

What Are the Outstanding Issues with Coronal Jets?

Alphonse C. Sterling

Main Additional Contributions:

Ronald L. Moore

Navdeep Panesar

David Falconer

Mitzi Adams

Supported by NASA's HGI program, NASA NPP program, and the MSFC/Hinode project.

Invited Talk

4. Eruptions in the solar atmosphere

What are the outstanding issues with coronal jets?

Alphonse C. Sterling¹

¹ *NASA/MSFC*

Solar coronal jets have been observed in X-rays since the early 1990s. Since then, high-cadence, high-resolution observations of them in the EUV with SDO/AIA, and similar advances in magnetic field information with SDO/HMI, resulted in a revolution in thinking about the mechanisms leading to and driving the jets. It now appears that at least many jets result when a small-scale filament (minifilament) erupts, and the field of that erupting minifilament undergoes magnetic reconnection with pre-existing surrounding field. Moreover, a primary - if not exclusive - mechanism for building the minifilaments and triggering them to erupt is cancelation of magnetic flux in the photosphere near the location from where the minifilament/fluxrope erupts. This presentation will discuss outstanding questions regarding coronal jets, such as the need to verify the above scenario with more data; confirming whether the same mechanism(s) drive jets in all solar regions, including active regions, quiet Sun, and coronal holes; and determining whether there is a threshold condition (or set of conditions) necessary for driven reconnection to result in explosive jets.

What Are the Outstanding Issues with Coronal Jets?

Alphonse C. Sterling

Main Additional Contributions:

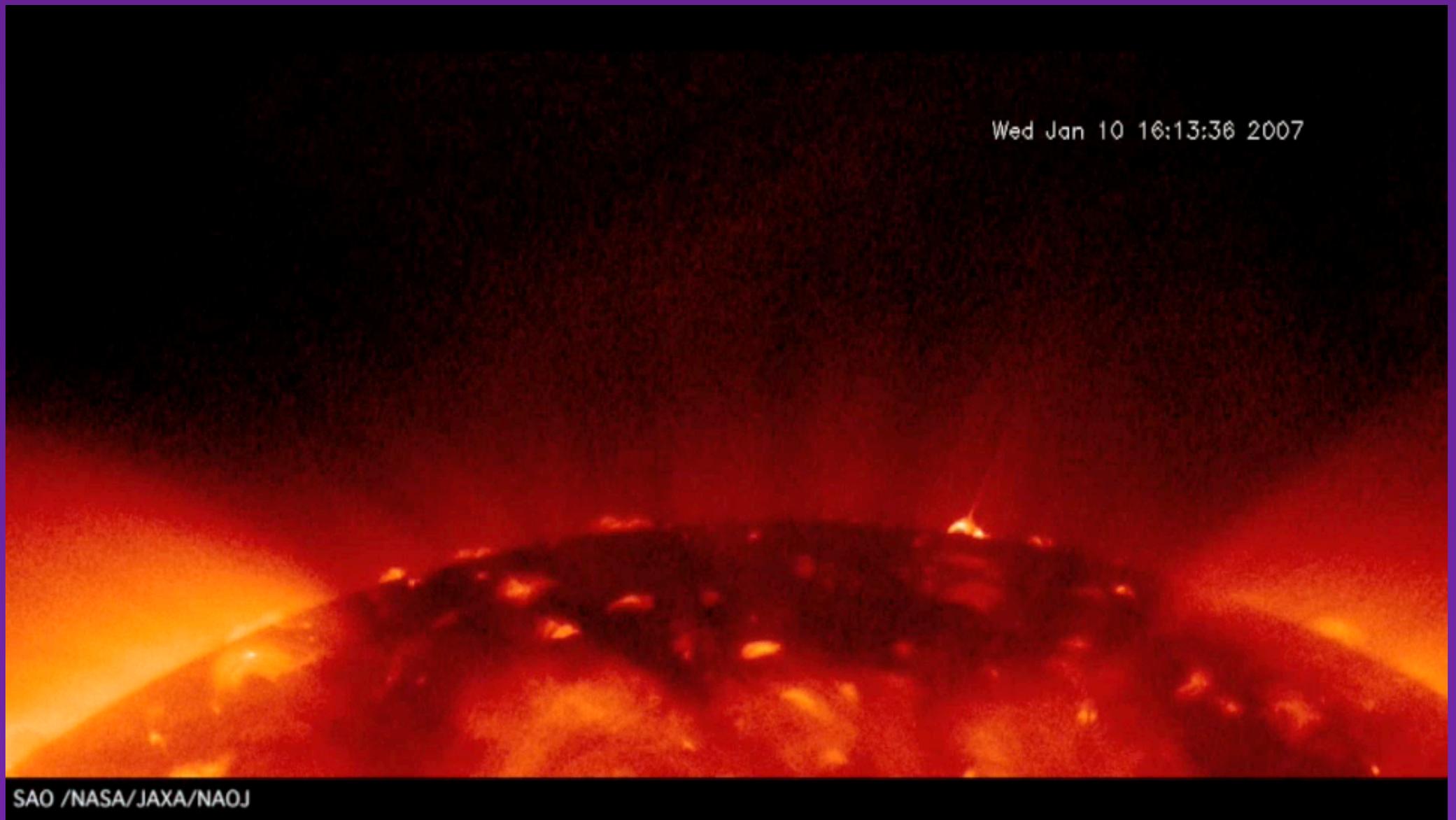
Ronald L. Moore

Navdeep Panesar

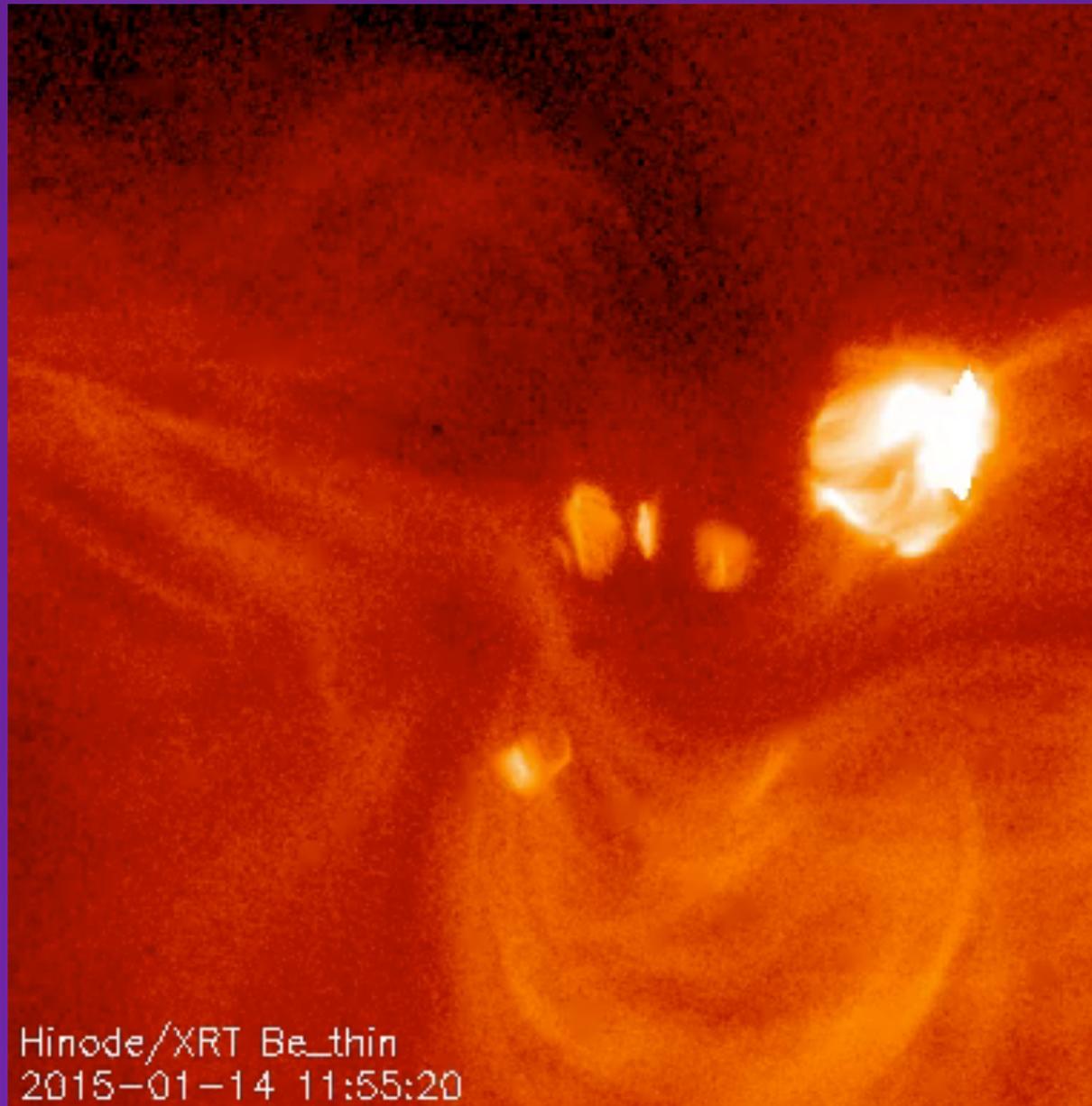
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Cirtain et al. (2007)



Sterling et al. (2017)

