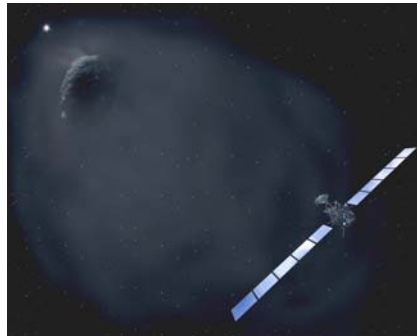


Rosetta - An example for a modern planetary mission

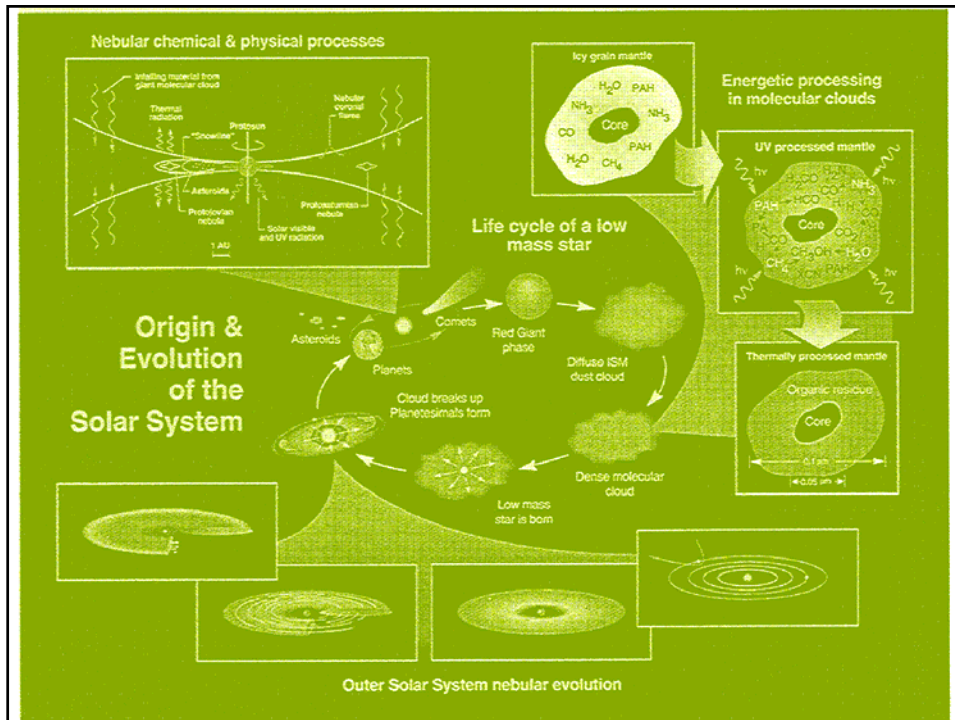
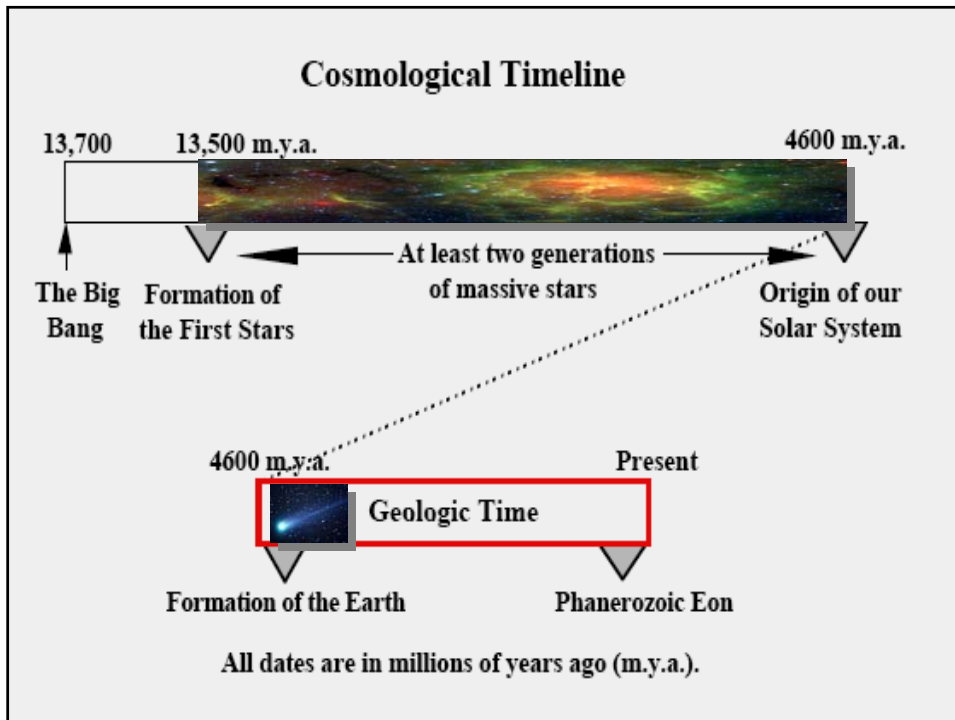
M. Hilchenbach

