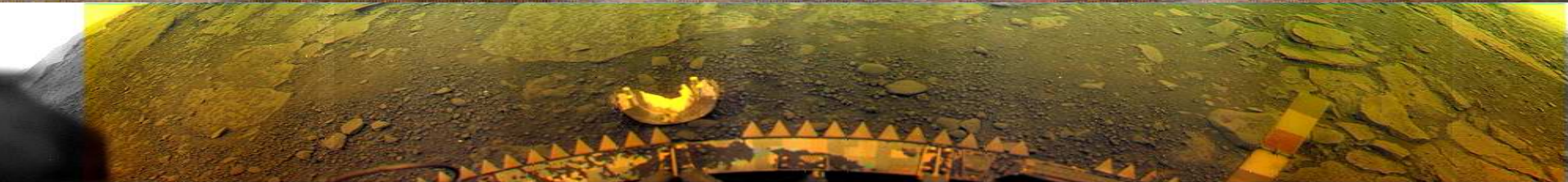






RADIATIVE ENERGY BALANCE IN THE PLANETARY ATMOSPHERES

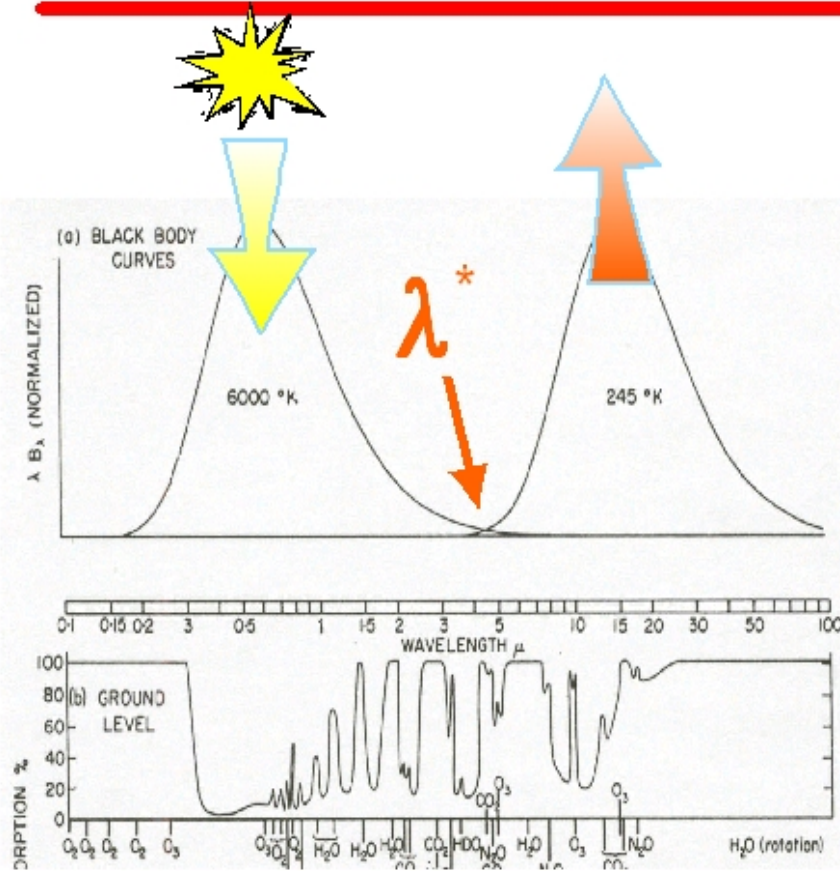


Outline of the talk

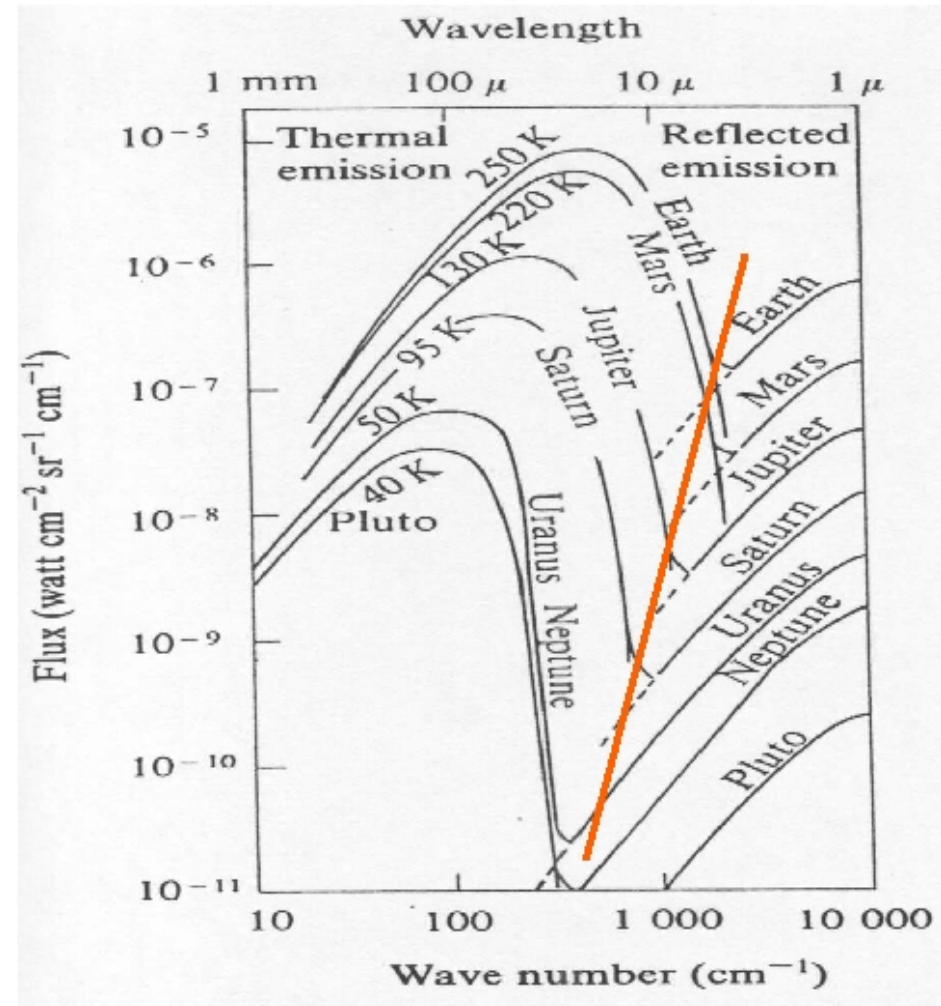
-  ***radiative energy budget***
-  ***forcing of the general circulation***
-  ***greenhouse effect***
-  ***balance of entropy***