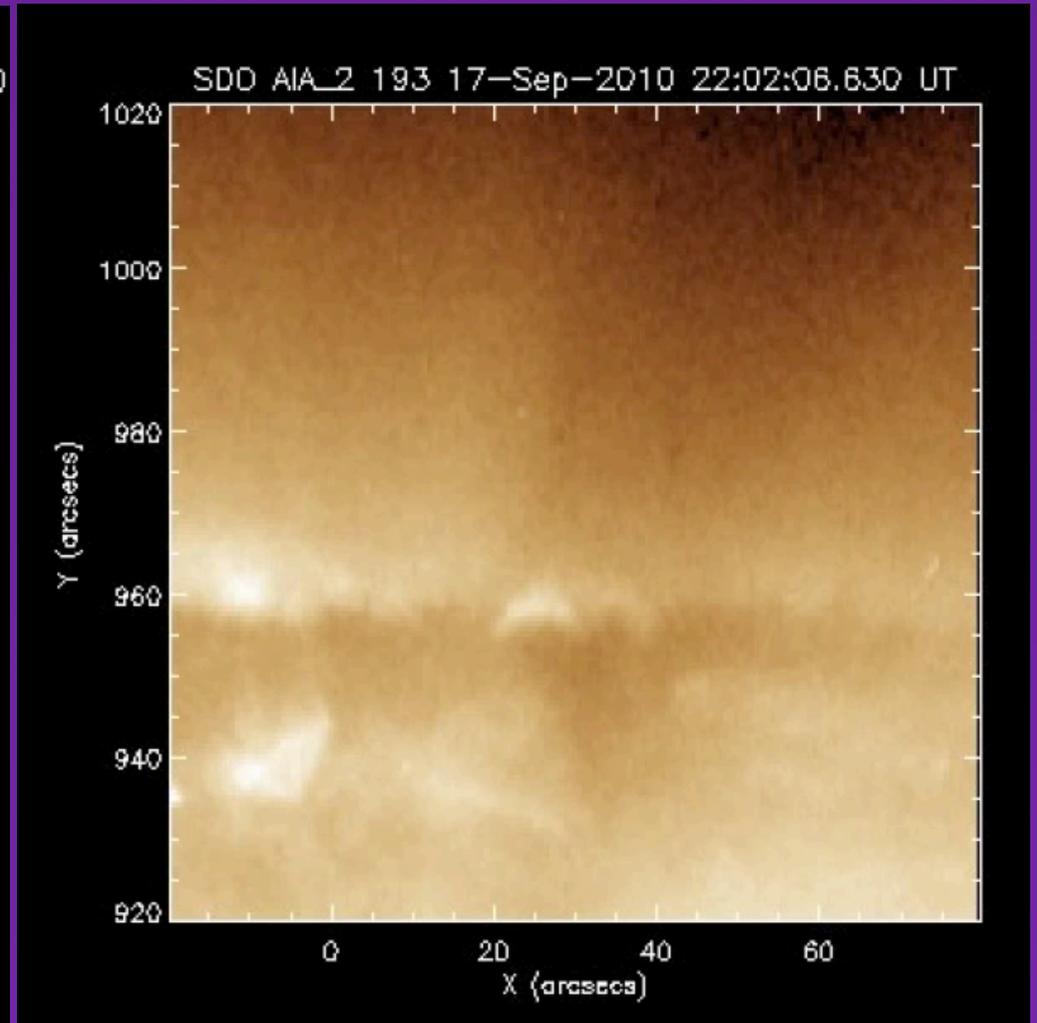
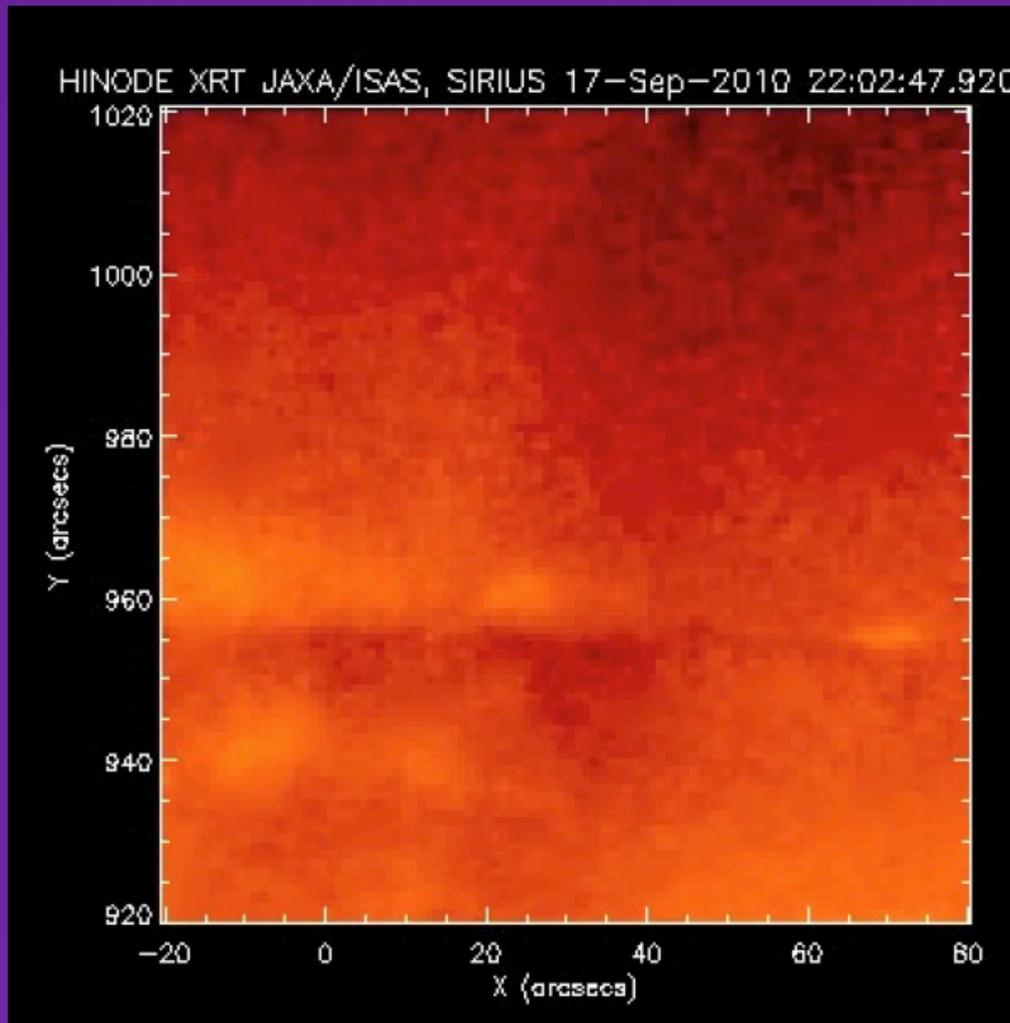
Introduction: Solar X-Ray Jets

- ◆ Observed since the Yohkoh days (Shibata et al. 1992; also Shimojo et al. 1996, etc. Reviewed by Raouafi et al. 2016.)
- ◆ Yohkoh (SXT) saw them mainly in active regions.
- ◆ Hinode/XRT found them to be plentiful in polar coronal holes (Cirtain et al. 2007; also Savcheva et al. 2007, etc.)
- ◆ Stereo EUVI+coronagraph (Nisticò et al. 2009, 2015).
- ◆ In polar coronal holes: size $\sim 50,000$ km x 8000 km; rate ~ 60 /day (Savcheva et al. 2007).
- ◆ Often have a “hot loop” at the jet’s base.
- ◆ Previously often-discussed mechanism is based on emerging flux (“emerging-flux model”). (Shibata et al. 1992; Yokohama & Shibata 1995, 1996; see also Moore et al. 2010.)
- ◆ Many of the above ideas deduced from SXR, and pre-SDO AIA observations.

Coronal Hole Jets: "Minifilament eruptions"

XRT

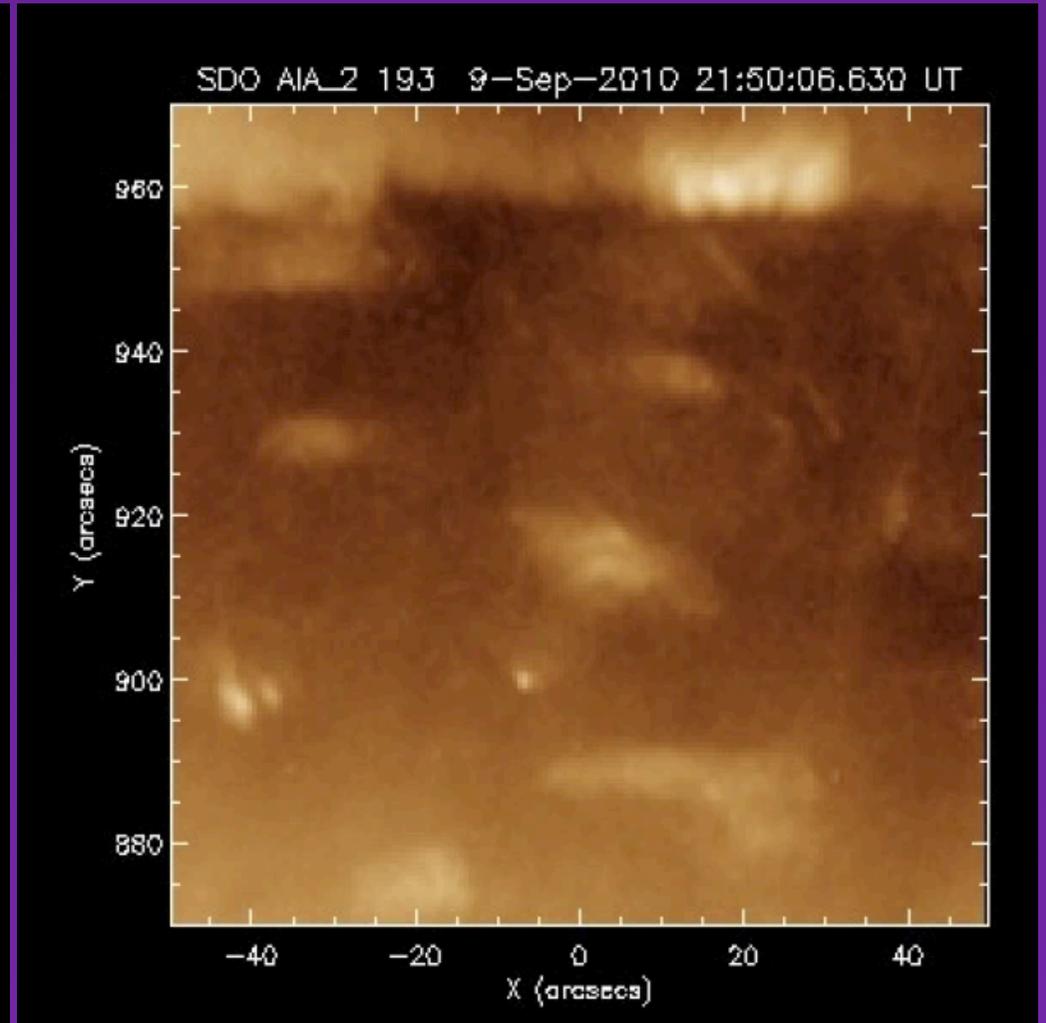
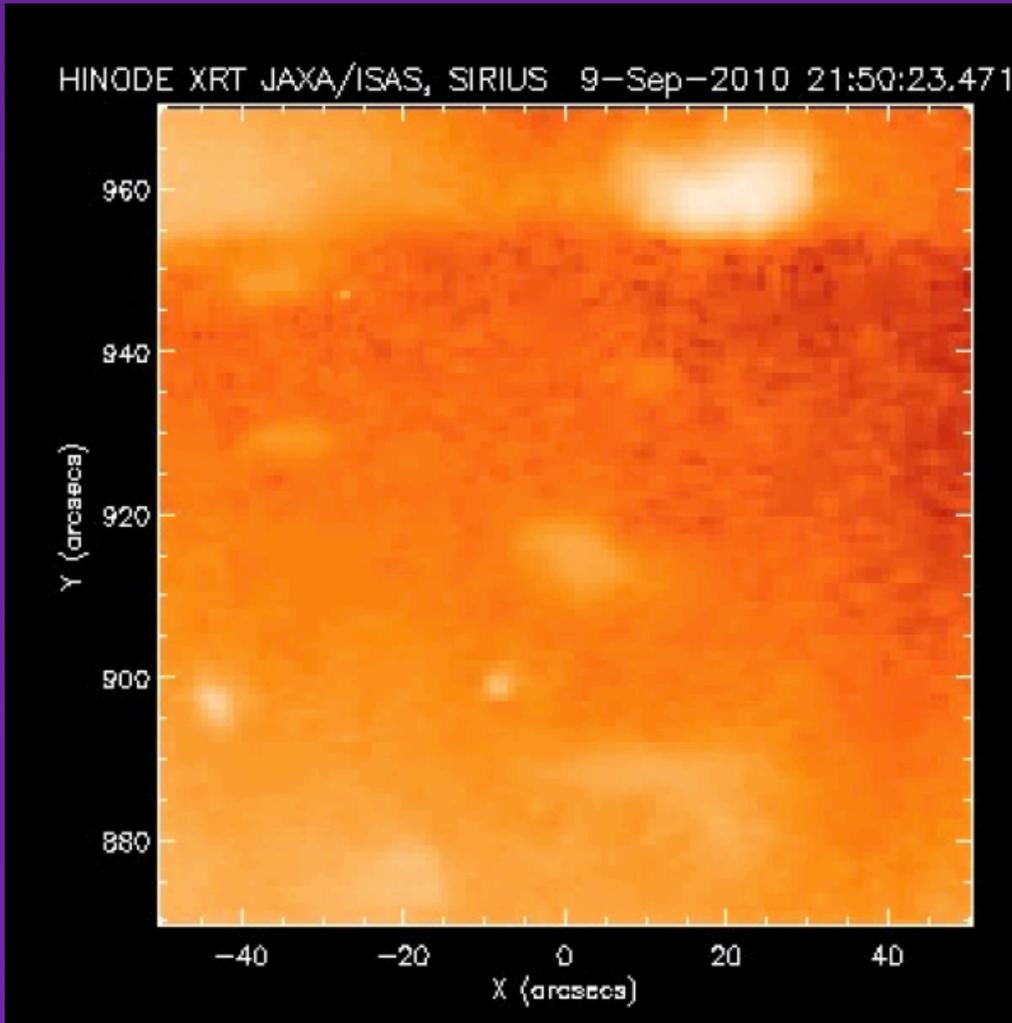
AIA 193