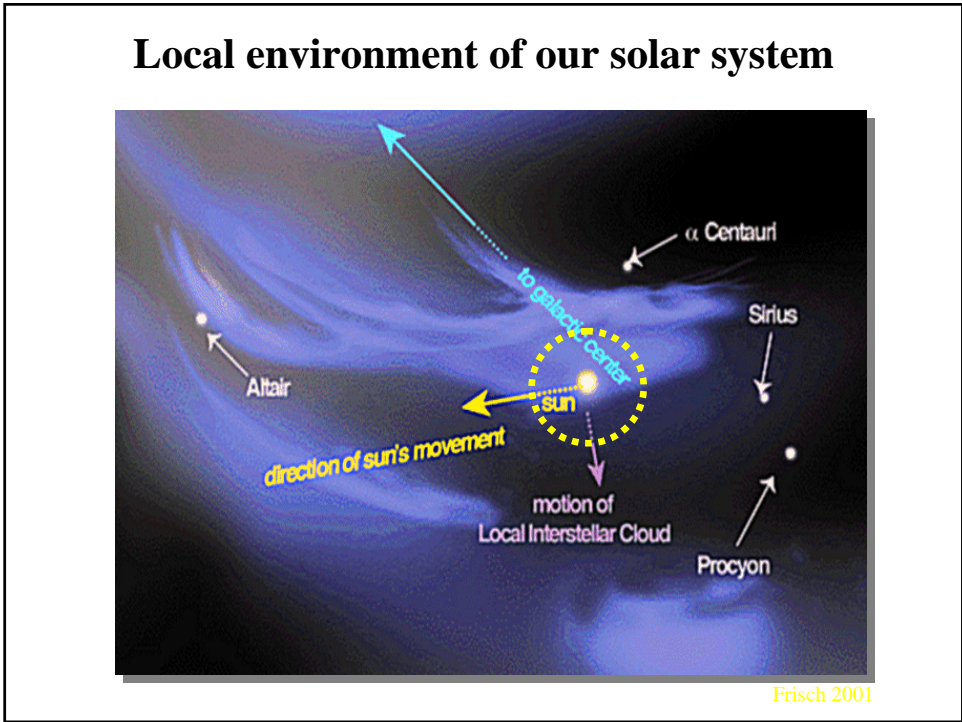
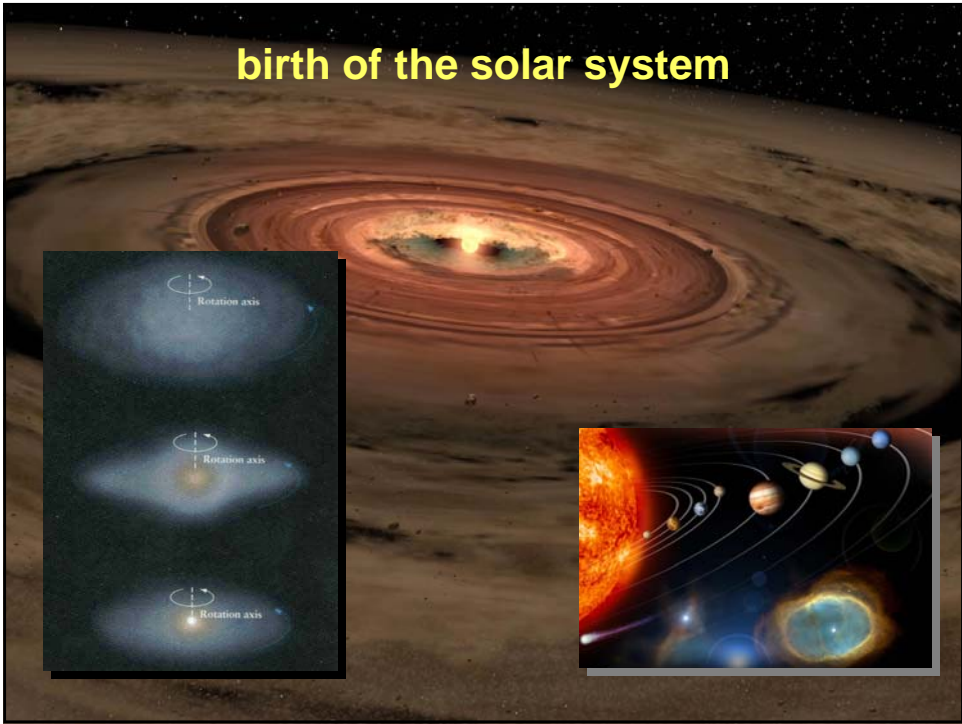
Max-Planck-Institut für Sonnensystemforschung
37191 Katlenburg-Lindau, Germany



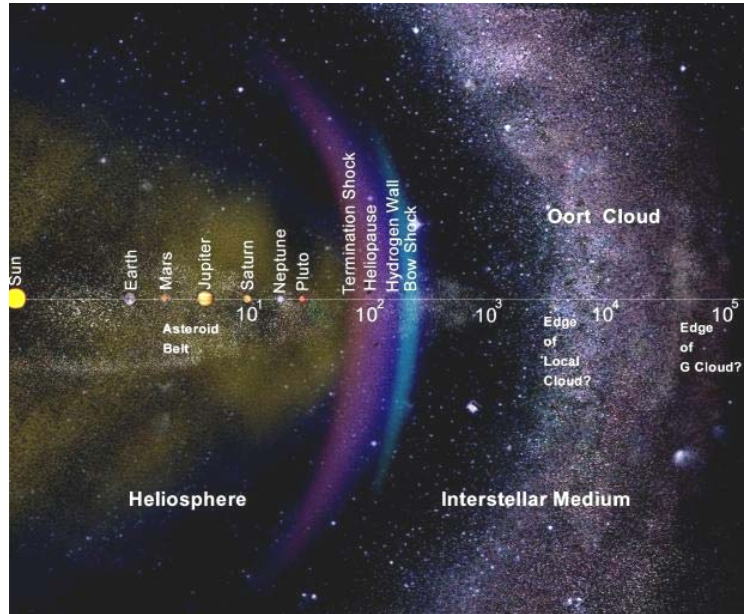
Outline

- timeline, matter and comets
- schedule and management approach for new missions
- Rosetta mission and payload
- first results from Rosetta science-phase