Radiative energy budget

Energy balance of planets



Planetary spectra



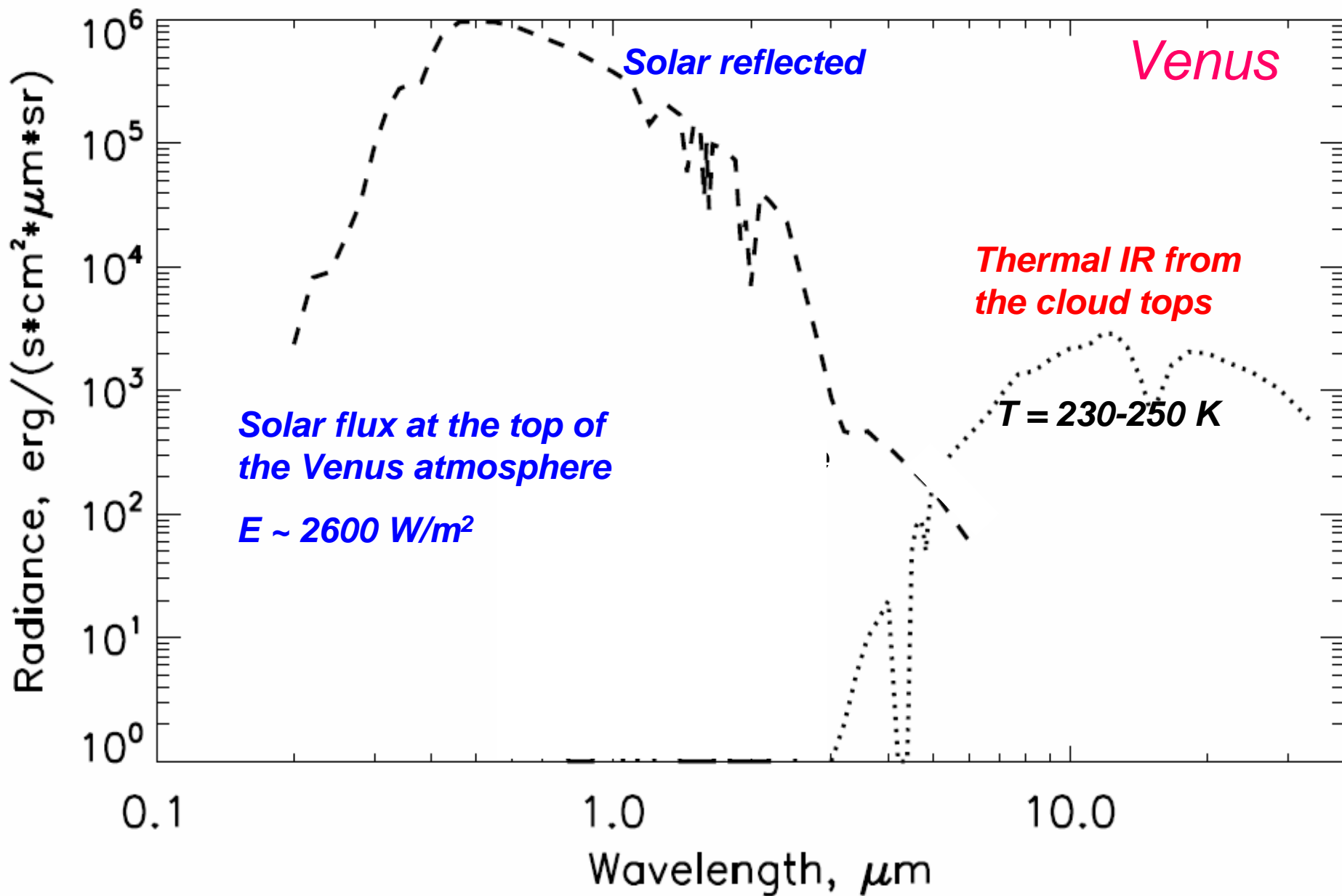
Effective temperature

$$E(1-A) = 4\sigma T_{\text{eff}}^4$$

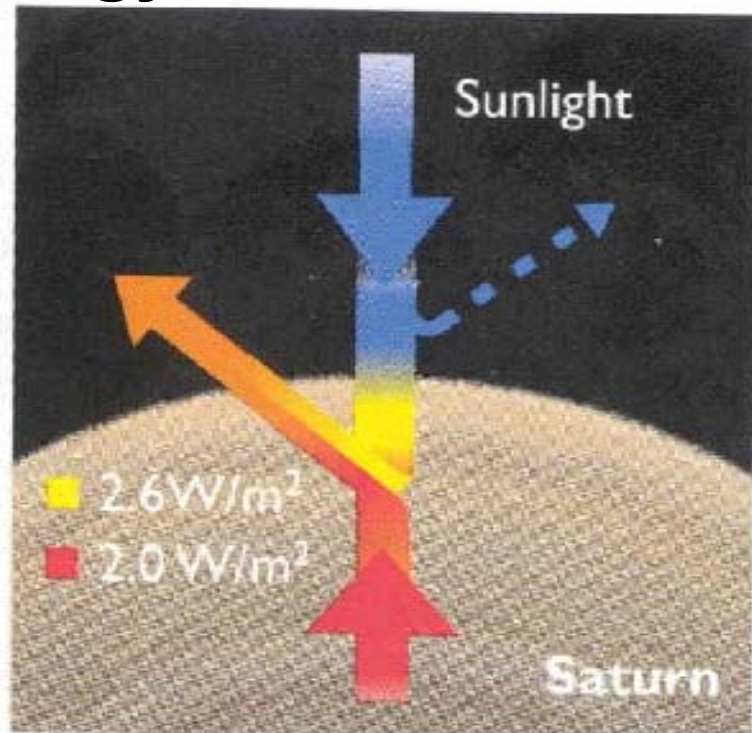
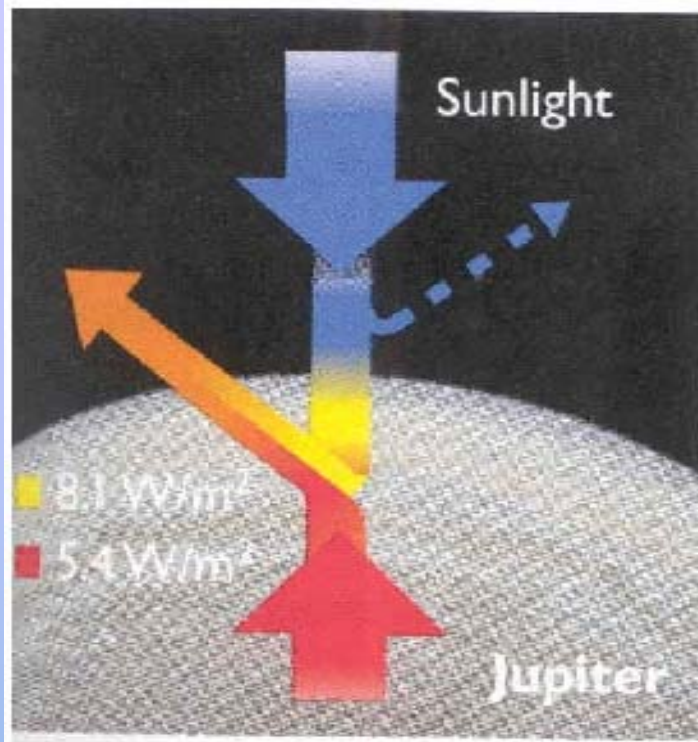
Solar constant