Sterling et al. (Nature, 2015): 20 Polar CH jets.

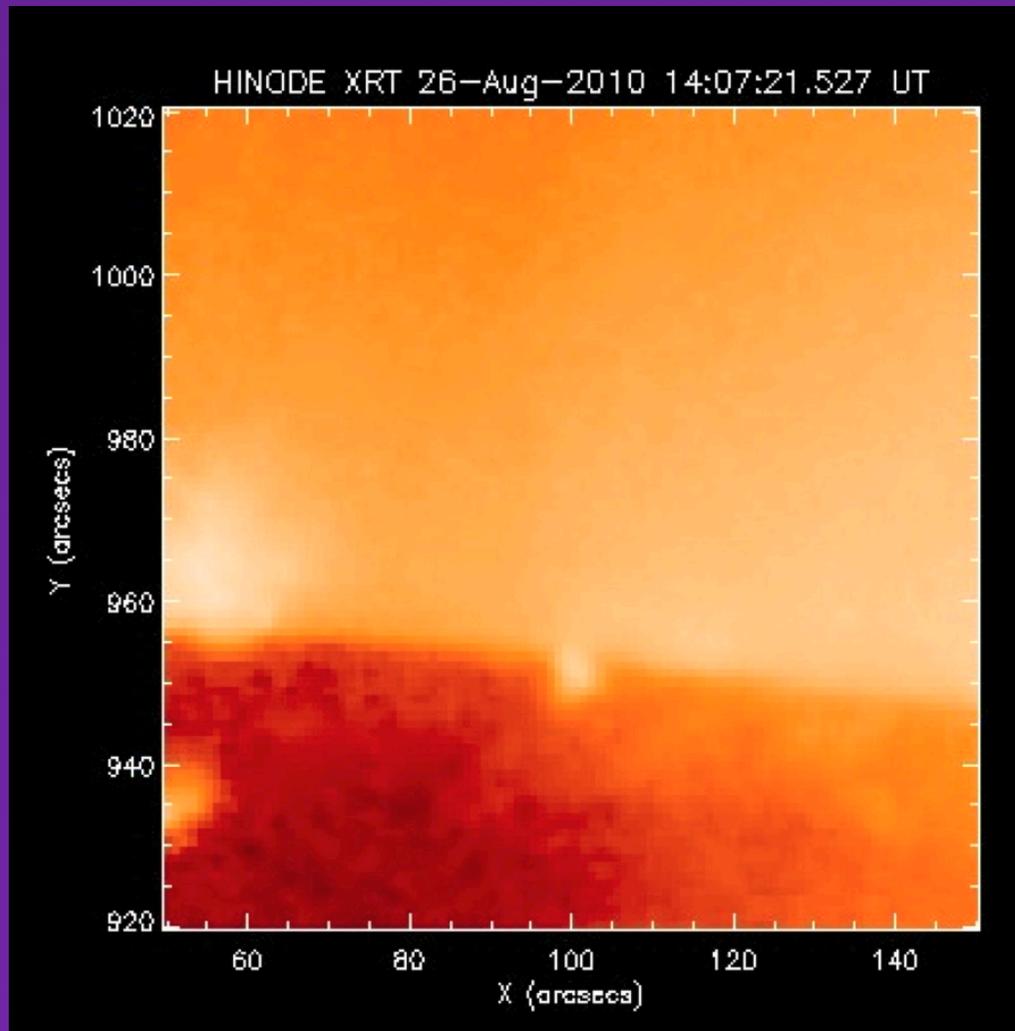
XRT

AIA 193

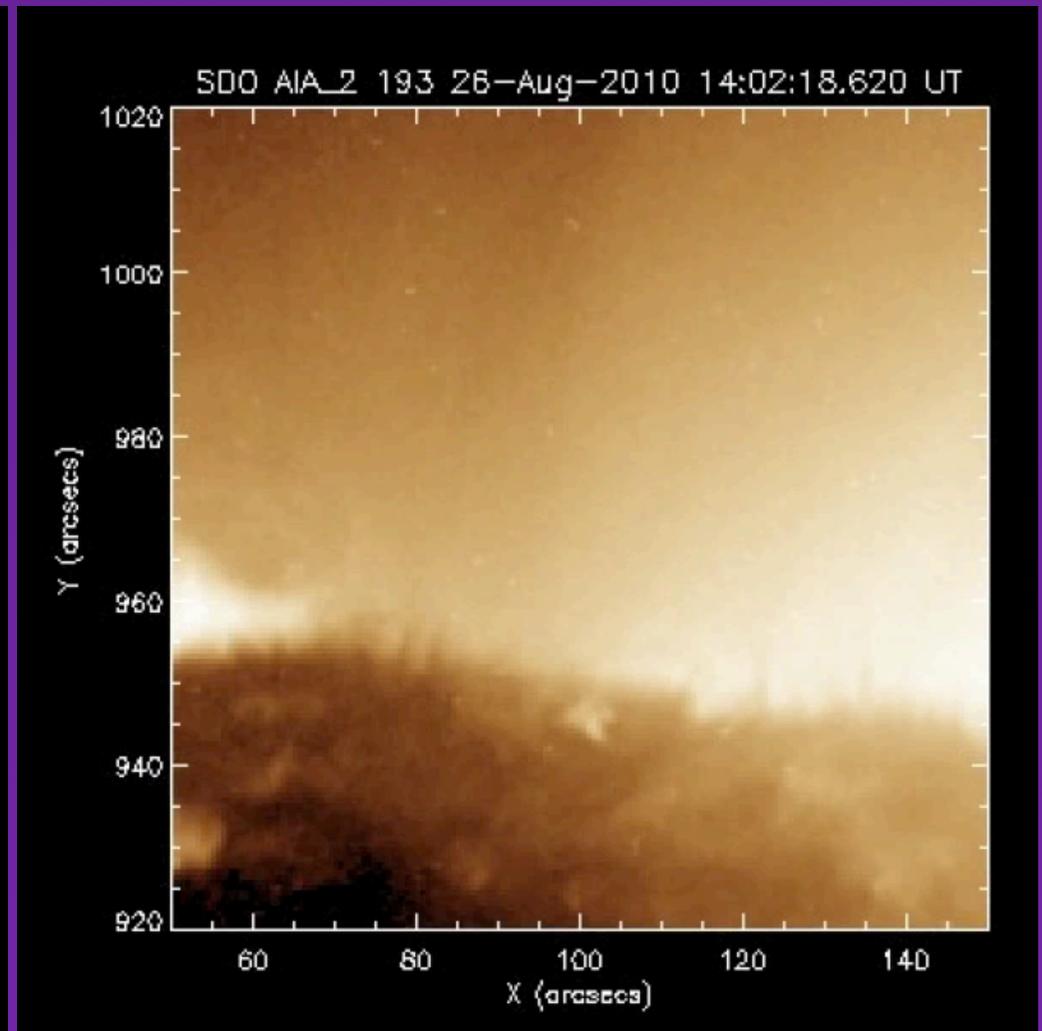


Event 12

XRT

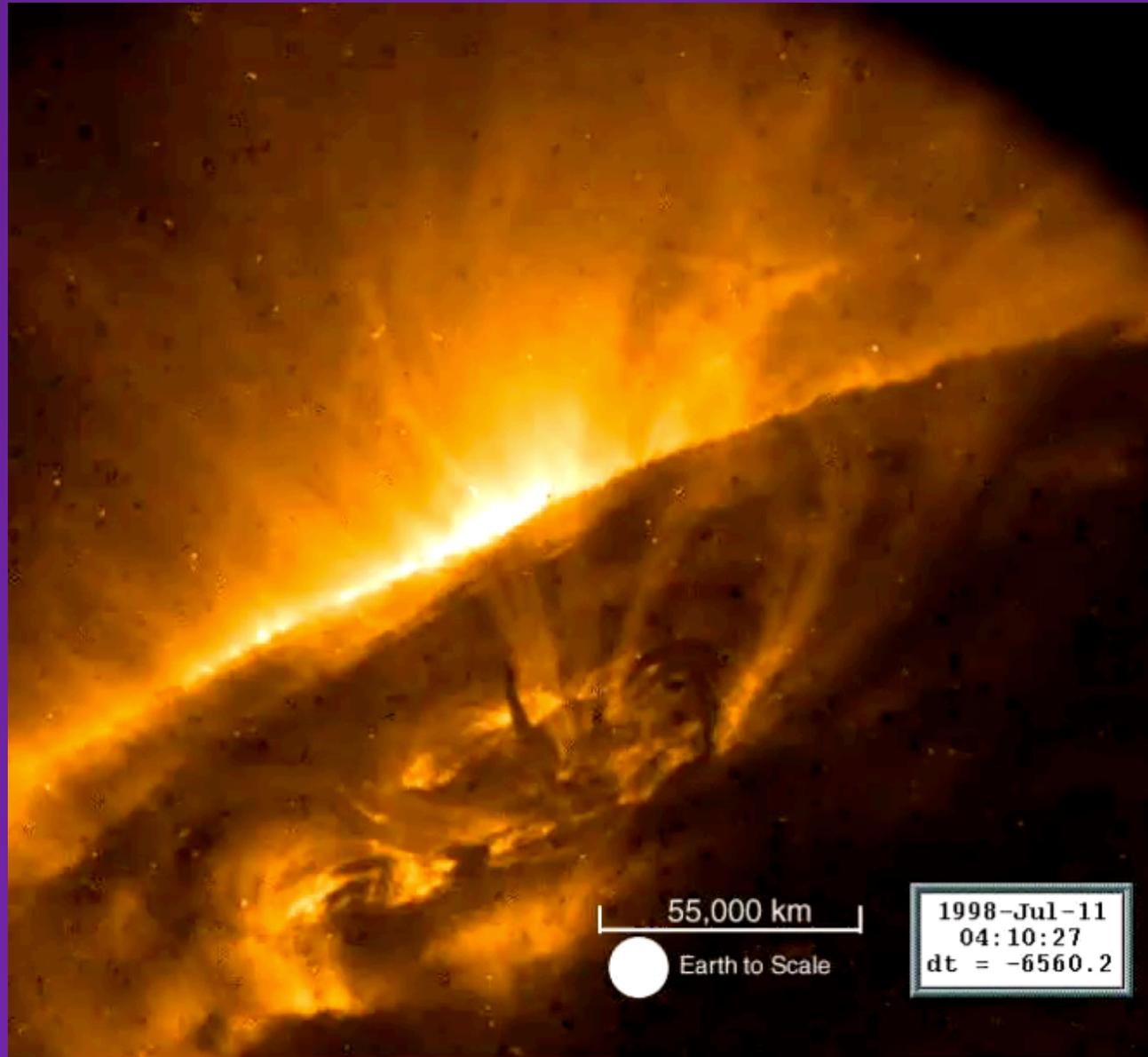


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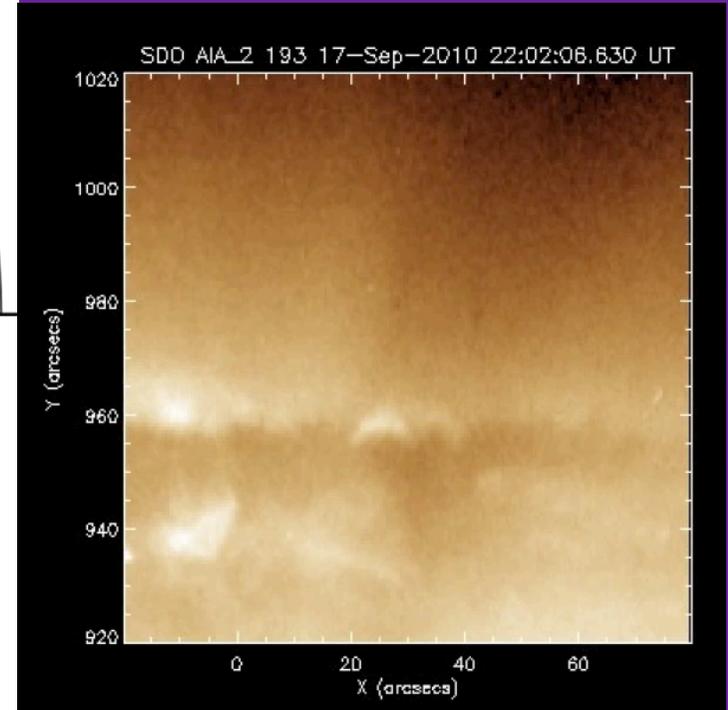
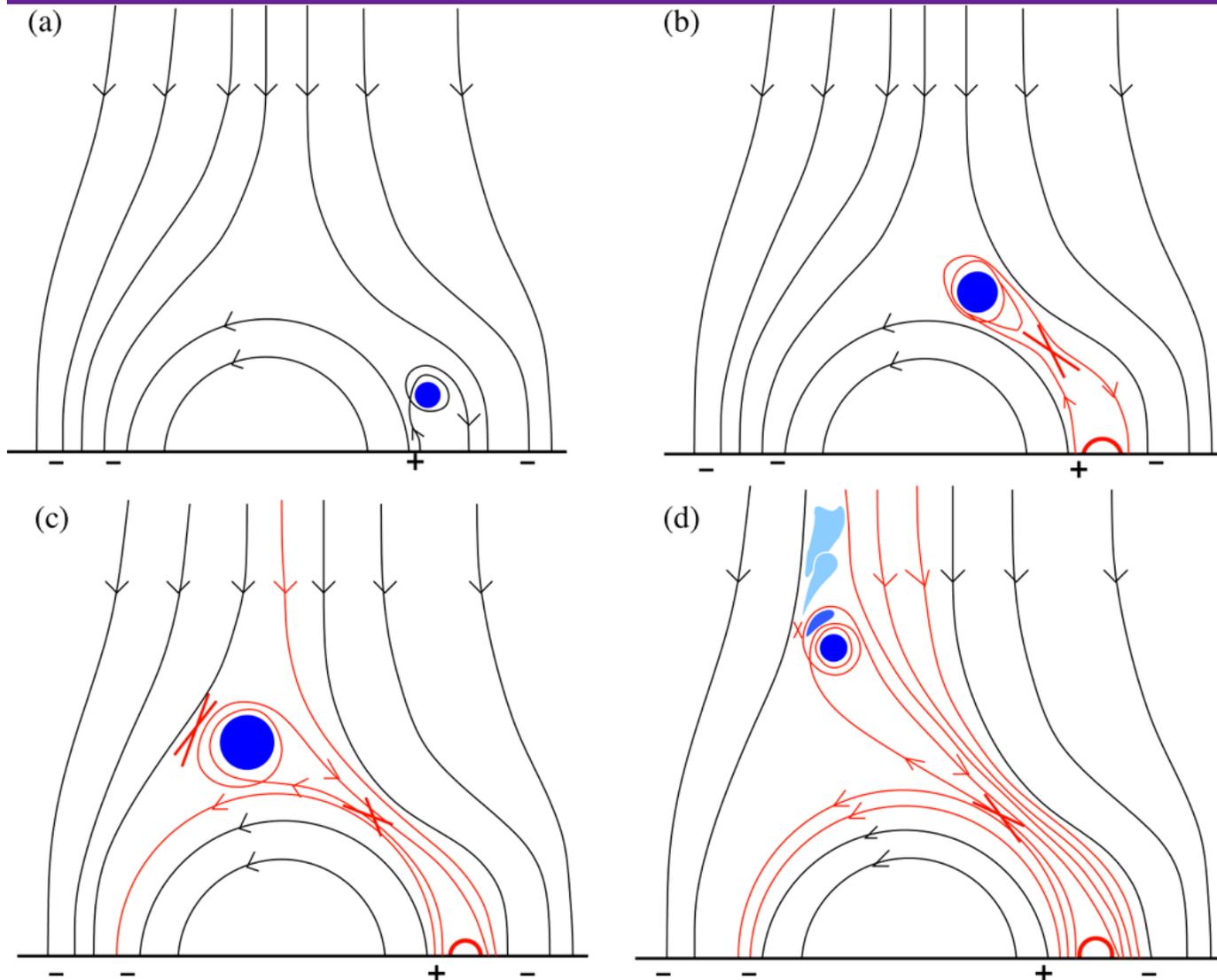