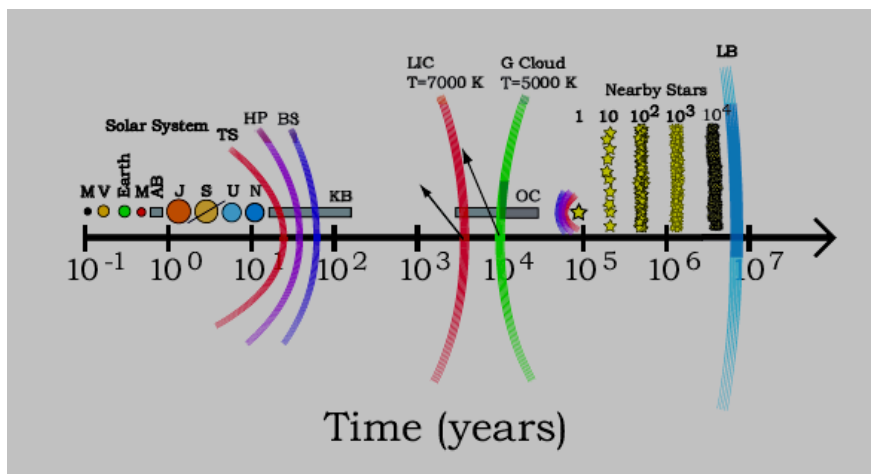




Local solar system regions



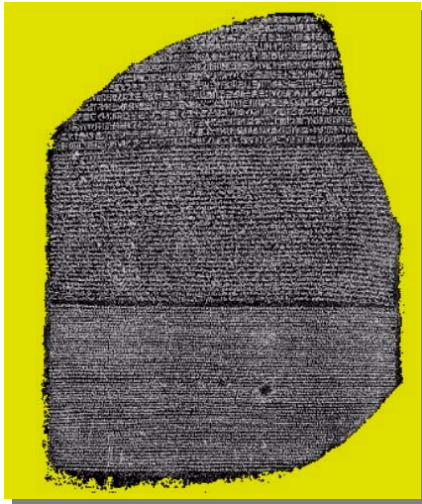
Local environment: future timeline



Redfill 1998

ROSETTA and PHILAE

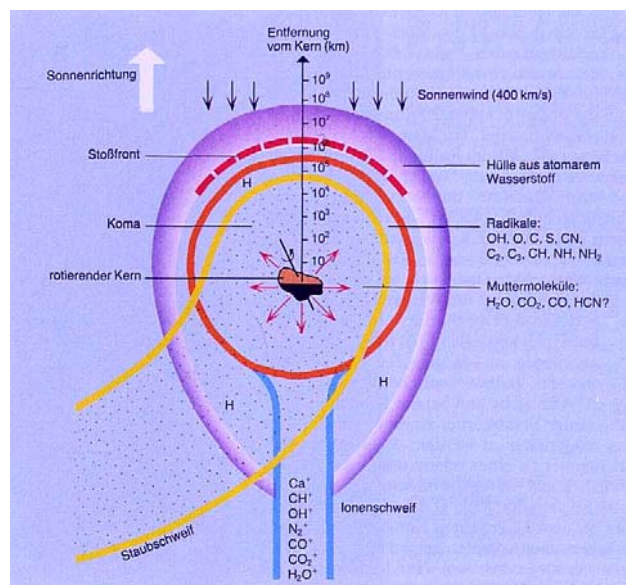
ROSETTA stone



Island Philae, column and inscription



Comets – greatest and tiniest cheaters in the solar system

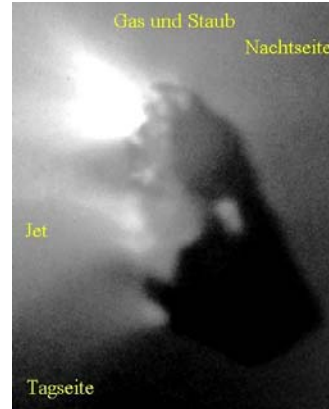




Comet Halley (March 13, 1986)

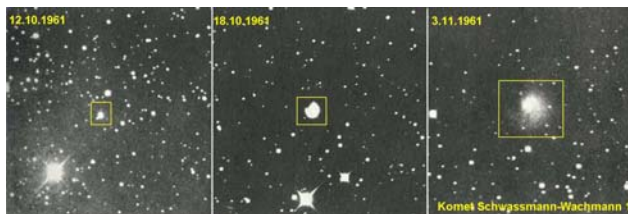
Famous comet: Halley

nucleus, gas and dust 1986



„dirty snowbal“ ?