$$E_{\oplus} = 1370 \text{ W/m}^2$$

Composite spectrum of the outgoing radiation

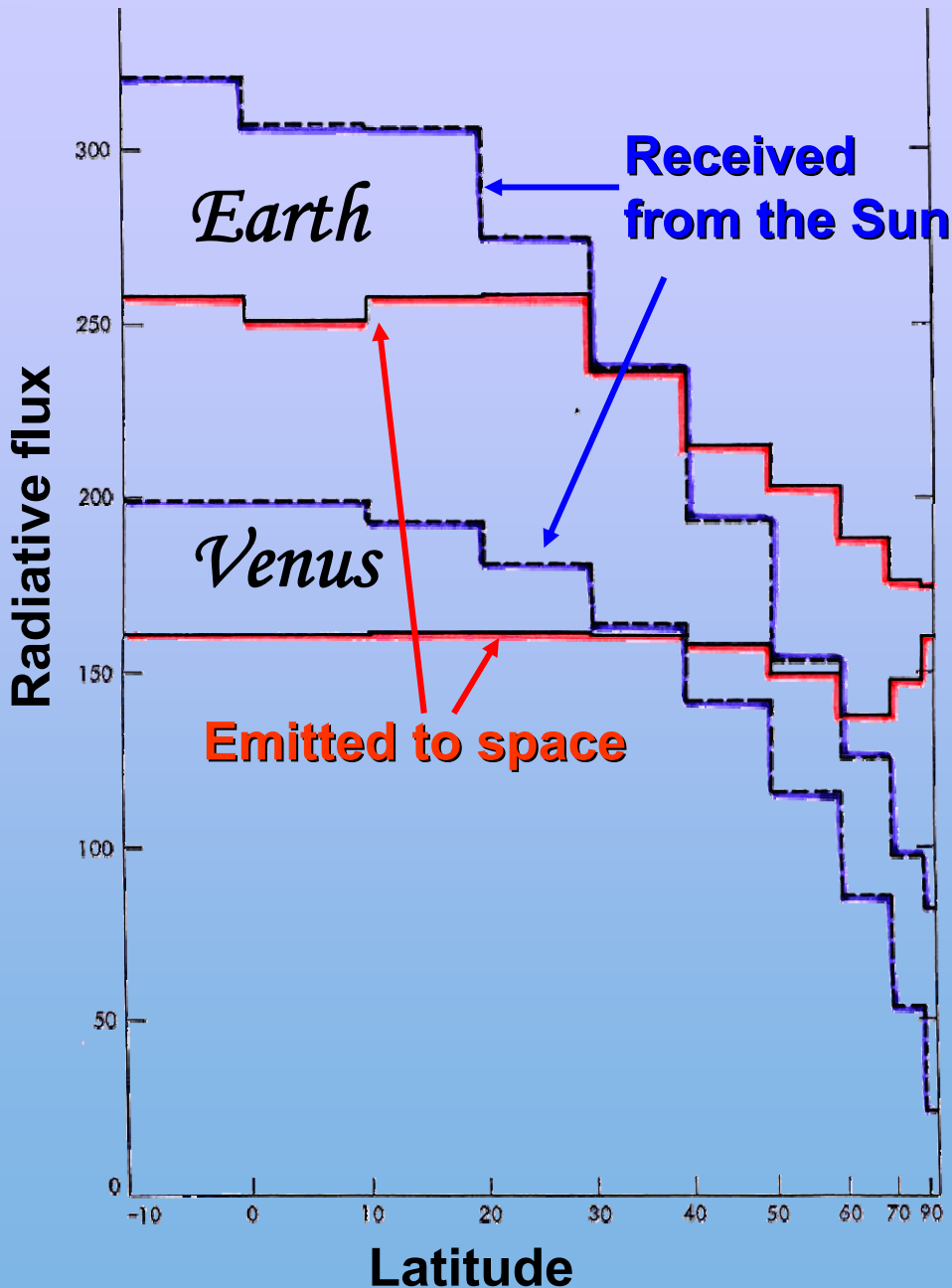


Sources of energy on Giants



	Earth	Jupiter	Saturn	Uranus	Neptune
D, au	1	5.2	9.56	19.22	30.11
Solar flux, W/m ²	1370	50.66	15	3.7	1.51
Energy balance, $F_{\text{IR}} / F_{\text{abs}}$	1	1.67	1.78	1.06	2.61

Latitudinal distribution of energy sources and sinks

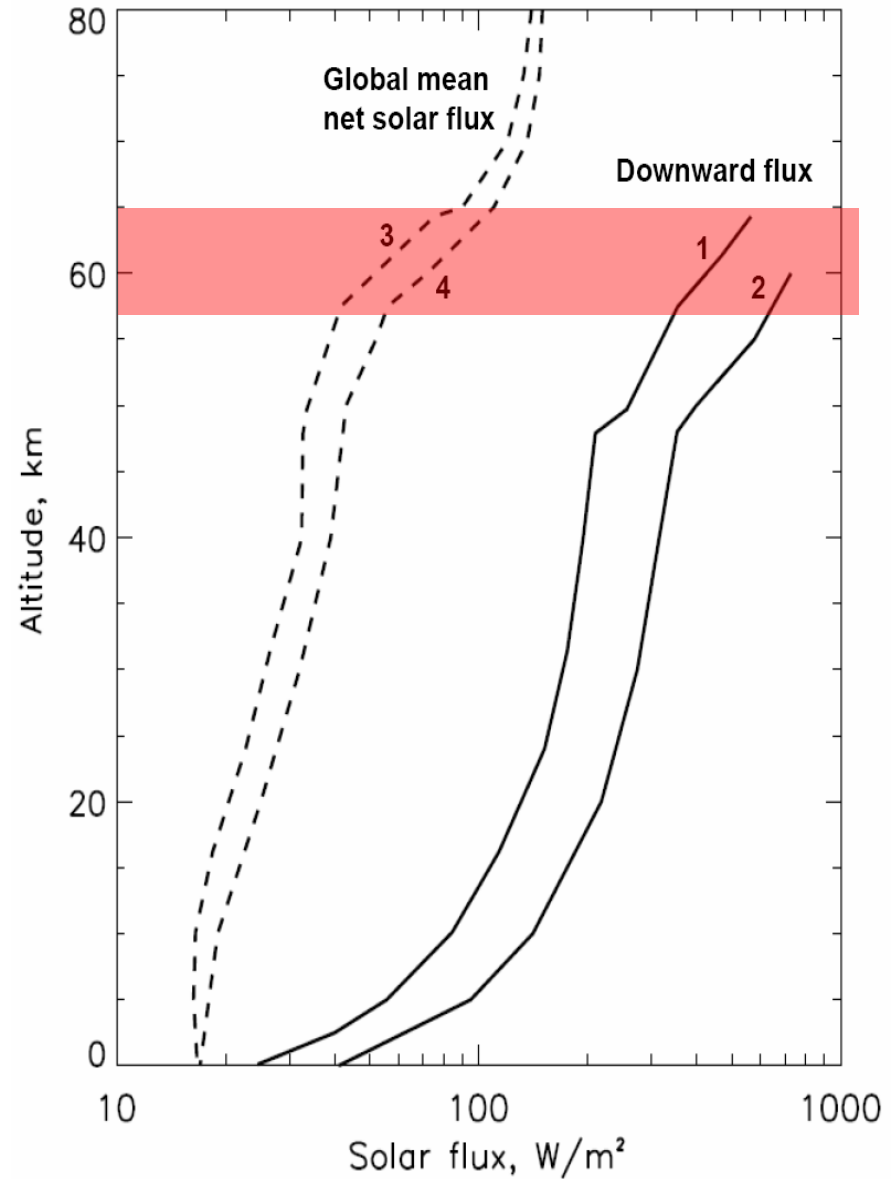
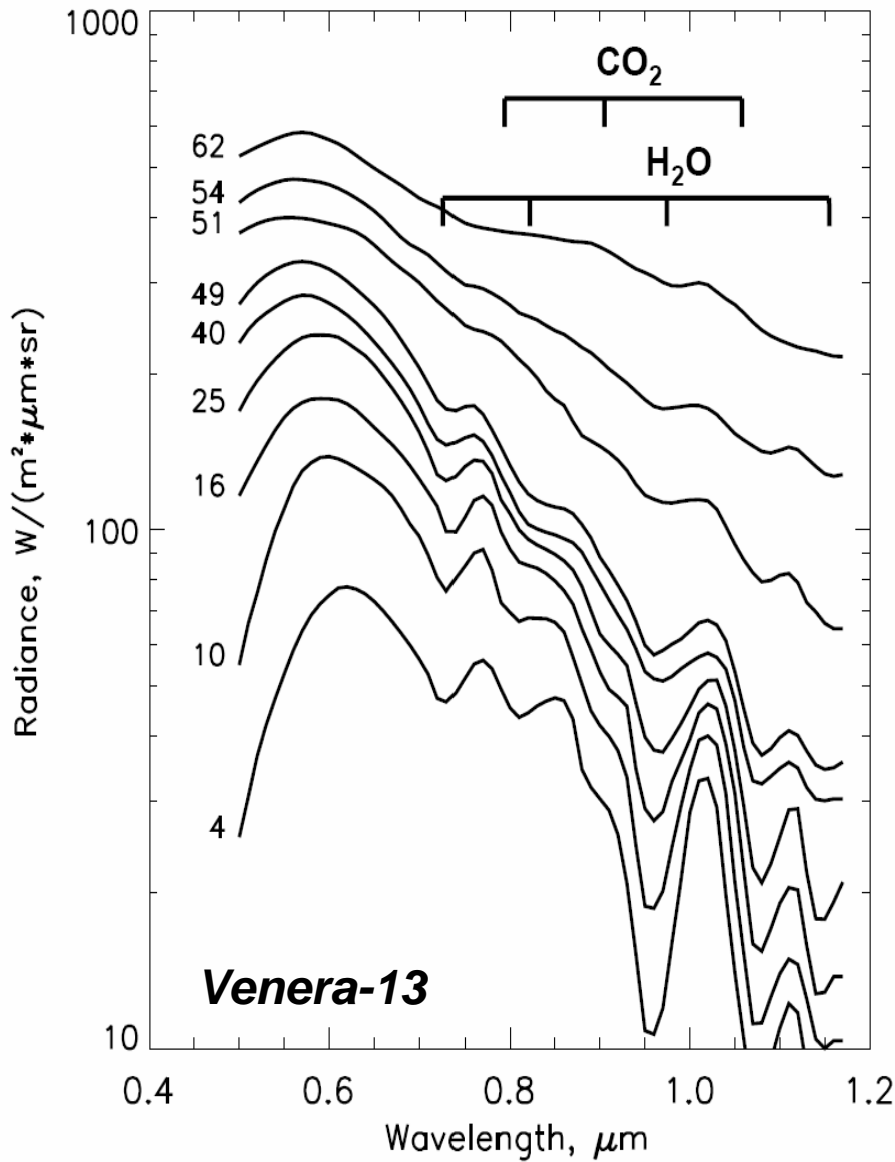


