New science windows with ground-based observations: the Daniel K Inouye Solar Telescope

Gianna Cauzzi^{1,2} for

the DKIST team



INAF – Osservatorio Astrofisico di Arcetri, Italy
 National Solar Observatory, USA

IRIS-9, Göttingen, 25-29 June 2018

Invited Talk

6. Science together with future facilities

New science windows with future ground-based observations - the Daniel K. Inouye Solar Telescope

<u>G. Cauzzi^{1,2}</u>

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 ² INAF-Osservatorio Astrofisico di Arcetri, Firenze (Italy)

The NSO-led, 4-m Daniel K. Inouye Solar Telescope (DKIST) is nearing completion on Maui, Hawaii, with first light foreseen for early 2020. The DKIST will herald a new era of ground-based solar observations, with its polarimetric capabilities, unprecedented spatial resolution, and novel diagnostics in the near IR. In particular, DKIST is expected to revolutionize coronal science, providing high-sensitivity, high-resolution measurements of the coronal polarized spectrum.

After a brief update on the current status of the project, I will present the efforts under way to prepare for the scientific exploitation of the new data during the first two years of operation, the so-called "Critical Science Plan". A number of thematic workshops are currently being held, involving a large fraction of the solar physics community, with the final goal of identifying science topics that can best be addressed with the DKIST capabilities. Many of these science projects, in particular the ones addressing the physics of the outer solar atmosphere, will benefit from coordination with IRIS, exploiting its high spatial and temporal resolution, and complementary diagnostics.

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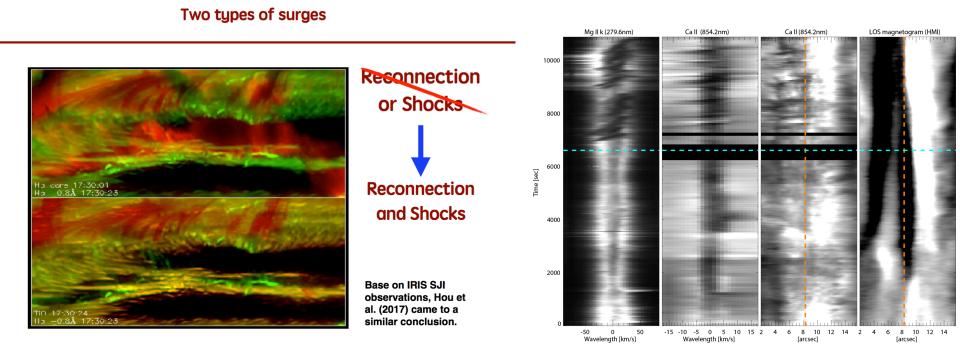


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IRIS and GBOs

Chromosphere/TR diagnostics, dynamics, high fidelity spectra

Polarimetry, high cadence, high resolution



J. Zhang, this meeting

M. Kubo, this meeting

Daniel K Inouye Solar Telescope (DKIST)



It's coming (Haleakala, Maui)

Daniel K Inouye Solar Telescope: The Team

– NSO

DKIST PI:

NSO/AURA

DKIST Director: T. Rimmele



<u>DKIST co-Is:</u> P. Goode, M. Knoelker, J. Kuhn, R. Rosner





Institute for Astronomy University of Hawaii

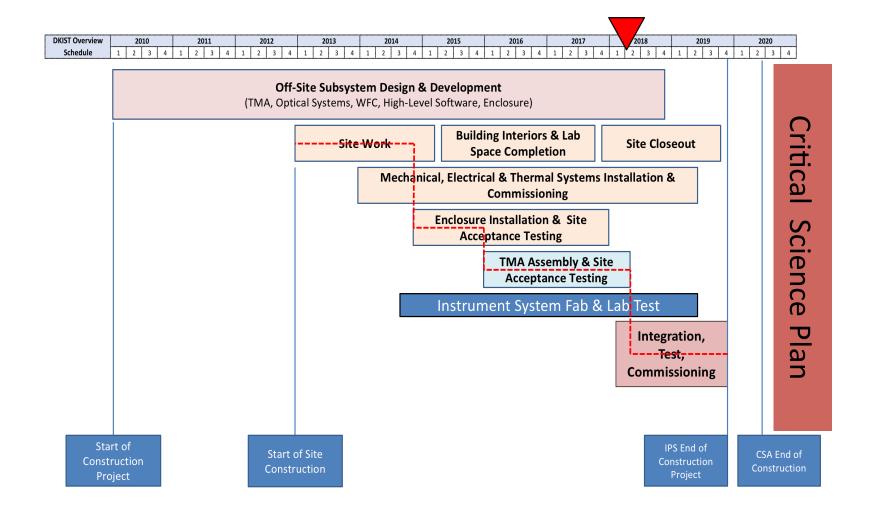


• CRYO-NIRSP PI: J. Kuhn

First light instruments

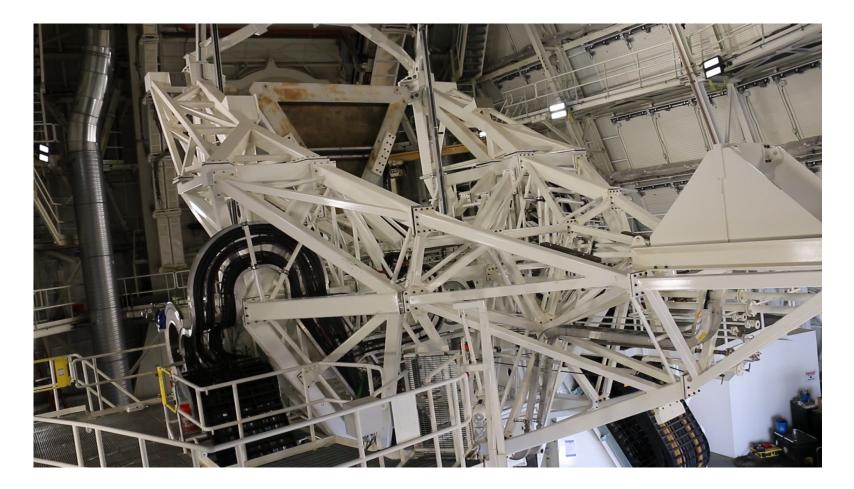
- DL-NIRSP PI: H. Lin
- High Altitude Observatory
 - ViSP PI: R. Casini
- KIS, Germany,
 - VTF PI: O. vd Luehe
- UK DKIST Consortium
 - Visible Detectors,
 - PI: M. Mathioudakis, QUB