Edmond Halley, painted by Thomas Murray



Schwassmann-Wachmann 1, outburst beyond Jupiter's orbit

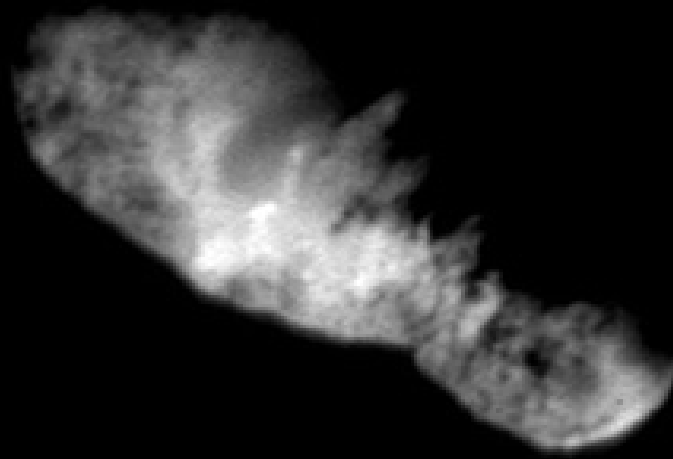
Comets



Comet West, 1976

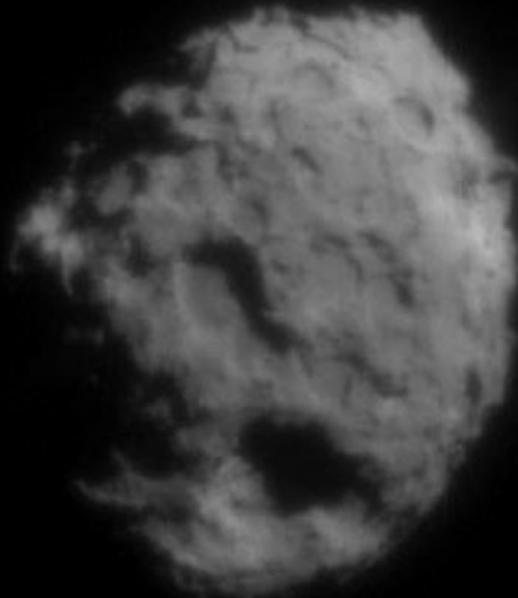


Cometary nucleus 19 P/Borely



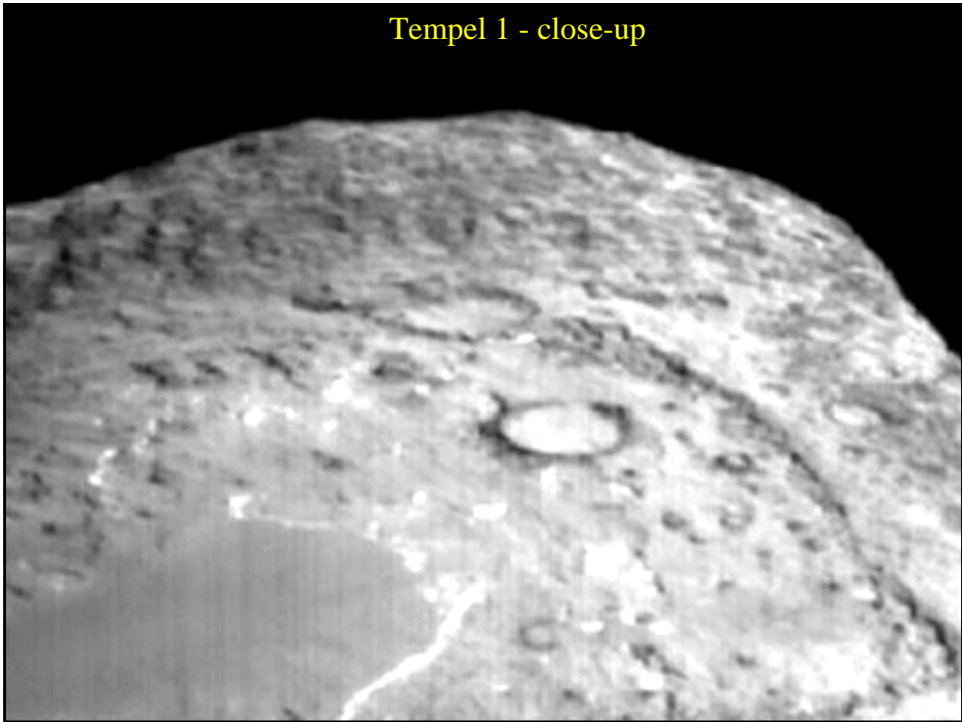
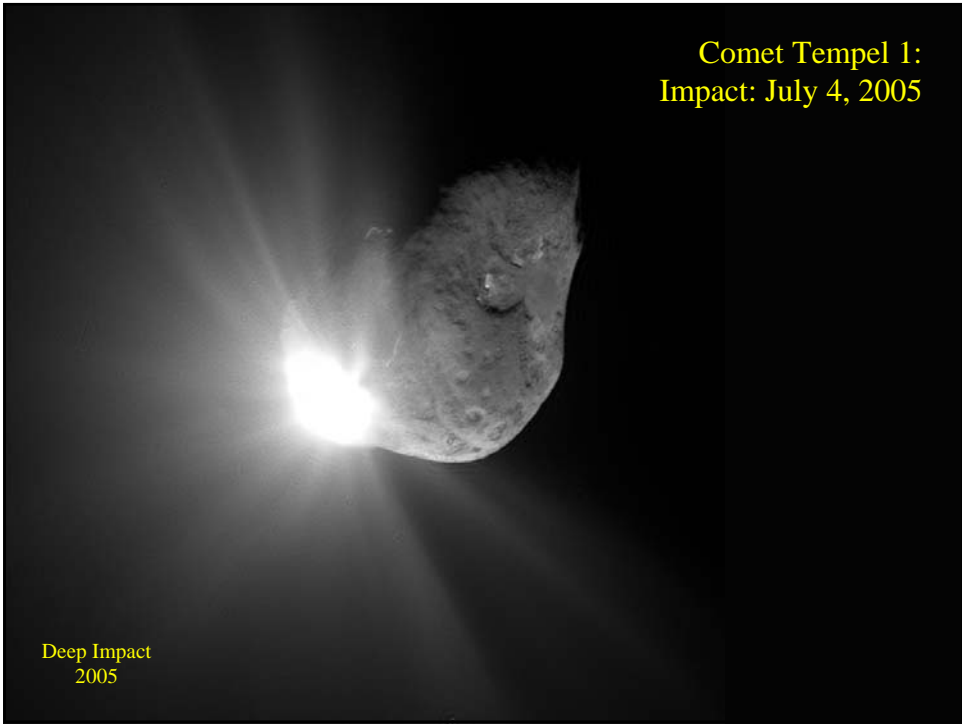
Deep space 1
2001