✚ *Venus gets less energy than the Earth !*

✚ *Net heating at equator, net cooling on poles*

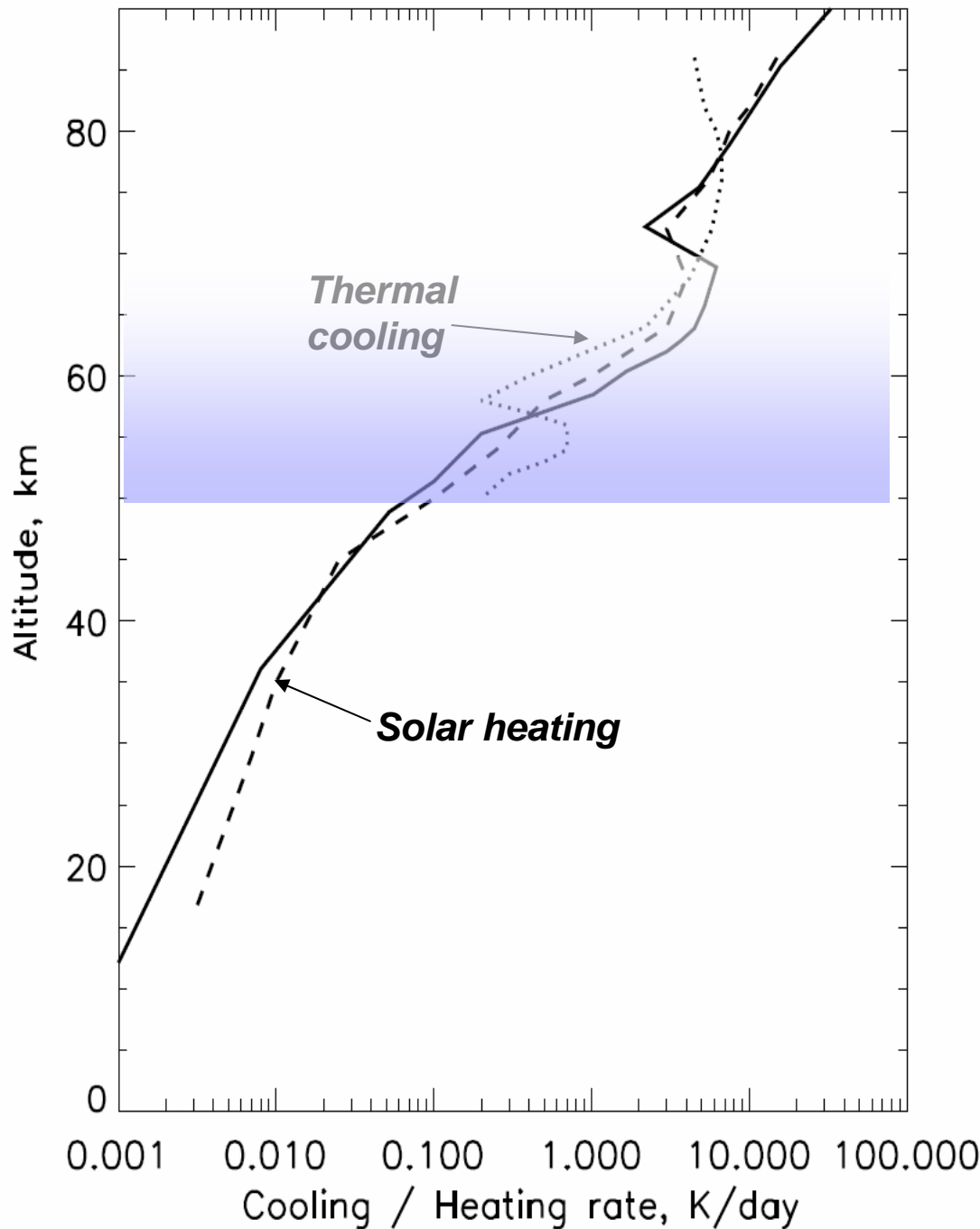
✚ *Latitudinal distribution of radiative balance implies energy transport by circulation*

Vertical distribution of deposited solar energy



Ekonomov et al., 1983

Global mean heating and cooling rates



✚ *half of solar energy deposited on Venus is absorbed by the unknown UV absorber in the cloud layer*

Tomasko et al., 1985

Crisp & Titov, 1997

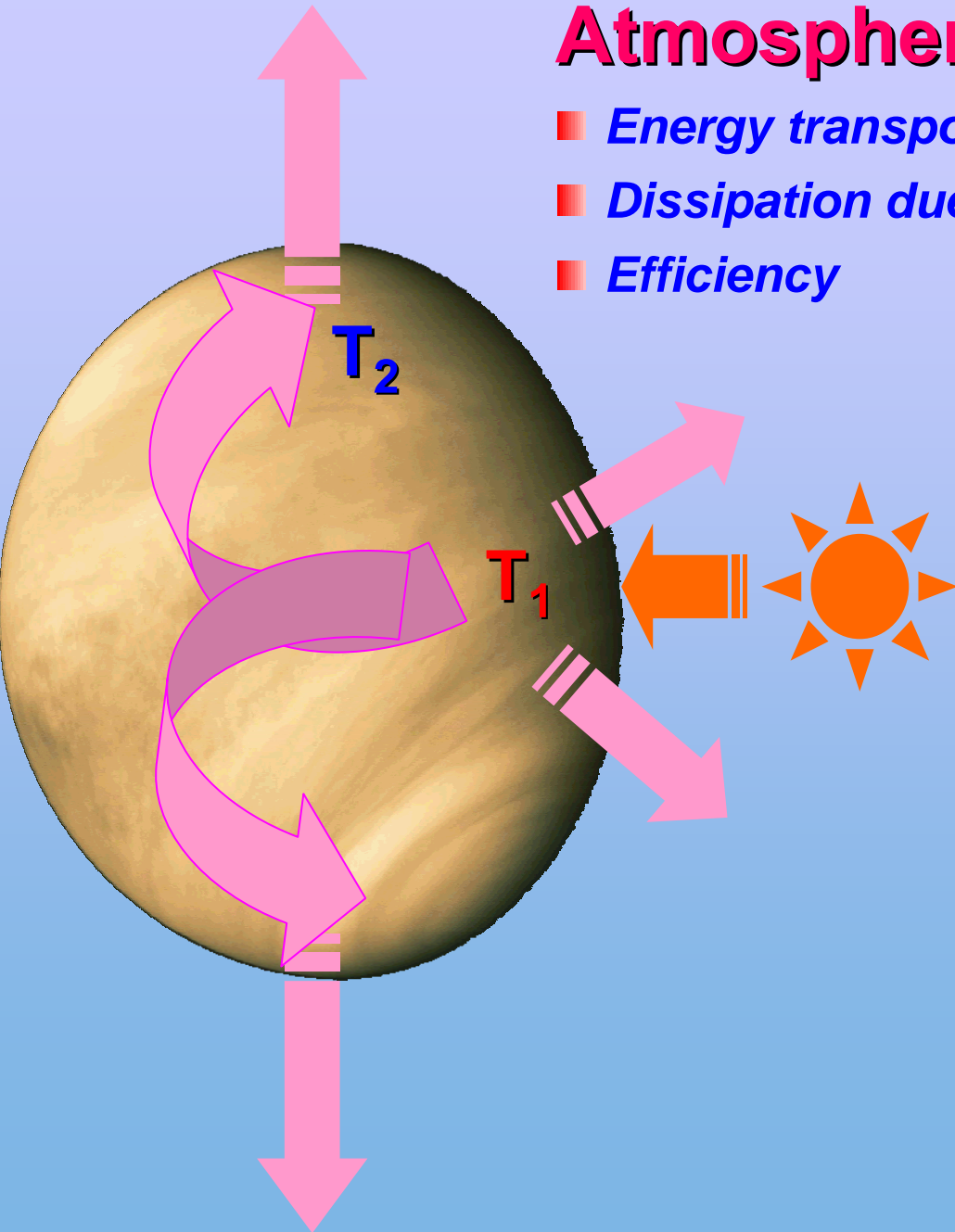
Mean deposition of solar energy on terrestrial planets [W/m²]