Event 3

“Normal” Filament Eruption (TRACE)



Minifilament-Eruption Model for (X-Ray) Jets



Sterling et al. (2015, 2016, 2017)

Quiet Sun jets work the same way (Panesar et al. 2016b)

Recently modeled by Wyper et al. 2017, 2018)

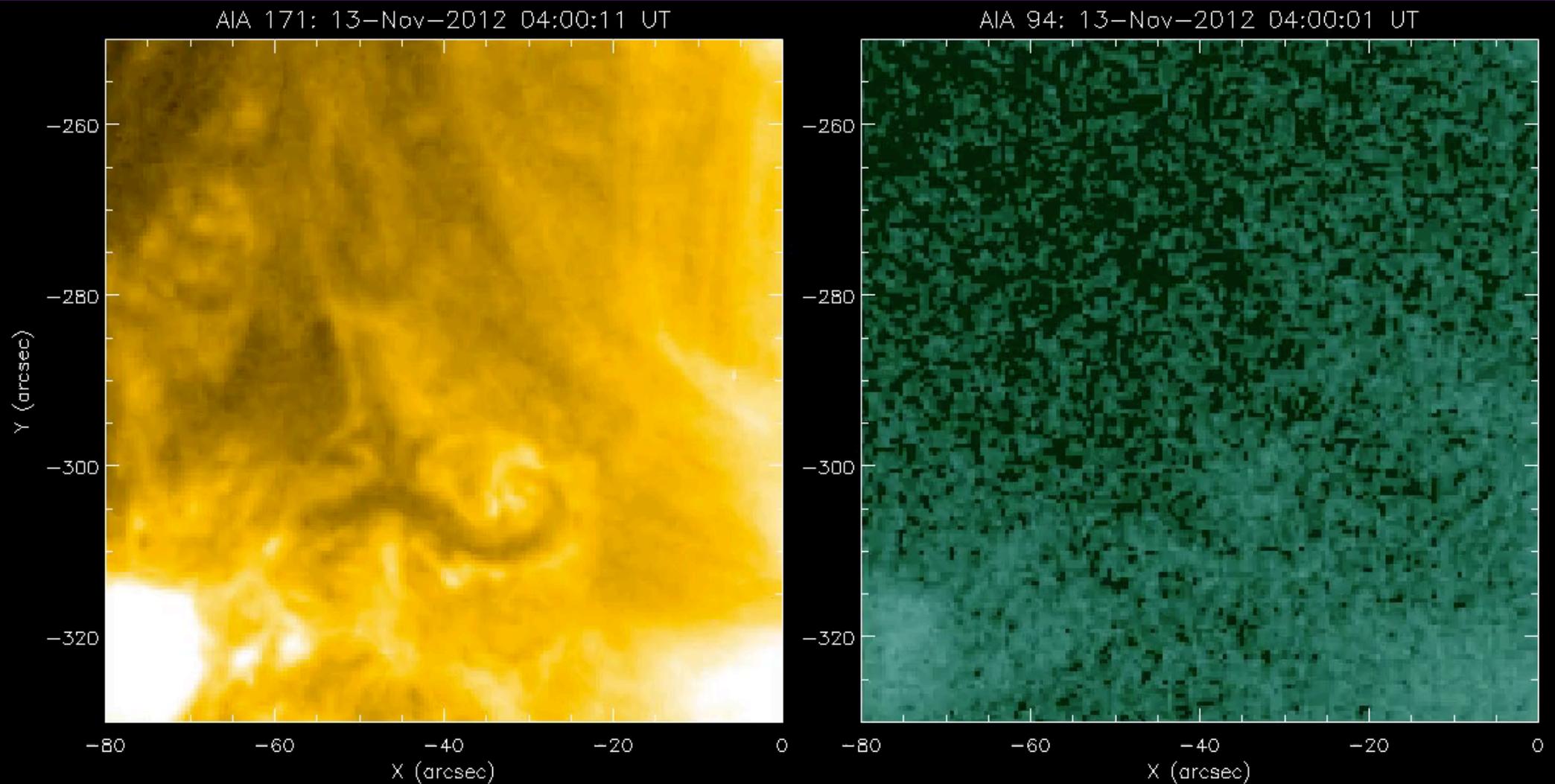
What Causes Miniature-Filament Eruptions?