The DKIST: status and timeline



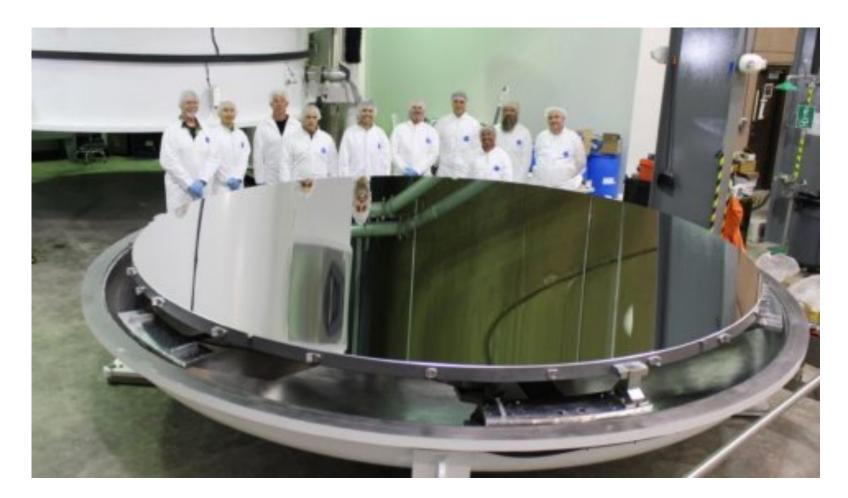
8 years of construction; 85% complete

Telescope mount



Science first light: mid 2020

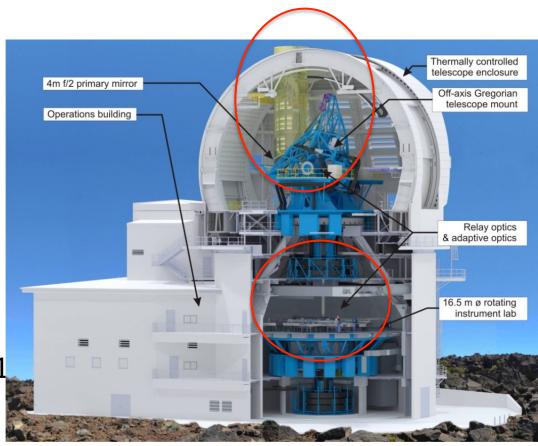
Telescope mirror



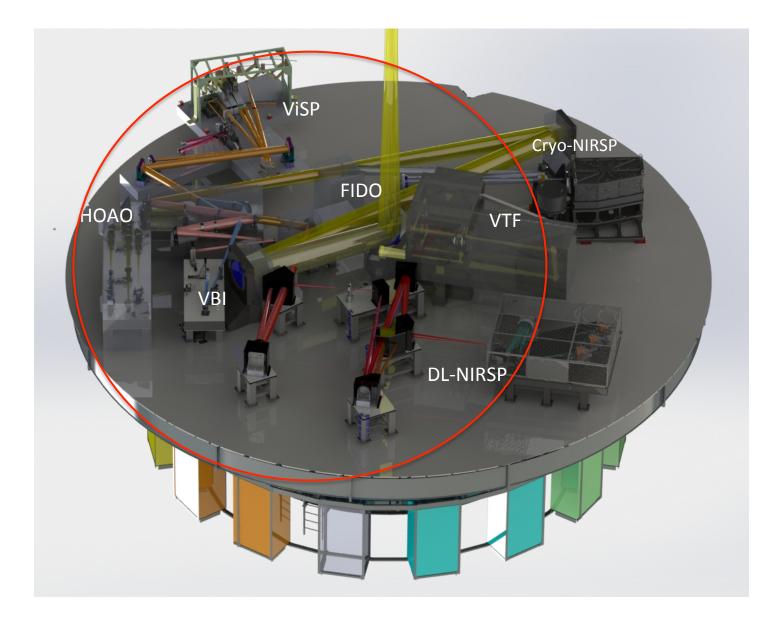
Science first light: mid 2020

DKIST: a transformational facility

- <u>4-meter aperture</u>, f/2
- Alt/az mount
- All reflecting optics
- Off-axis design
- Low-scattered light
 - Coronagraph
 - Lyot stop & limb occulter
 - In situ clean & wash of M1
- Integrated adaptive optics
- High-precision polarimetry



First light instruments



DKIST Instrument Suite Overview

Instrument Name	Acronym	Wavelength Range	Analogs	
Visible Broadband Imager	VBI (blue, red)	390 – 550 nm 600 – 860 nm	Hinode/BFI; ROSA High cadence, high spatial res.	
Visible Spectro-Polarimeter	ViSP	380 – 900 nm	SPINOR, Hinode/SP, IRIS Scanning spectrograph, high spectral fidelity	
Diffraction-Limited Near IR Spectro-Polarimeter	DL-NIRSP	500 – 900 nm 900 – 1350 nm 1350 – 1800 nm	SPIES, Image Slicer @ GREGOR True IFU, variable spatial resolution / FOV	
Visible Tunable Filter	VTF	520 – 870 nm (590 – 870 nm)	IBIS, CRISP, GFPI, HMI Imaging spectro-polarimeter	
Cryogenic Near IR Spectro- Polarimeter (with context imager)	Cryo-NIRSP	1000 – 5000 nm	CYRA (BBSO) Cryogenic, scanning spectrograph, novel IR diagnostics	

ALL instruments (but VBI) are meant to perform polarimetric obs.

What is so special about DKIST

<u>4 meter diameter:</u>

- super-high resolution (~20 km @ 400 nm)
- "light bucket" -> high cadence
- "light bucket" -> high precision polarimetry (goal: S/N = 10^4 in Q/I)

All reflective:

• access to new IR diagnostics (mostly coronal; but not only)

<u>Off-axis:</u>

• low scattered light (coronal physics; but also sunspots!)

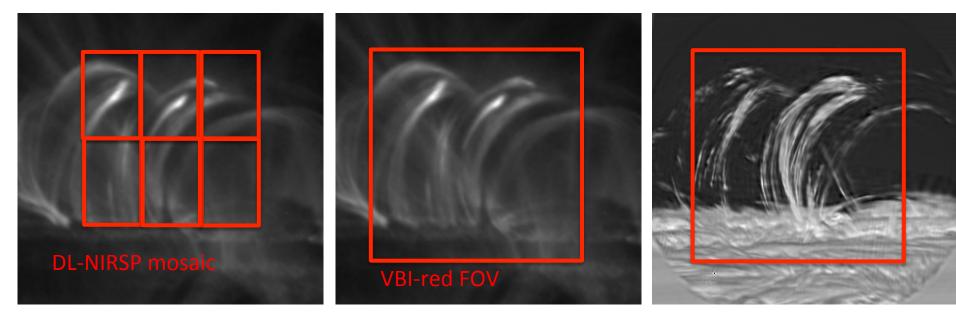
Multi-instrument:

• multi-height diagnostics

Science ideas ("use cases")

Collective behavior of spicules? Coronal rain? (T. Van Doorsselaere)

Limb observations with VBI blu (Call K) + red channel (FeXI 789.2; Halpha), + DL-NIRSP chromospheric (coronal?) magnetic maps 0.05"- 1" resolution, few s – 2 min cadence