	Venus	Earth	Mars
Atmosphere	130	70	~0
Surface	20	170	125

Atmospheres as heat engines

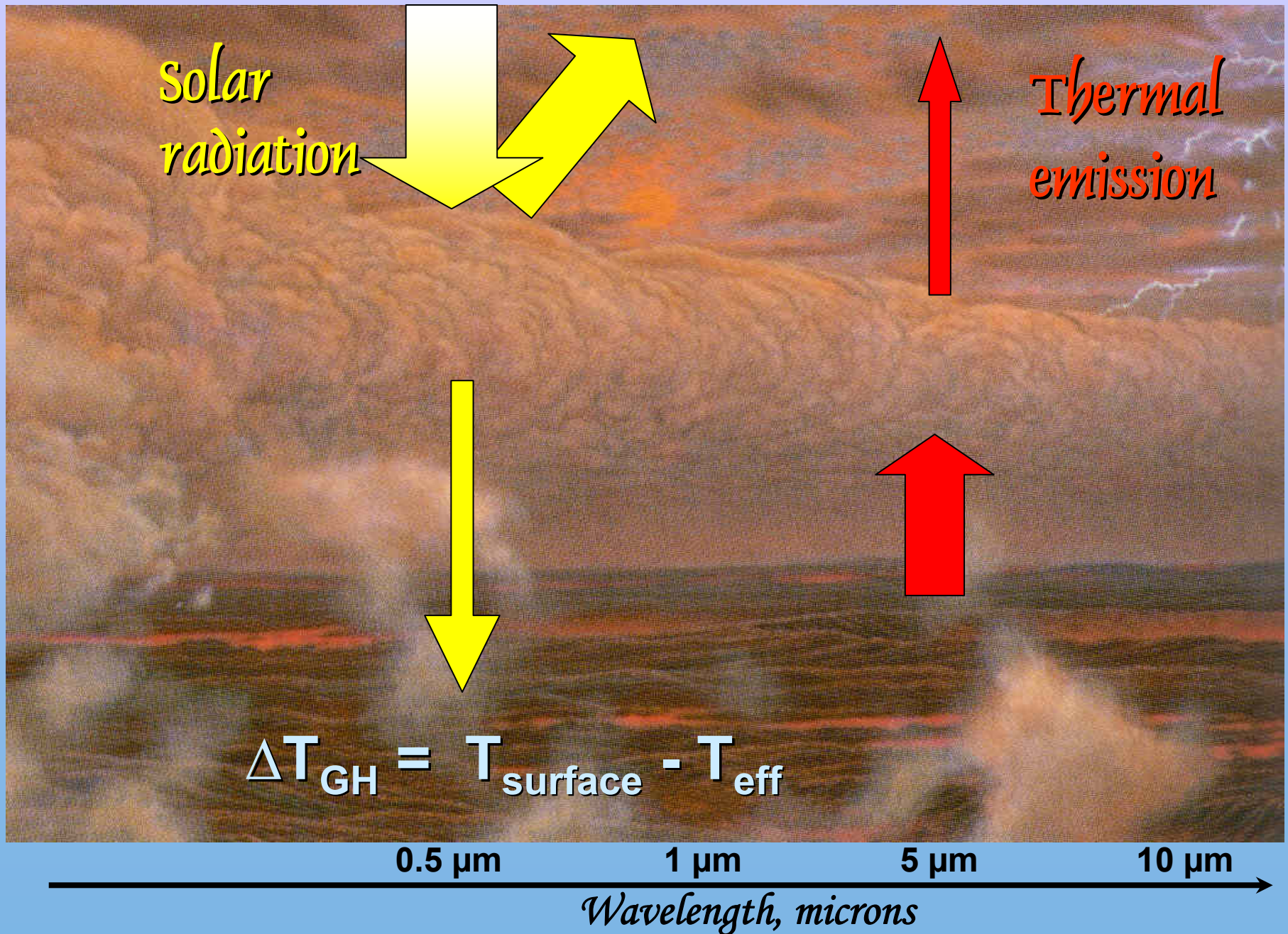
- *Energy transport by atmospheric motions*
- *Dissipation due to friction*
- *Efficiency*