- ◆ Did not look on-disk in this study, due to polar view. But....
- ◆ Adams et al. (2014) found no emerging flux in the jet region. Filament erupted from location where flux canceled. (Also, Hong et al. 2014.)
- ◆ Several other found cancelation leading to jets (e.g., Hong et al. 2011; Huang et al. 2012; Young & Muglach 2014a,b).
- ◆ Some others found jets from location of emerging flux+flux cancelation (e.g., Liu et al. 2011; Shen et al. 2012, 2017; Hong et al. 2012; Li et al. 2015).

Quiet Sun Jets — Similar to PCH jets

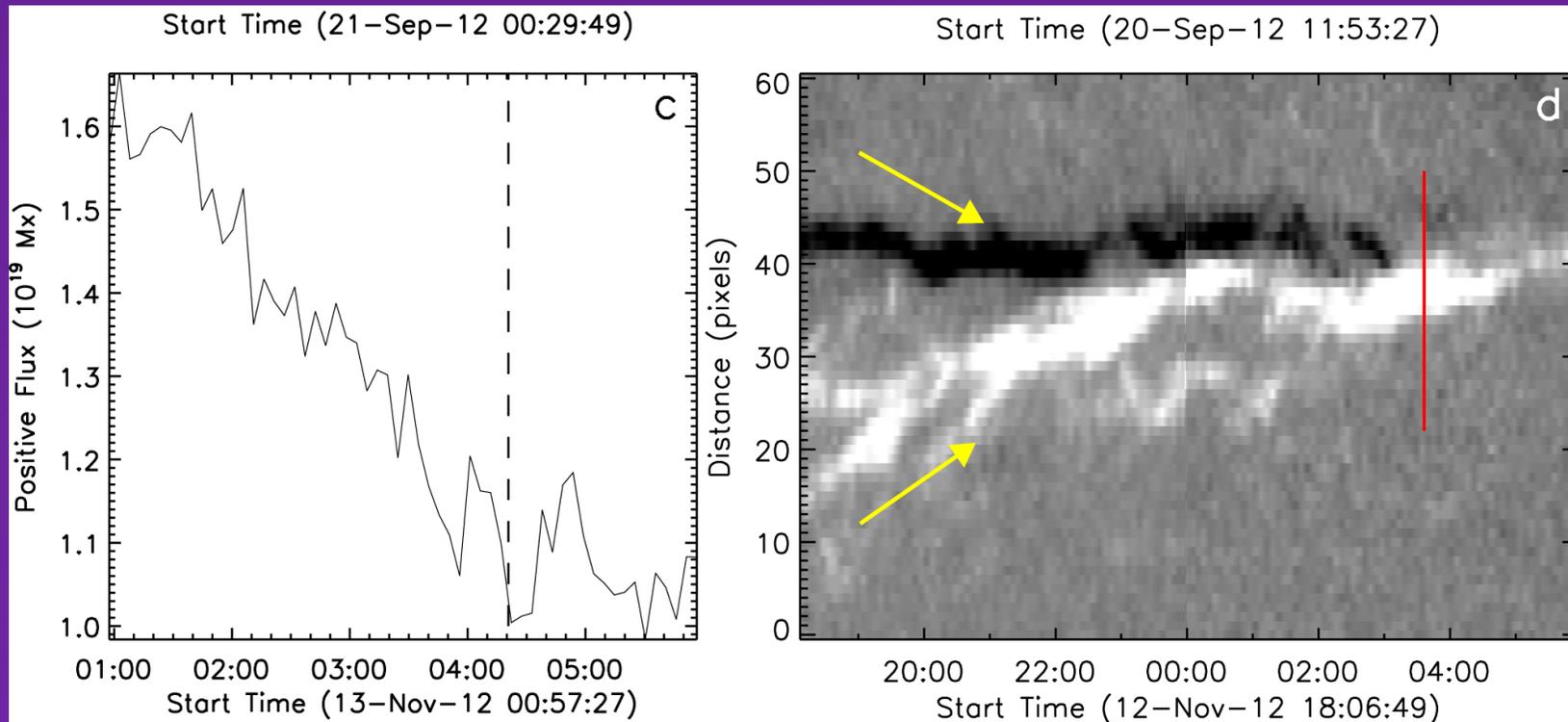
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(Panesar et al. 2016b)

Same for QS jets: Occur at cancelation sites.



(Ave. Cancellation rate: $\sim 10^{18}$ Mx/hr.)

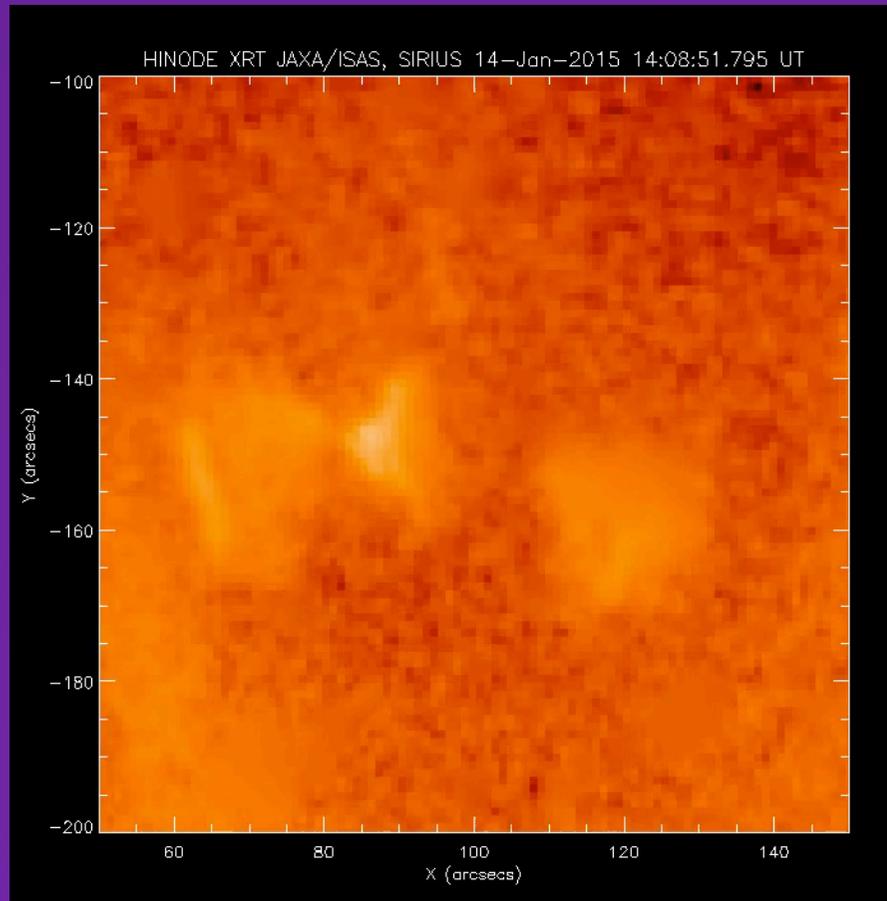
Panesar, Sterling, & Moore (2016b) — 10 jets.

Active Region Coronal Jets

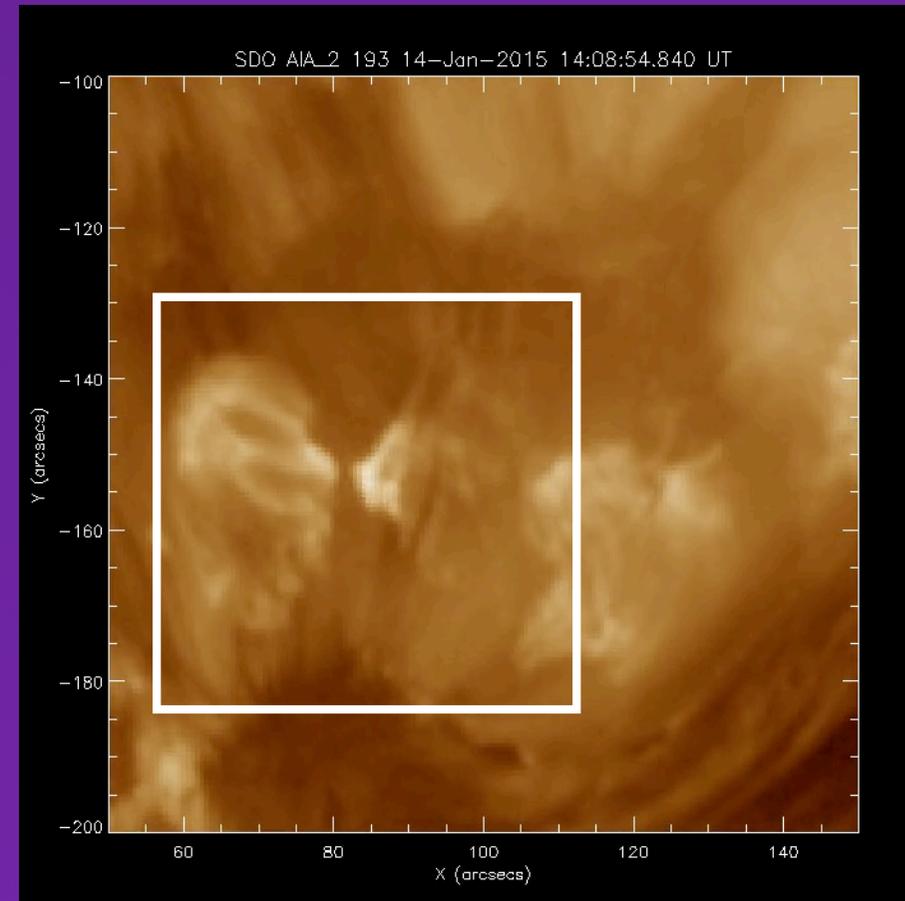
- Yohkoh studies (Shibata et al., Shimojo et al., many others).
- Raouafi et al. (2016).
- Mulay et al. (2017a, b) - AR-jet temps/emissions.
- Hong et al. (2017) — Minifil. eruption —> AR jet & Type III burst (also Shen et al. 2017, Moroccan et al. 2017).
- Panesar et al. (2016a); Sterling et al. (2016, 2017).