Cometary nucleus 81 P/Wild-2



Stardust
2004

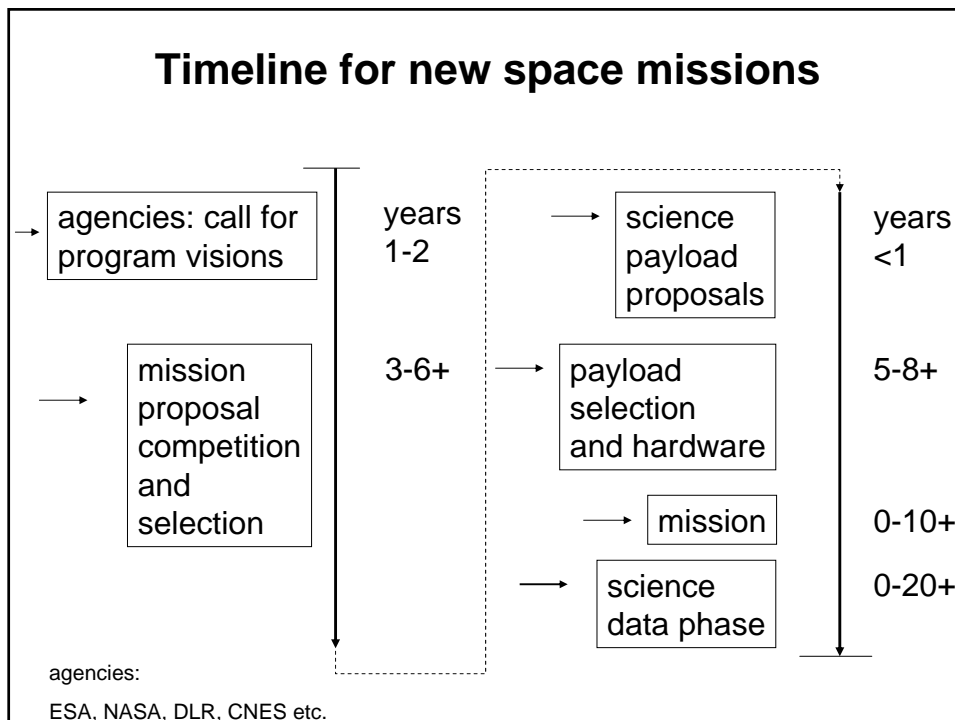
~ 5 km



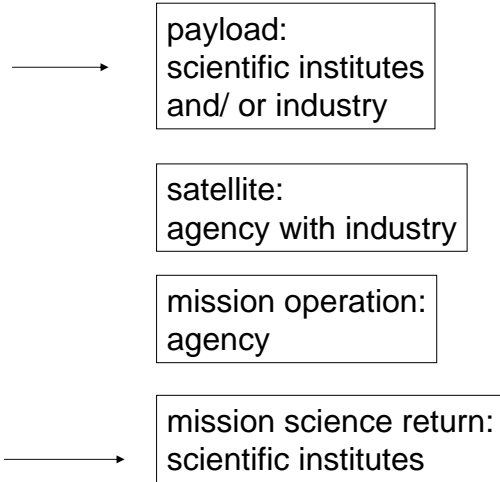
Outline

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Timeline for new space missions