$$\varepsilon \leq 1 - \frac{T_2}{T_1}$$

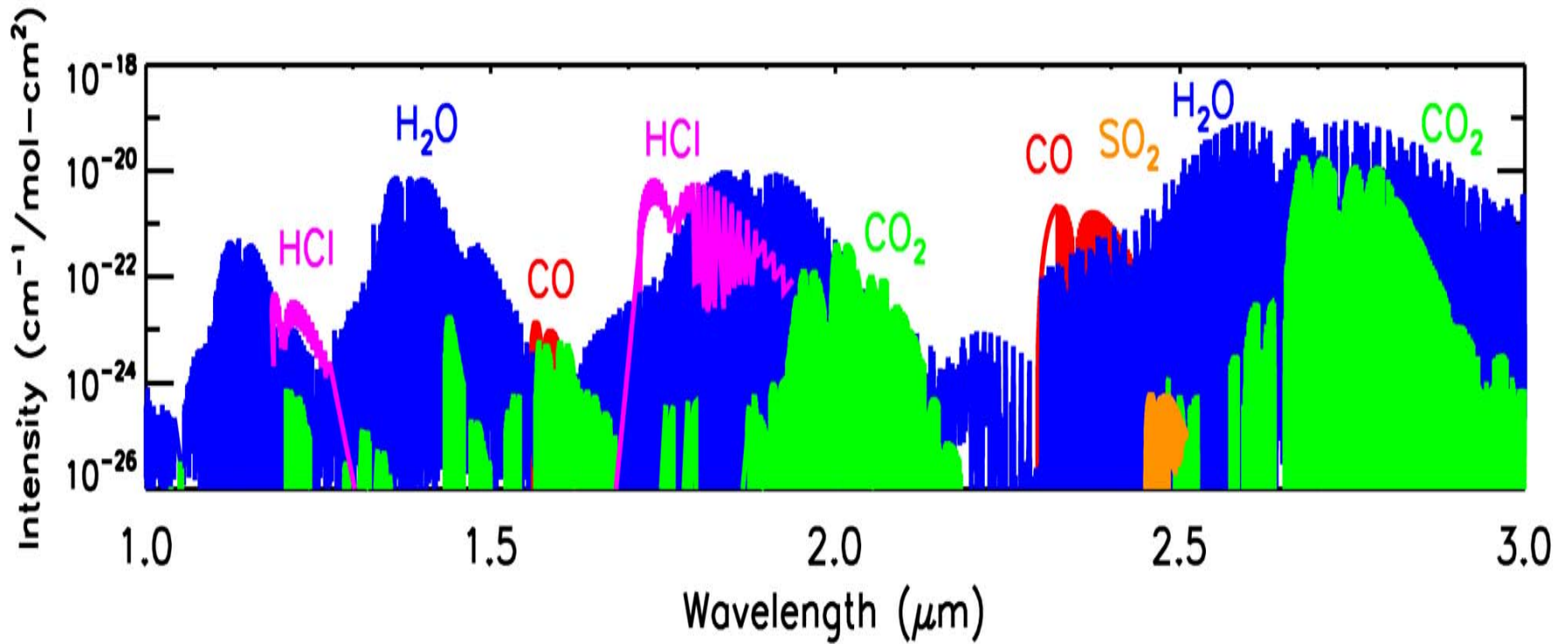


Greenhouse effect

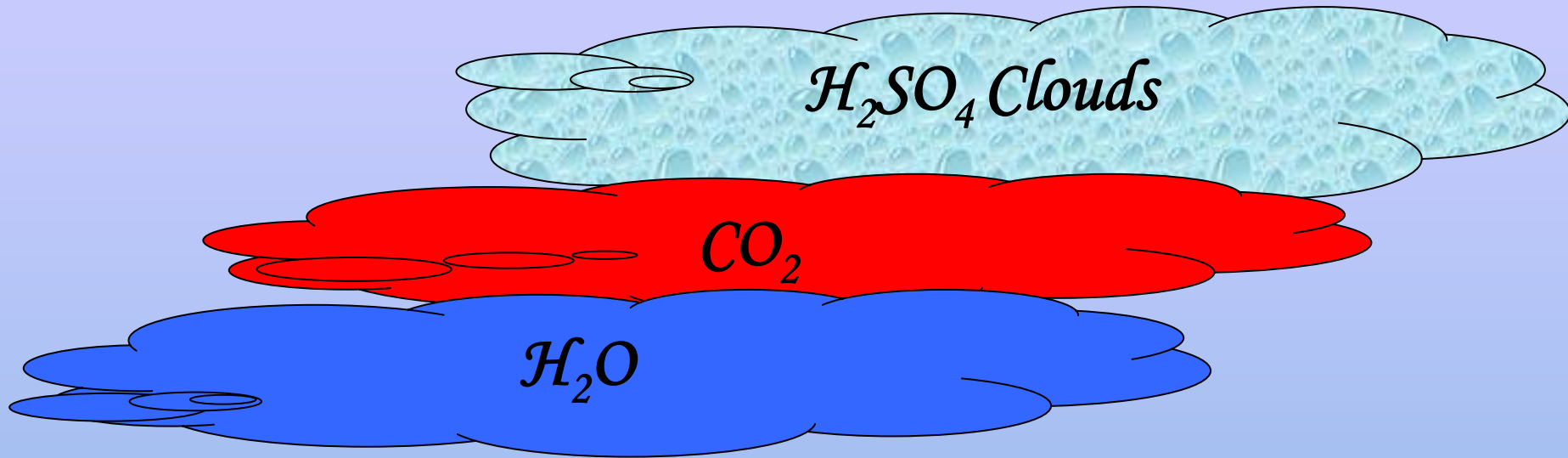
Basics of the greenhouse effect



Absorption bands of the atmospheric gases



Venus – the queen of greenhouse



$CO_2 - 420K$

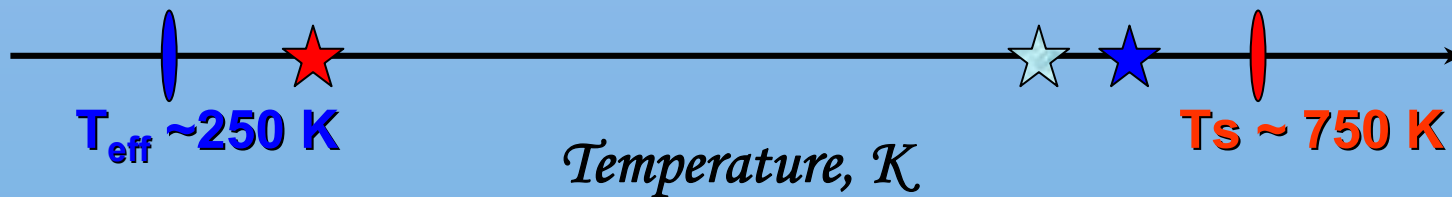
$CO_2 - 12K$

$H_2O - 70K$

$CO - 3K$

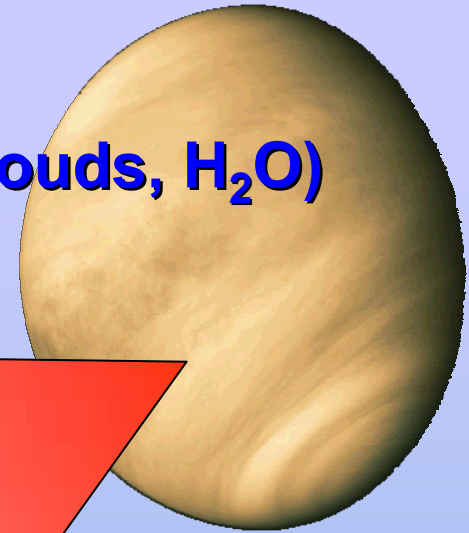
$Clouds - 140K$

$SO_2 - 3K$

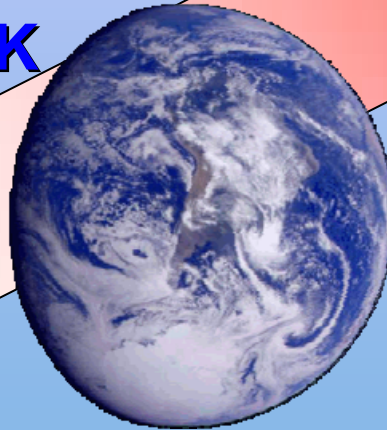


Greenhouse effect on the terrestrial planets

Venus (CO₂, clouds, H₂O)
 $\Delta T \sim 500K$



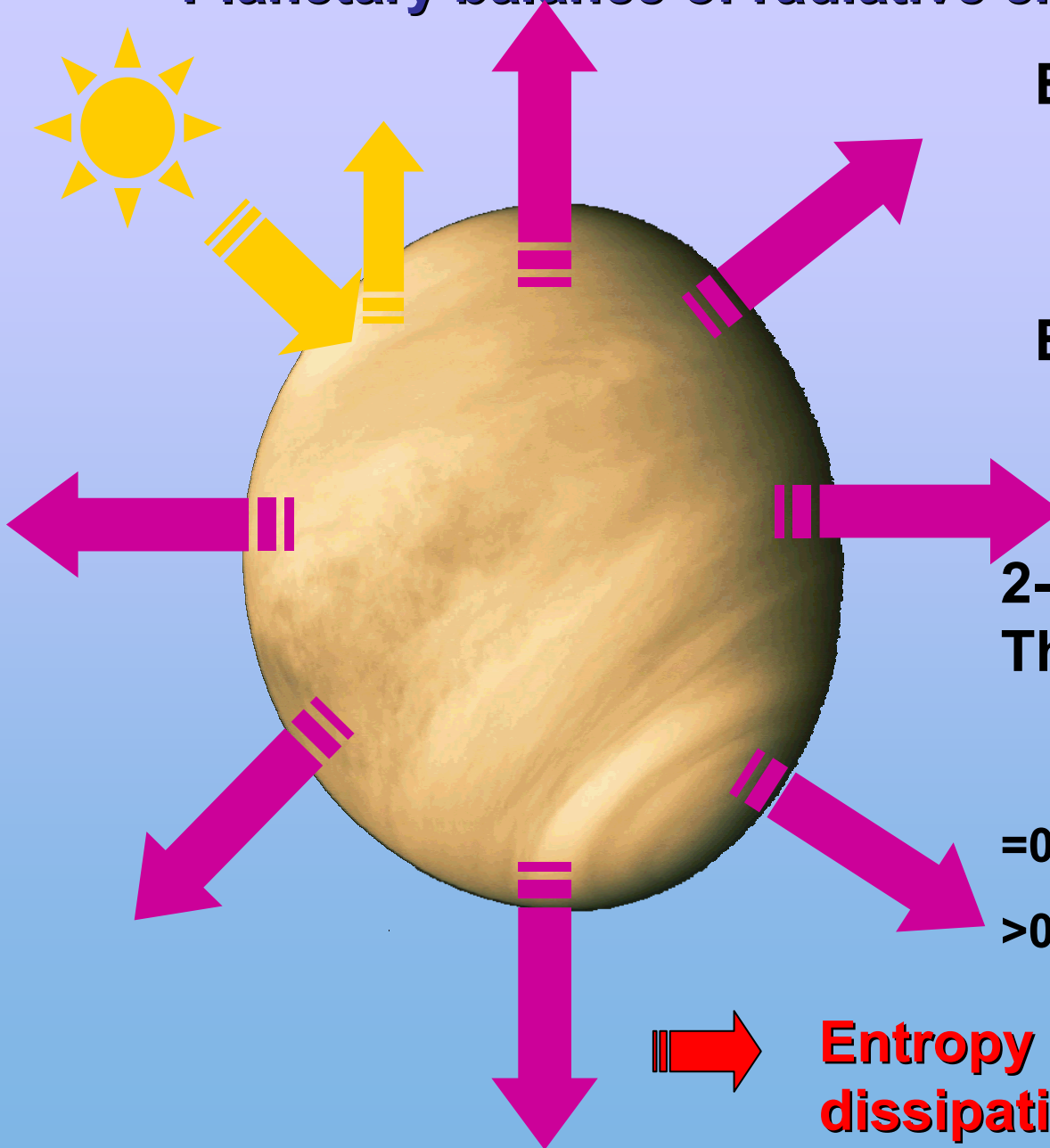
Earth (H₂O, clouds)
 $\Delta T \sim 40K$