An Example: AR Jets

- 14 Jan 2015 (NOAA AR 12259), AIA, HMI, Hinode, IRIS.
- Sterling et al. (2017)



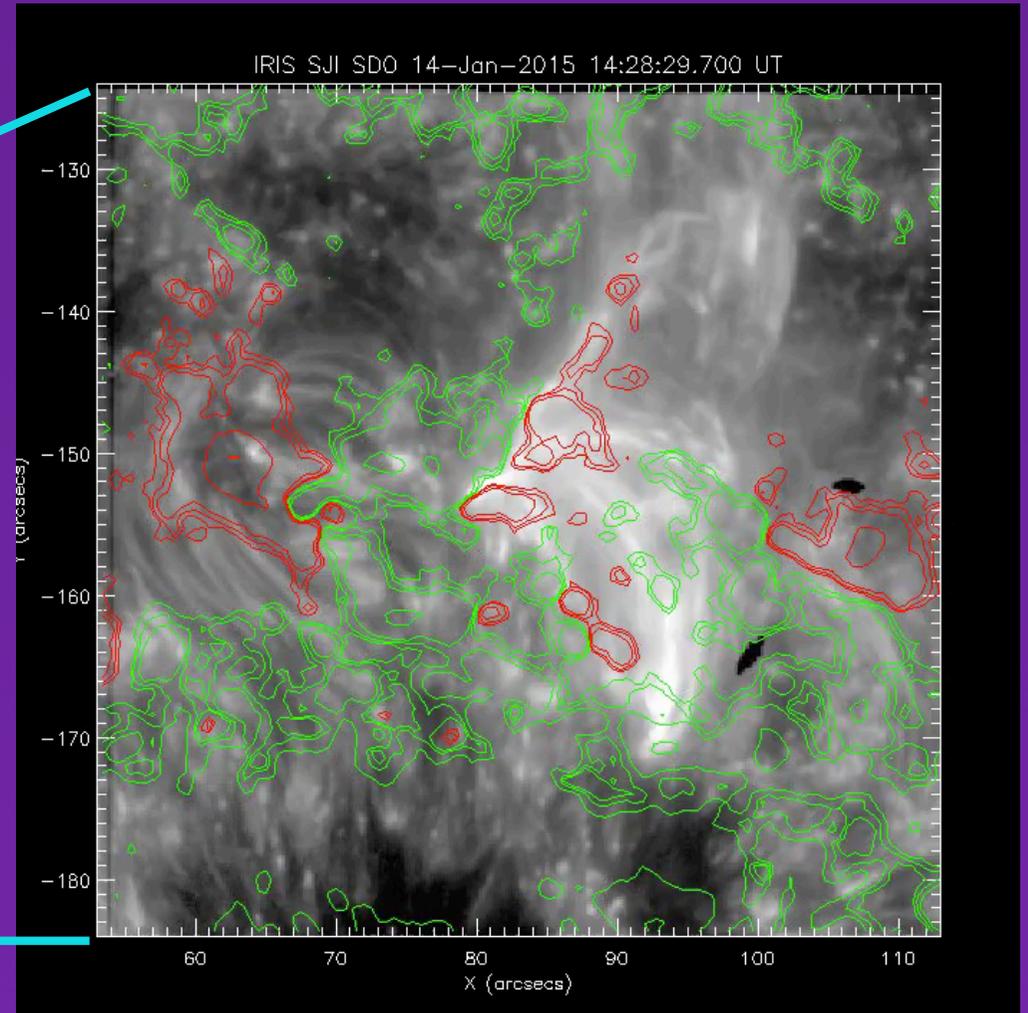
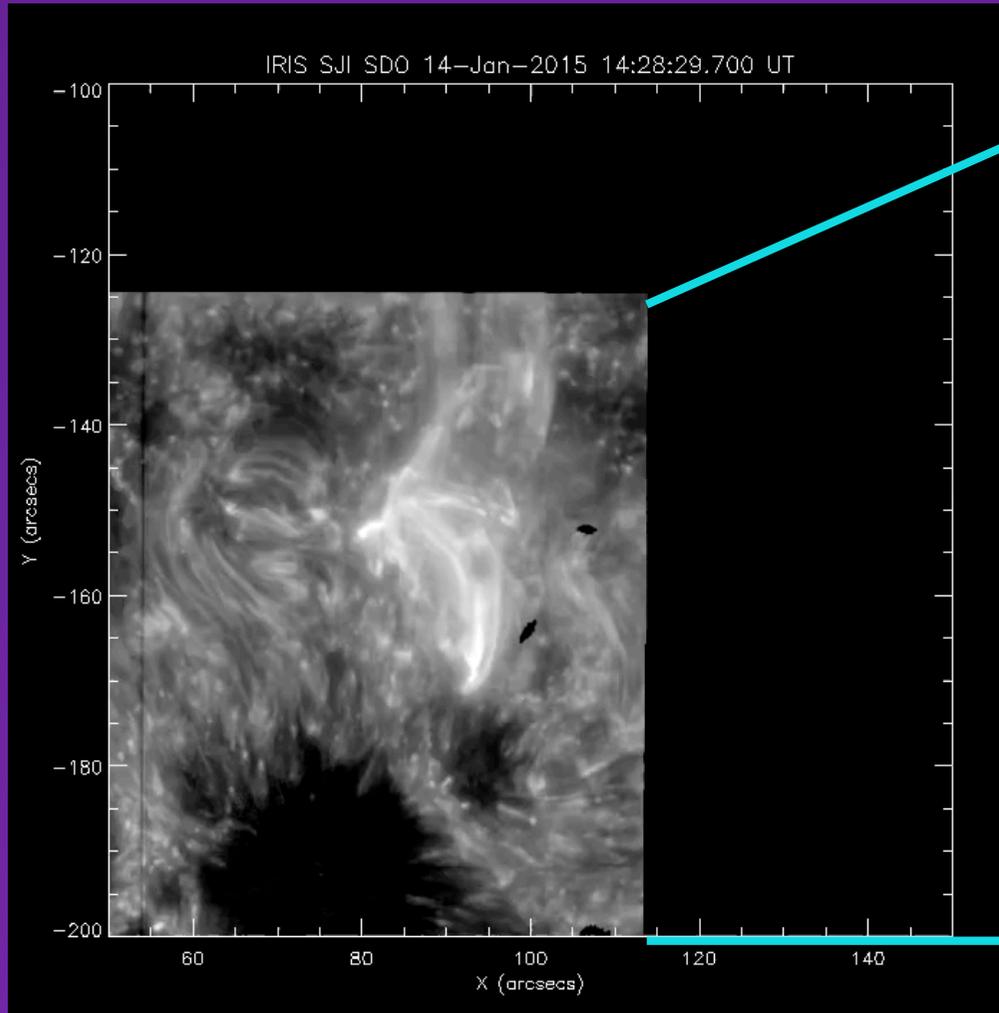
Hinode/XRT



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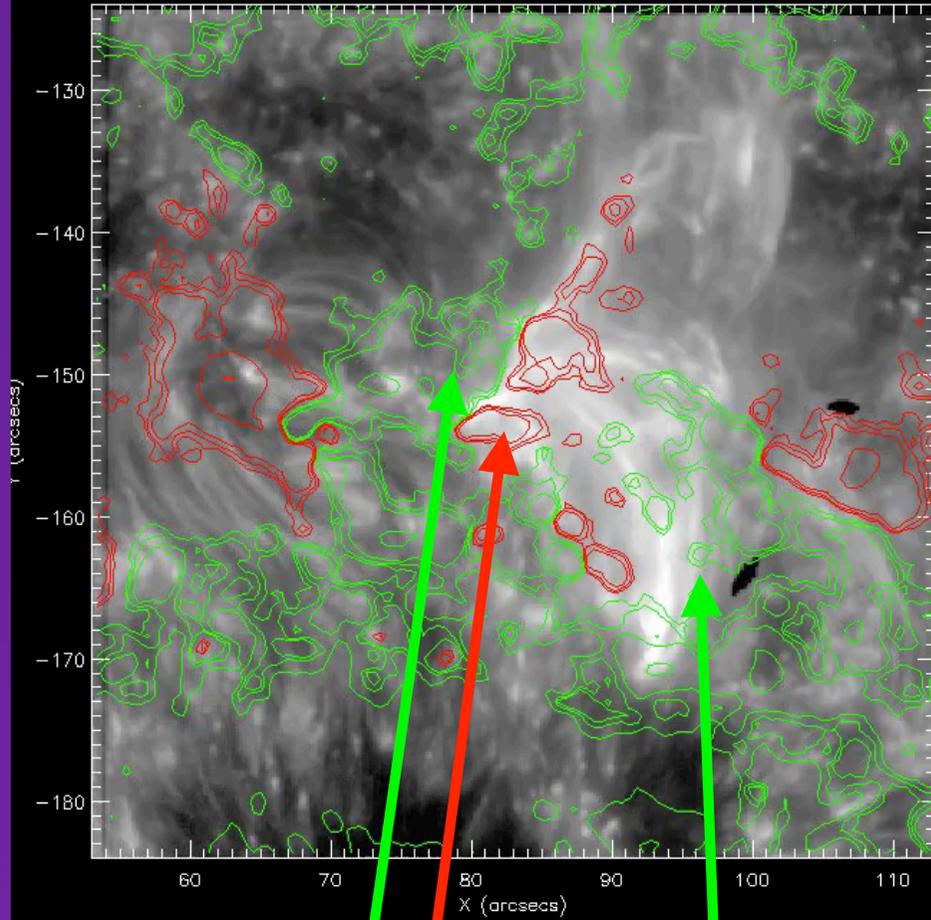
Minifilament hard to see (absent?). Work the same way??