Space missions: General management approach



Design Depends on Individual Who Defines Problem

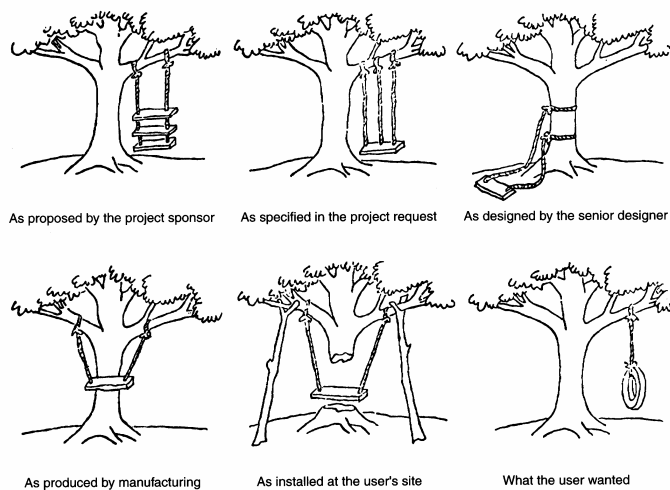


FIGURE 1.4

Note how the design depends on the viewpoint of the individual who defines the problem.

Outline

- timeline, matter and comets
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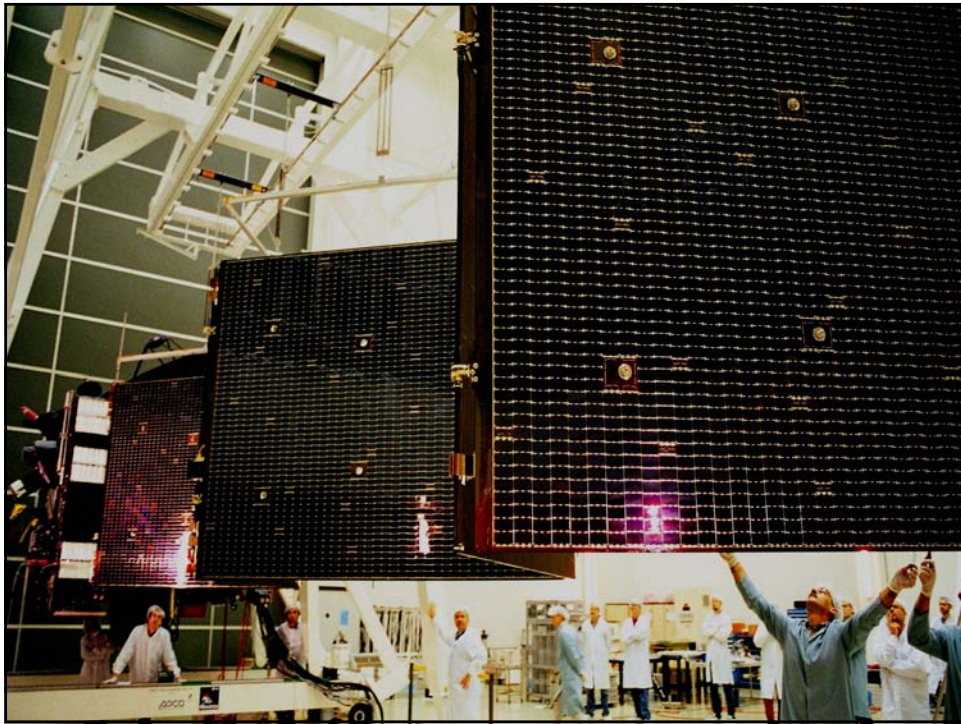
Science goals of Rosetta mission

- a global characterization of the nucleus,
- the determination of its dynamic properties,
- the surface morphology and composition,
- the determination of chemical, mineralogical and isotopic compositions of volatiles and refractories in the cometary nucleus,
- the determination of the physical properties and interrelation of volatiles and refractories in the cometary nucleus,
- studies of the development of cometary activity and the processes in the surface layer of the nucleus and inner coma, that is dust/gas interaction,
- studies of the evolution of the interaction region of the solar wind and the outgassing comet during perihelion approach.



Table 2 Spacecraft Properties

Size: main structure	2.8 x 2.1 x 2.0 m ³
Span of solar arrays	32 m
Launch mass: - total	2900 kg
- propellant	1720 kg
- science payload	165 kg
- lander PHILAE	100 kg
Solar array output	850 W at 3.40 AU 395 W at 5.25 AU
Propulsion subsystem	24 bipropellant 10 N thrusters
Operational life time	12 years
Prime contractor	Alenia Spazio, Italy



Rosetta and Rosetta Lander Philae



Rosetta (solar simulation test)