Mars (dust, CO₂)
 $\Delta T \sim 5K$

Entropy balance

Planetary balance of radiative energy and entropy



Energy balance:

$$E_{\text{Solar}} - E_{\text{ThIR}} = 0$$

Entropy:

$$\Delta S = E/T$$

2-d Law of
Thermodynamics:

$$\Delta S \geq 0$$

=0 - reversible processes

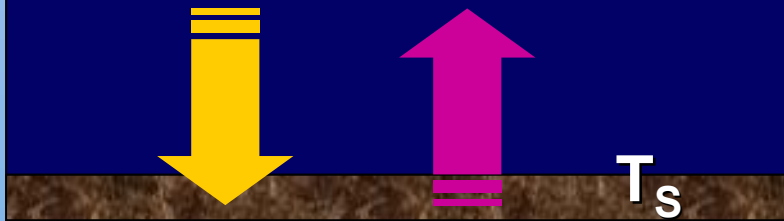
>0 - irreversible processes

**Entropy is a measure of
dissipative processes**

Flux of radiative entropy

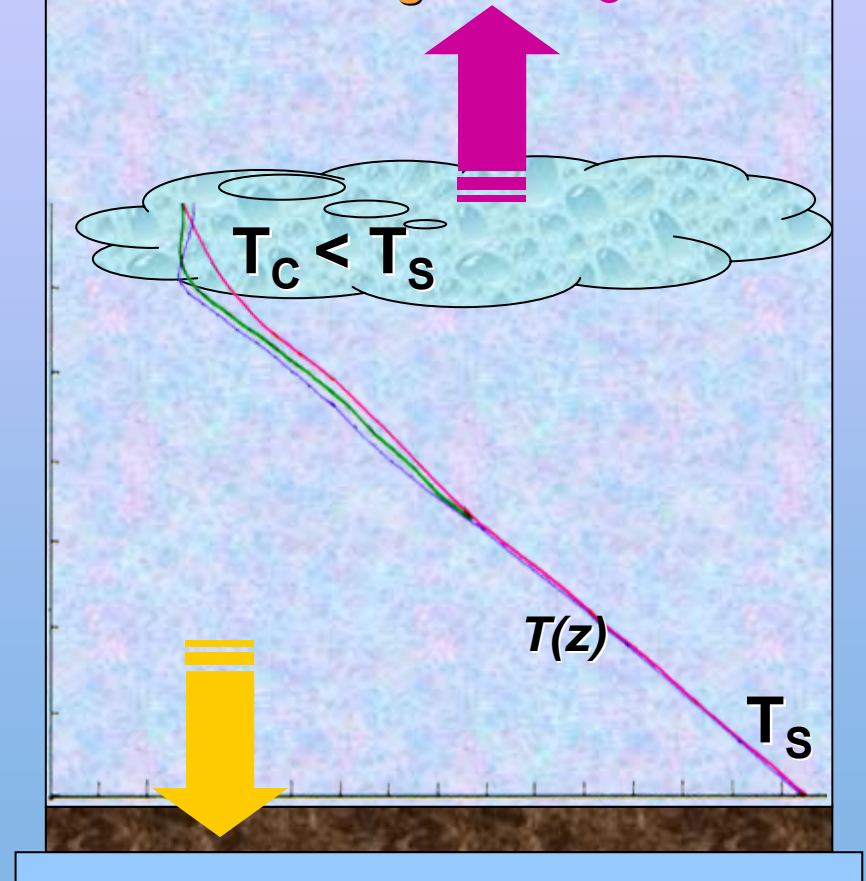
No atmosphere

$$\Delta S = E/T_s - E/T_s = 0$$



Dense atmosphere

$$\Delta S = E/T_s - E/T_c < 0$$



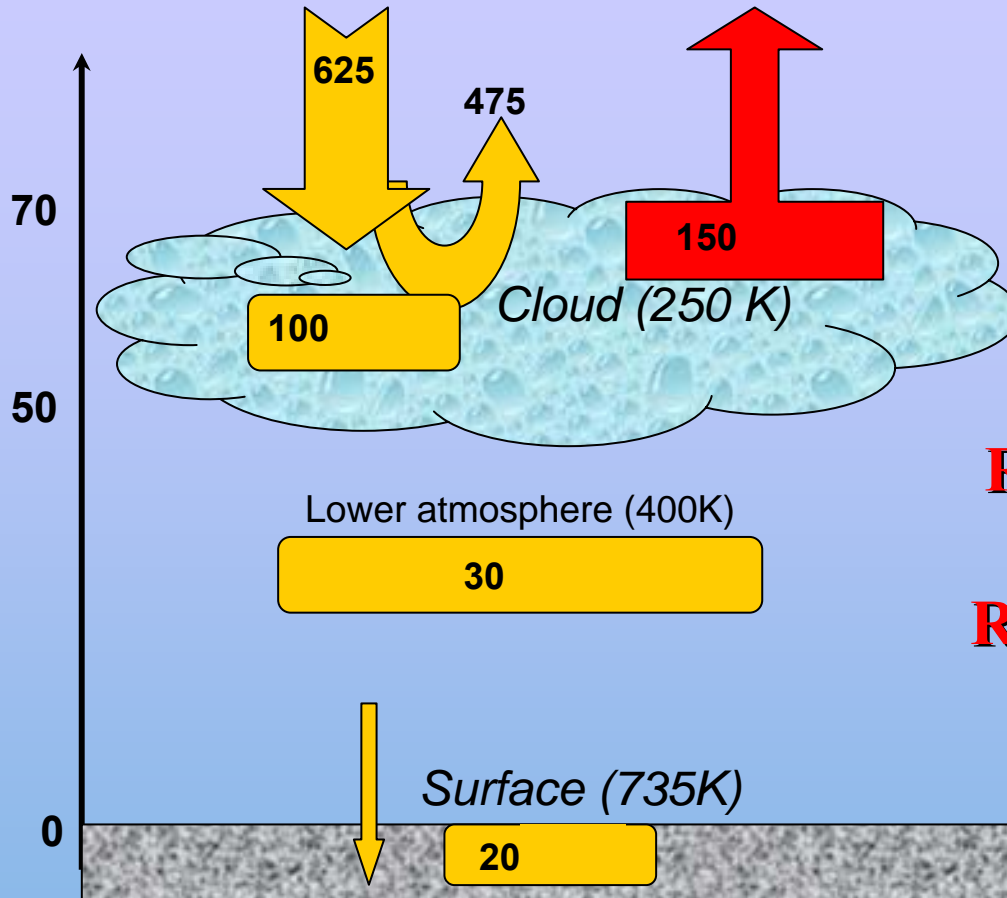
Planets receive negative entropy from the Sun