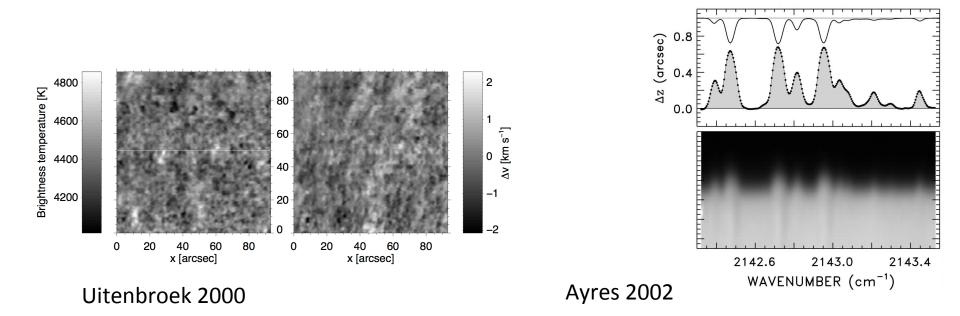
SDO/AIA 171 – 900 km resolution

IBIS H α – 130 km resolution

Science ideas ("use cases")

Cold bubbles in the chromosphere? (J. Martinez Sykora, A. Sukhorukov)

Cryo-NIRSP observations ON DISK, of CO band at 4.7 μ m (Te < 3700 K). Resolution = 0.3"-0.5"





<u>Aim:</u> By science first light, to be ready, as a community, to execute a set of observations exploiting the DKIST capabilities to address critical, compelling science.

Planned: two years (mid 2020, mid 2022)

<u>How (1)</u>: Bottom-up approach, community based. NSO is actively engaging the community, informing it of the DKIST capabilities and offering practical tools to develop "proto-proposals".

CSP Workshops



DKIST Critical Science Plan Workshops

The National Solar Observatory is currently supporting a number of Critical Science Plan Development workshops, whose main purpose is to advance the DKIST Critical Science Plan (CSP) in preparation for start of operations. Each CSP Development Workshops is expected to advance Science Use Case development by the user community by focusing on a particular research area.

- Workshops science
- Workshops Logistics
- DKIST web pages
- DKIST in a nutshel
- DKIST instruments
- DKIST Critical Science Plan

The DKIST Critical Science Plan (CSP) Nine CSP Workshops: https://www.nso.edu/cspw



CSP Workshops



9 thematic workshops
8 US Universities
3 Foreign Institutes
200+ scientists
200+ "science use cases"



<u>How</u> (2): Compile an extensive set of complementary Science Use Cases, detailing the topics to be investigated, and the type of DKIST observations necessary to address the science.

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<u>How</u> (2): Compile an extensive set of complementary Science Use Cases, detailing the topics to be investigated, and the type of DKIST observations necessary to address the science.



- Updated the Filter Sequence & Spectral Lines
- Updated the Total Time of Observation
- Changed the Summary to 'Changing coronal magnetic fields over limb active regions with ViSP, DLNIRSP and Cryo-NIRSP'

5 hours ago Comment Vote Watch

Mari Paz Miralles updated 4 fields of UC-211 - DKIST Pseudostreamer Study: Magnetic Topology, Plasma Properties, Magnetic Reconnection

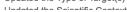
- Updated the Abstract
- Updated the Program Type
- Updated the Scientific Context
- Updated the Observing Coordination

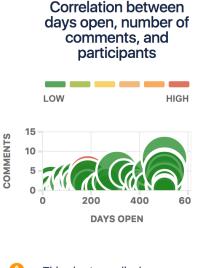
11 hours
ago Comment Vote Watch



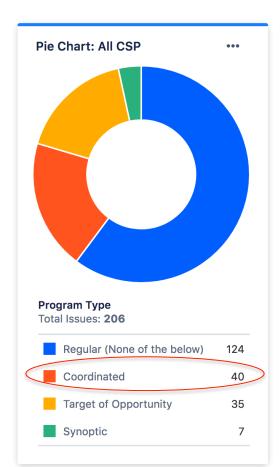
Sarah.Gibson updated 19 fields of UC-209 - DKIST prominence cavity study

- · Updated the Program Type
- Updated the Instrument Set Definition
- Updated the Type of Target(s)





This chart can display only the first 200 matching issues. Refine your search to return fewer issues.



<u>Feedback</u>: The Science Working Group (DKIST SWG) will provide feedback to PIs, including suggestions for team consolidation, comments on best use of DKIST capabilities, etc. Final goal is to articulate the community vision for essential DKIST science in the early years.

Warning:

The CSP Science Use Cases are NOT proposals, and the Science Working Group is NOT a Time Allocation Committee.

Useful links; email

- <u>https://dkist.nso.edu/CSP</u>
- https://dkist.nso.edu/CSP/instruments
- https://www.nso.edu/cspw.php
- https://nso-atst.atlassian.net
- mark.rast@lasp.colorado.edu; gcauzzi@nso.edu

Thanks !



http://dkist.nso.edu/