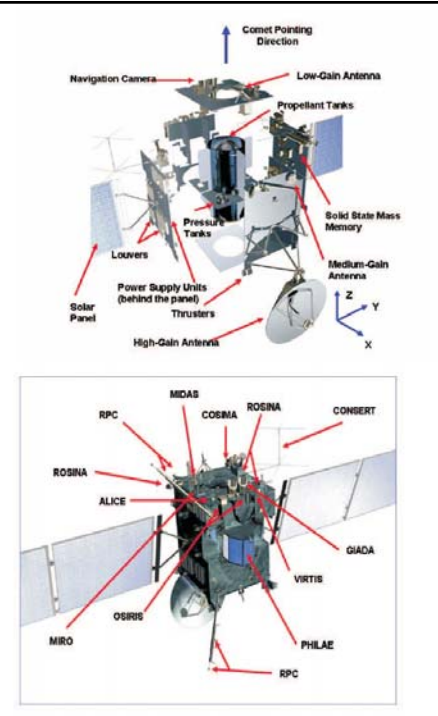


Rosetta Lander Philae attached to orbiter



Table 3 The Payload of the ROSETTA Mission

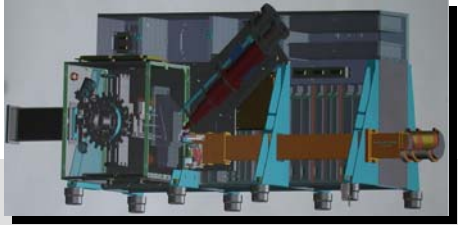
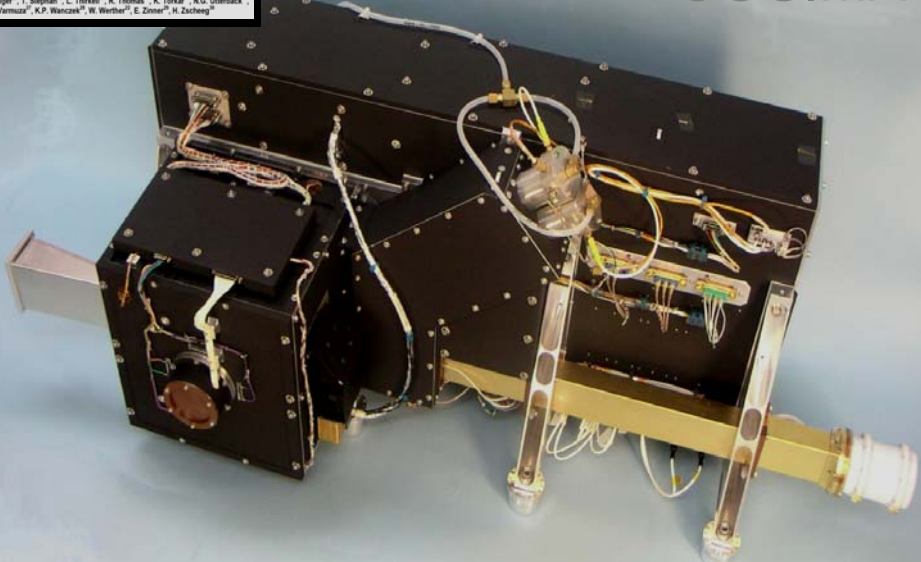
Instrument Name	Scientific Objectives	Principal Investigator
OSIRIS	Multi-Colour Imaging with a Narrow and a Wide Angle Camera	Horst-Uwe Keller MPS, Lindau Germany
ALICE	UV-Spectroscopy (70 nm - 205 nm)	Alan Stern SRI, Boulder, USA
VIRTIS	VIS and IR Mapping Spectroscopy (0.25 - 5 μm)	Angioletta Coradini IAS-CNR, Rome, Italy
MIRO	Microwave Spectroscopy (1.3 mm and 0.5 mm)	Sam Gulikis JPL, Pasadena, USA
ROSINA	Neutral Gas and Ion Mass Spectroscopy DFMS: 12-200 AMU $M/\Delta M \approx 3000$ RTOF: 12-350 AMU $M/\Delta M > 1000$	Hans Balsiger Universität Bern Switzerland
COSIMA	incl. Gas Pressure Sensor Dust Mass Spectrometer (SIMS, $m/\mu\text{m} \approx 2000$)	Martin Hilchenbach (formerly Jochen Kissel) MPS, Lindau Germany
MIDAS	Grain Morphology with an Atomic Force Microscope at nm Resolution	Willi Riedler IWF, Graz, Austria
CONSERT	Radio Sounding and Nucleus Tomography	Wlodek Kofman CEPHAH, Grenoble France
GIADA	Dust Velocity and Impact Momentum Measurement, Contamination Monitor	Luigi Colangelo INAF, Naples, Italy
RPC	Langmuir Probe (LAP) Ion and Electron Sensor (IES) Flux Gate Magnetometer (MAG) Ion Composition Analyser (ICA) Mutual Impedance Probe (MIP) Plasma Interface Unit (PIU)	Anders Eriksson (formerly Rolf Boström) IRF Uppsala, Sweden Jim Burch SRI, San Antonio, USA Karl-Heinz Glassmeier IGEP, TU Braunschweig Germany Rickard Lundin IRF, Kiruna, Sweden Jean-Gabriele Trotignon LPCE/CNRS, Orleans France Chris Carr Imperial College, London England Martin Pätzold Universität Köln Germany
RSI	Radio Science Experiment	



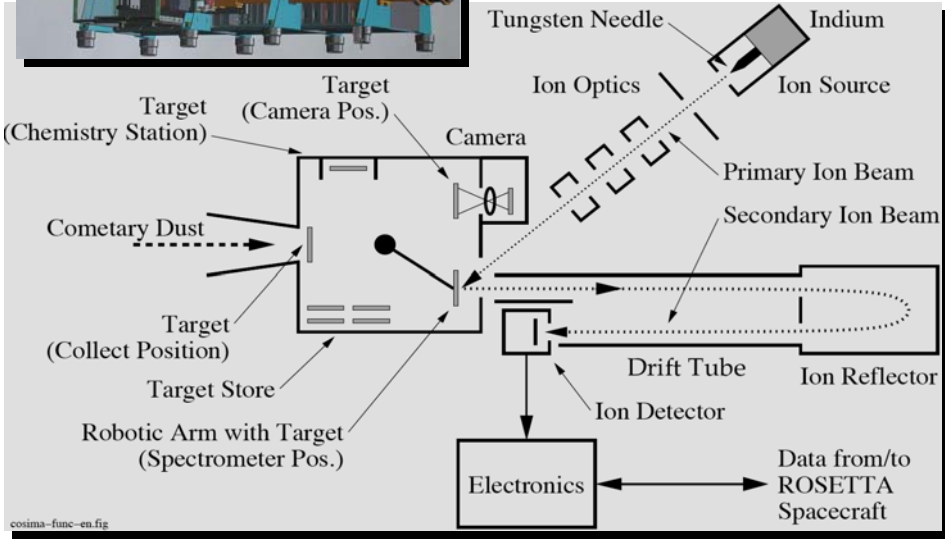
COSIMA, a High Resolution Time-of-Flight Secondary Ion Mass Spectrometer for the Analysis of Cometary Dust Particles