Coronal Jets in Active Regions

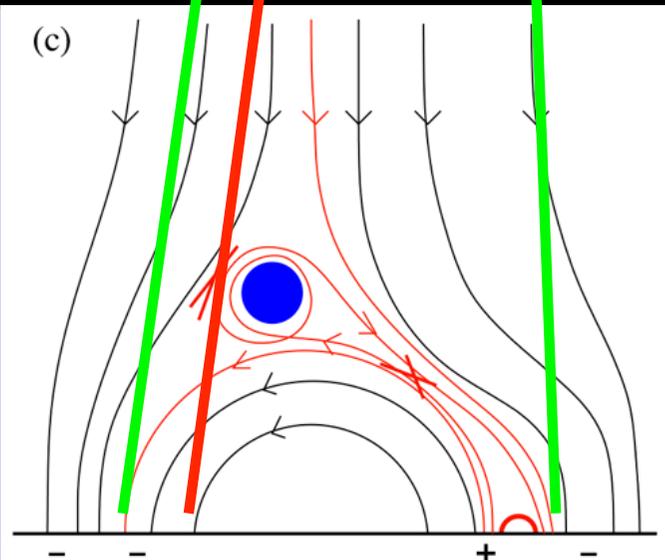
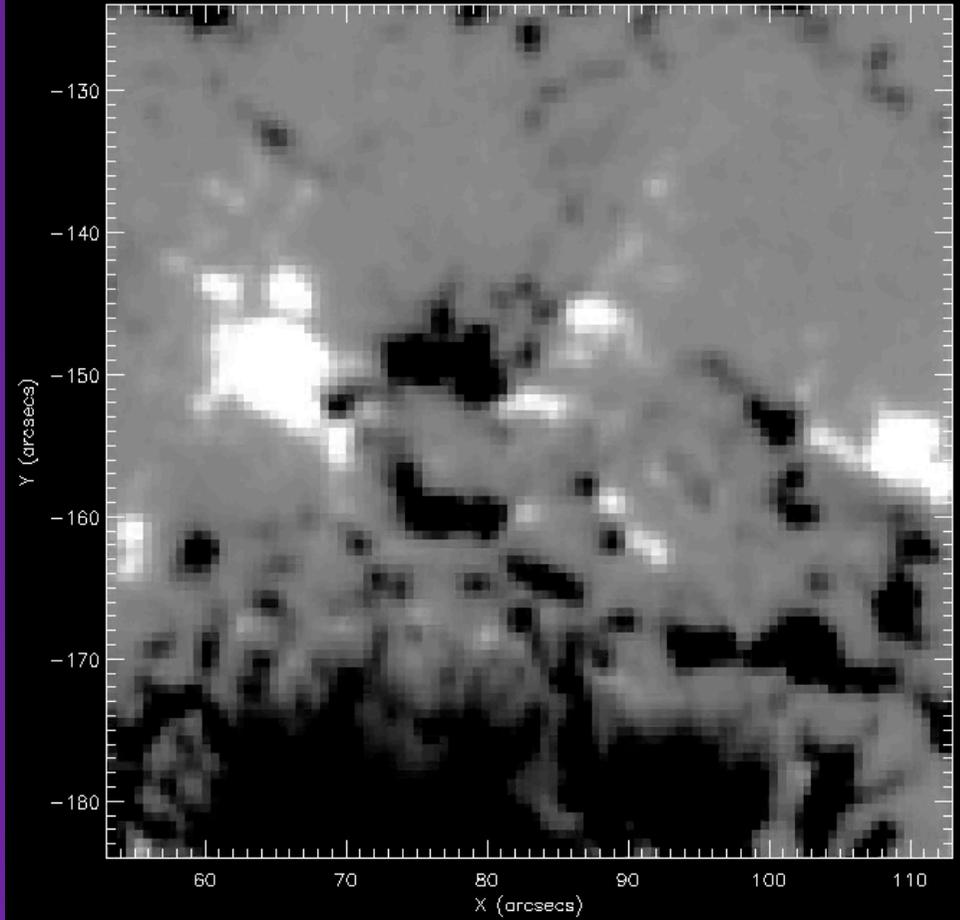


Sterling et al. (2017)

IRIS SJI SDO 14-Jan-2015 14:28:29.700 UT

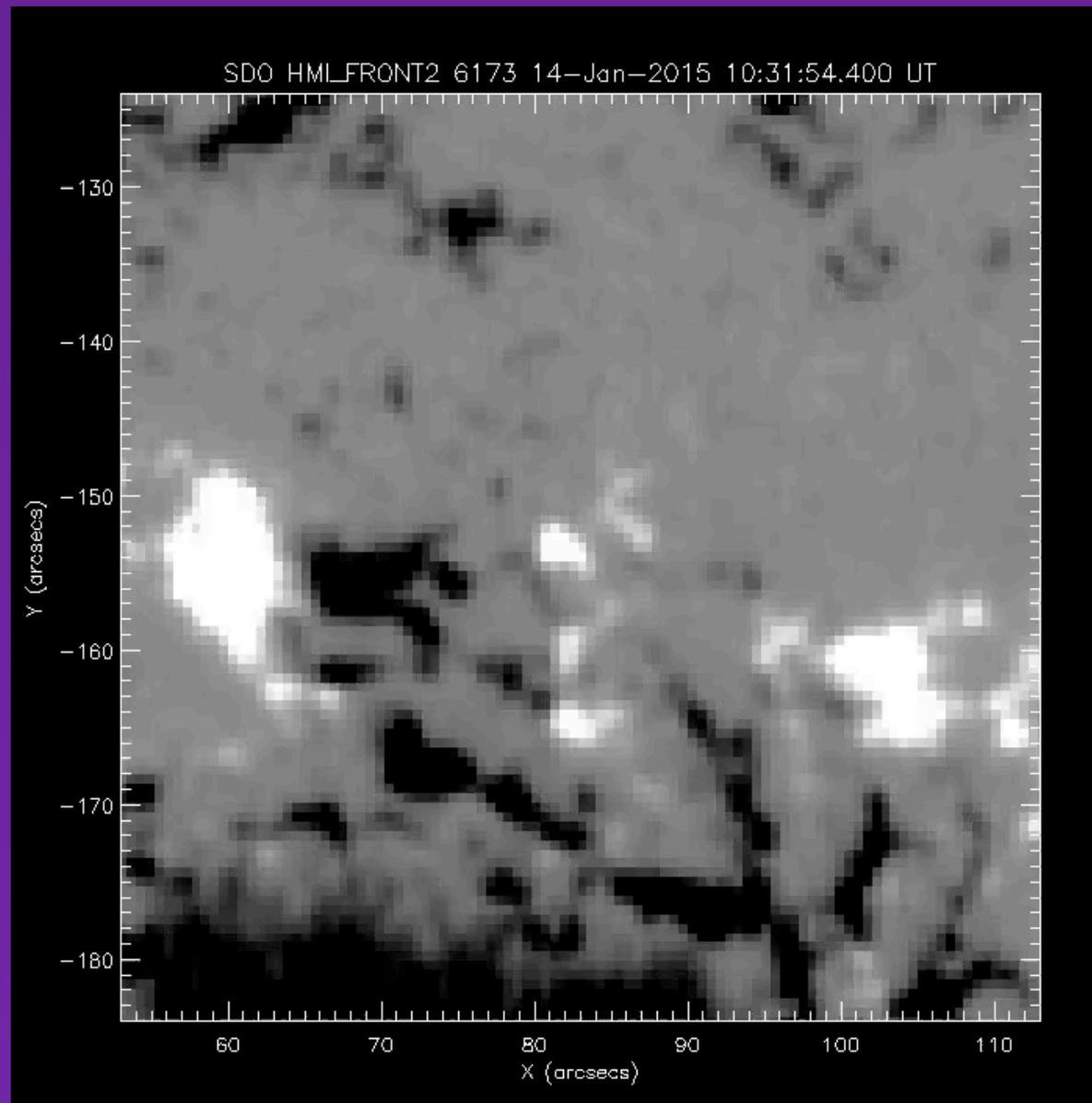


SDO HMLFRONT2 6173 14-Jan-2015 14:29:39.300 UT



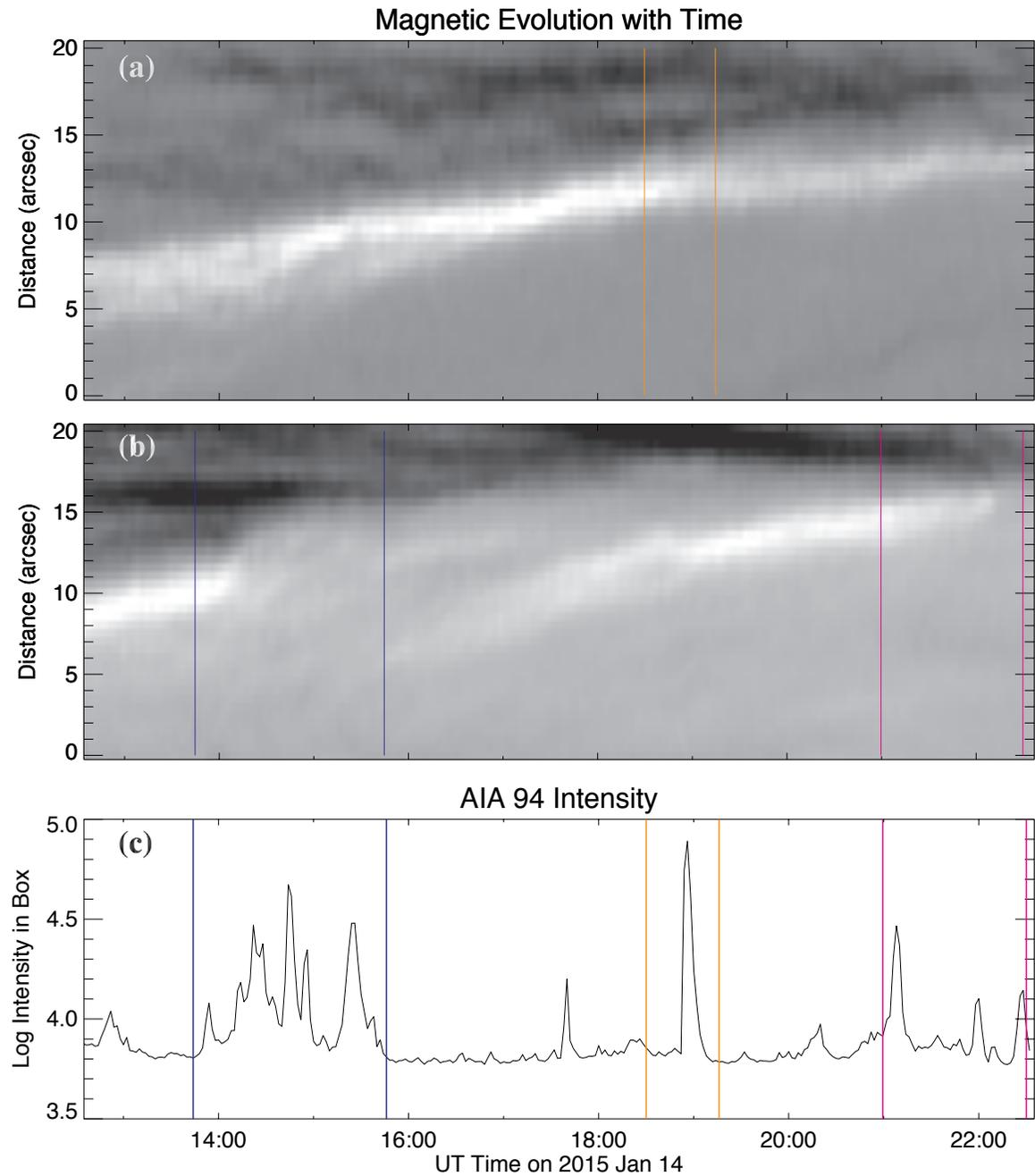
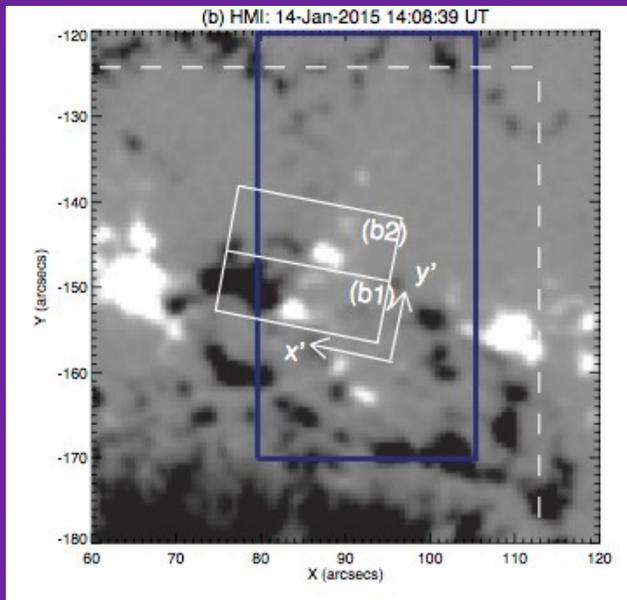
Sterling et al. (2017)