Solar

Thermal

Radiative

Energy / Entropy balance on Venus



Radiative energy balance

$$\Delta E \approx 0$$

Radiative entropy balance

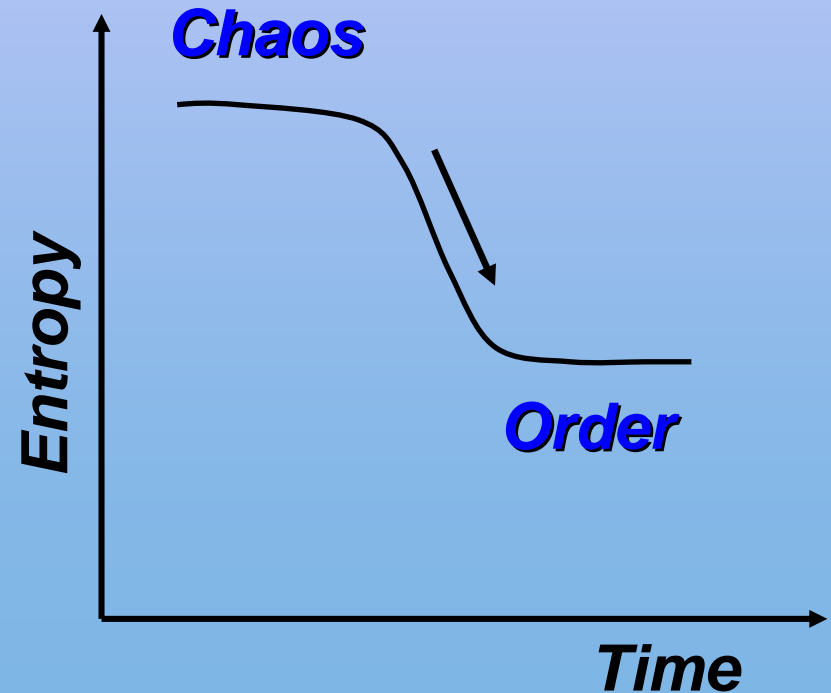
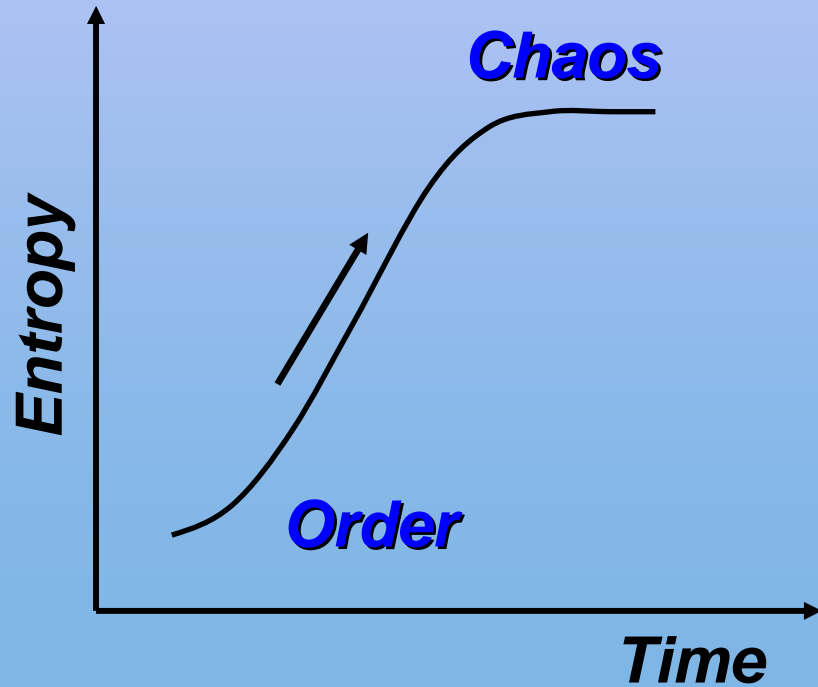
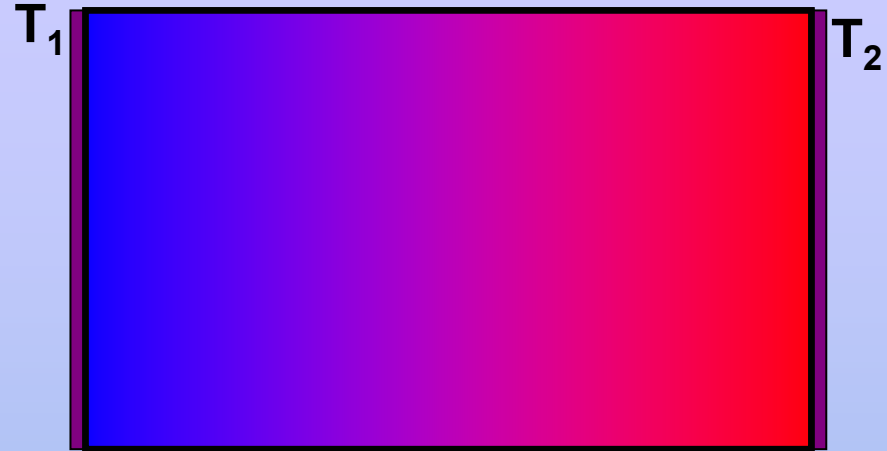
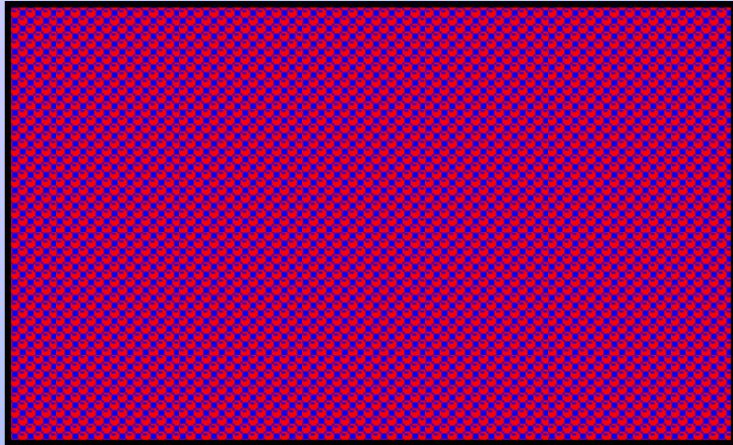
$$\Delta S \approx -100 \text{ mW/m}^2/\text{K}$$

Entropy balance on Earth and Venus

	Earth (Goody, 2000)	Venus
Net radiative sink	-70	-100
Moist convection	+55	0
Mechanical dissipation	+12	~1
Net balance	-3	-100

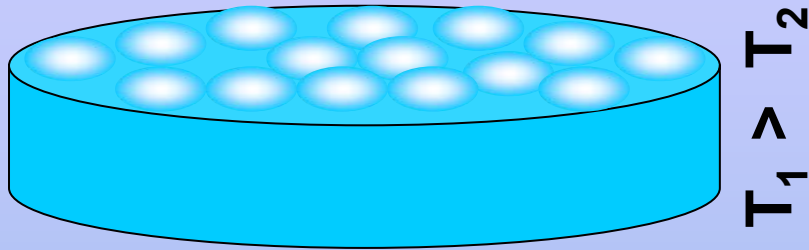
Dissipative processes in the Venus atmosphere - ????

Equilibrium and non-equilibrium thermodynamics



Non-equilibrium dissipative systems

Benard convection



- ✚ critical temperature gradient
- ✚ high level of order
- ✚ high entropy production

