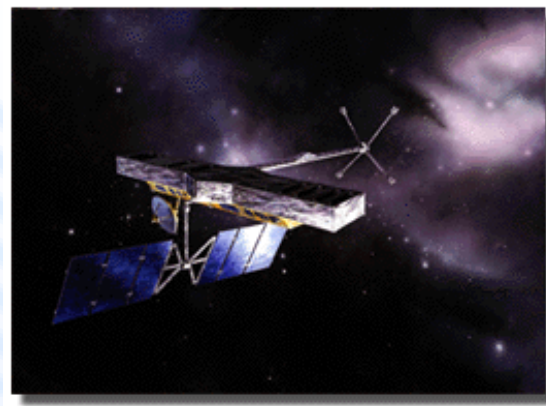
- complete census of giant planets to 200 – 500 pc
- orbital characteristics of several thousand systems

Present: Using FGS on HST (GL876b mass determination)

Soon: VLT-PRIMA

2011/12: GAIA

201?: SIM-PlanetQuest



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# DARWIN-Lite or Super-PEGASE

- First mission in CV1525 is too small an envelope for DARWIN proper (300 Meuro)
  - CV plan allows for several (sub-)missions or experiments to answer a Theme
- Investigate Darwin-lite:
- Cheaper mission with less encompassing science case
  - Smaller DARWIN or larger PEGASE: "Where shall the twain meet?"
- Outcome of initial study by Univ. of Liège:
- **Science case not compelling**
  - **Cost & complexity NOT significantly reduced!**
  - **TE-SAT recommends NO Darwin-lite!**



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# Where do we go from here?

## 0. Get DARWIN into Cosmic Vision program

- Role of Europlanet = ? But must be significant

## 1. Continue development of enabling technology

## 2. Broaden interest base: Again a role for Europlanet

- Solar system community
- Biology
- Paleontology
- Geophysics

## 3. Work on preparation

- Study of target systems: Kuiper-Edgeworth belts, exo-zodiacal dust, inclination, etc...
- Models of planetary evolution