HMI of jetting region

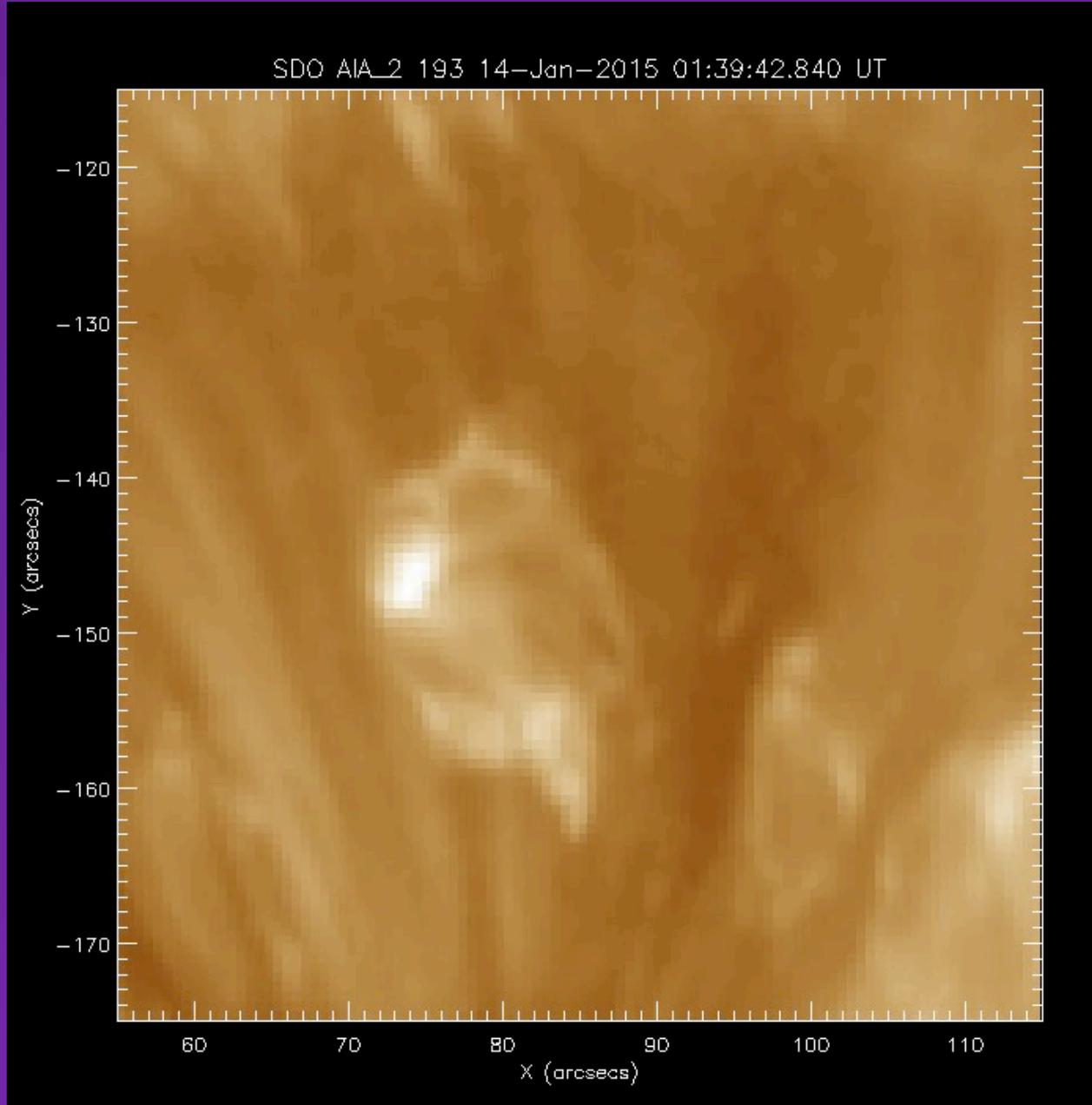


Jets occur at *flux cancellation* locations!

AR jets (Sterling et al. 2017)



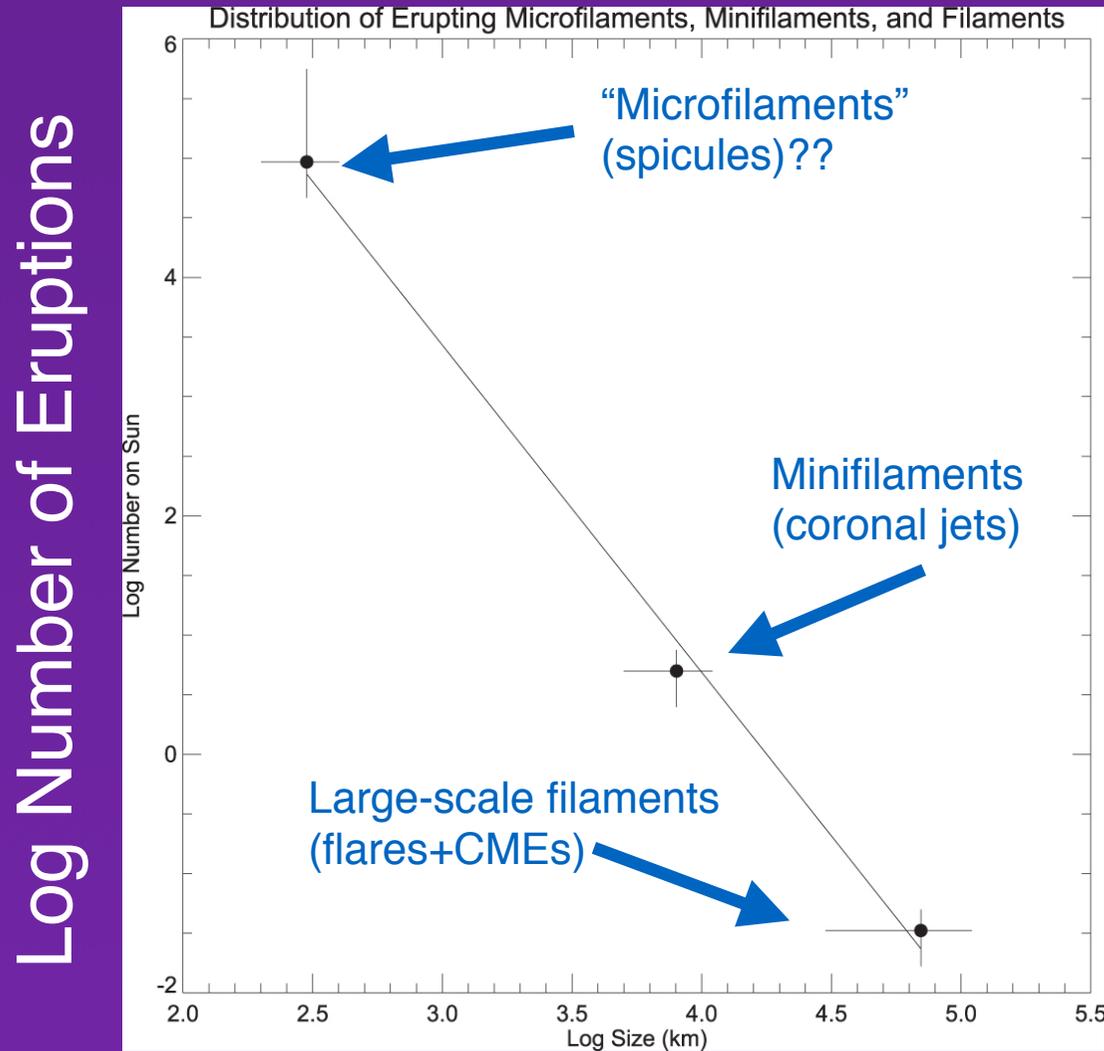
Minifilament “strand” visible from neighboring region, slightly different time.



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Do Jets Exist on Smaller Size Scales?

Filament-Like Feature Eruptions on Smaller Scales??



Log “Filament” Size

Sterling & Moore (2016)

“Jetlets” in plumes (Raouafi & Stenborg 2014)?
 (“Jetlets” in more general network?? Panesar talk).

Some Outstanding Questions

- What causes jets? Strong evidence that it is flux cancelation in quiet Sun and CHs (20+ events; Panesar talk). Still must study more! (Shear only?? Kumar et al. 2018.)
- AR jets: Minifilaments sometimes less obvious (absent?). Also, “brightest” bright points sometimes in unexpected locations. (Result of complex field, multiple eruptions? Sterling et al. 2016, 2017.) Frequently see cancelation+emergence.
- How do jets scale to smaller structures (contribute to coronal heating? (Moore et al. 2015)).
- If most jets result from flux cancelation, what about larger eruptions? (Field complexities might disguise the key processes; Sterling et al. 2018.)
- Role of twist in powering jets, and “narrow CMEs.”

Summary

- Approaching a good understand of jets, especially in QS and CHs: At least many jets are miniature filament eruptions triggered by flux cancelation.
- AR jets are similar, but they can be more complicated. (Due to complex dynamic field?)
- More observations needed: Factors besides flux cancelation? Can field complexities explain uncertain aspects of AR jets?
- How does jet physics scale to different sizes? (Large eruptions? ``Jetlets''? Spicules??)
- Needed: Jet simulations based on minifilament eruptions (e.g. Wyper et al), and *flux cancelation!*

Image:
Alphonse Sterling
21 August 2017,
Lewisville, Idaho