J. Kissel¹, K. Altwegg², B.C. Clark³, L. Colangelo⁴, H. Cottin⁵, S. Czopiel⁶, J. Ehl⁷, G. Engelund⁸, H.M. Fehringer⁹, B. Feuerbacher¹⁰, M. Fomenkova¹¹, A. Glasmeier¹², J. M. Greenberg¹³, E. Guzmán¹⁴, G. Harshbarger¹⁵, H. Hanel¹⁶, M. Hübner¹⁷, H. von Hoerner¹⁸, H. Höfner¹⁹, K. Hornung²⁰, E.K. Jessberger²¹, A. Koch²², F.R. Krueger²³, H. Krüger²⁴, G. Kurat²⁵, Y. Langevin²⁶, P. Parigger²⁷, F. Raulin²⁸, F. Rodemann²⁹, J. Ryan³⁰, E. R. Schmitt³¹, R. Schulz³², J. Silen³³, W. Singer³⁴, T. Stephan³⁵, L. Tosi³⁶, R. Thomas³⁷, K. Torok³⁸, K.G. Usterhus³⁹, K. Varnum⁴⁰, K.P. Wanczak⁴¹, W. Wierher⁴², E. Zinner⁴³, H. Zeeberg⁴⁴

Rosetta COSIMA



COSIMA- functional diagram



Stardust (Comet 81P/Wild-2): In-situ and (future) laboratory measurements

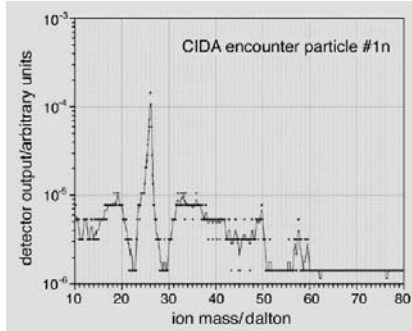
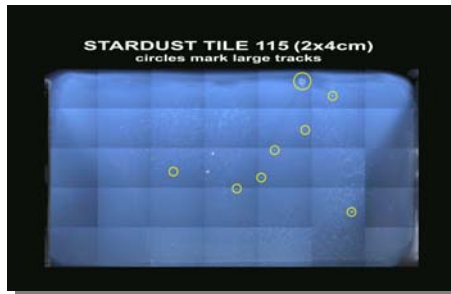


Fig. 2. A negative-ion spectrum, converted like Fig. 1. The dominating ion is CN^- , as is typical for nitrogen organic chemistry. The oxygen region ($m/z = 16, 17$) is surprisingly low; however, the SH^- ($m/z = 33, 35$) is high. The shift of the H^- line is a known instrumental effect.

Kissel et al 2004



cometary dust,
captured in aerogel,
(sample return
Jan 2006)

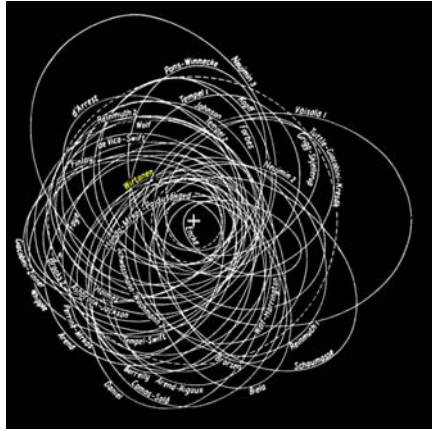


Tsou 2006

Selection of target comet for the Rosetta mission



Selection of target comet for the Rosetta mission - II



2003

conditions:
 periodic comet, „new“ in the inner solar system and
 target comet should not fall apart*, and rocket should be
 ready for launch in time....

*unpredictable...

New target comet for Rosetta in 2003 :

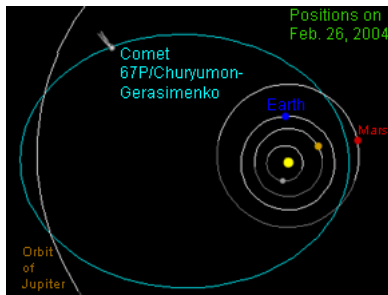
67P/Churyumov-Gerasimenko

size	=	3 x 5 km
perihelion	=	1.292 AU
aphelion	=	5.730 AU



67P/Churyumov-Gerasimenko
 February 1st 2003

Approaching comet 67P/Churyumov-Gerasimenko



Journey to comet (2004-2014)



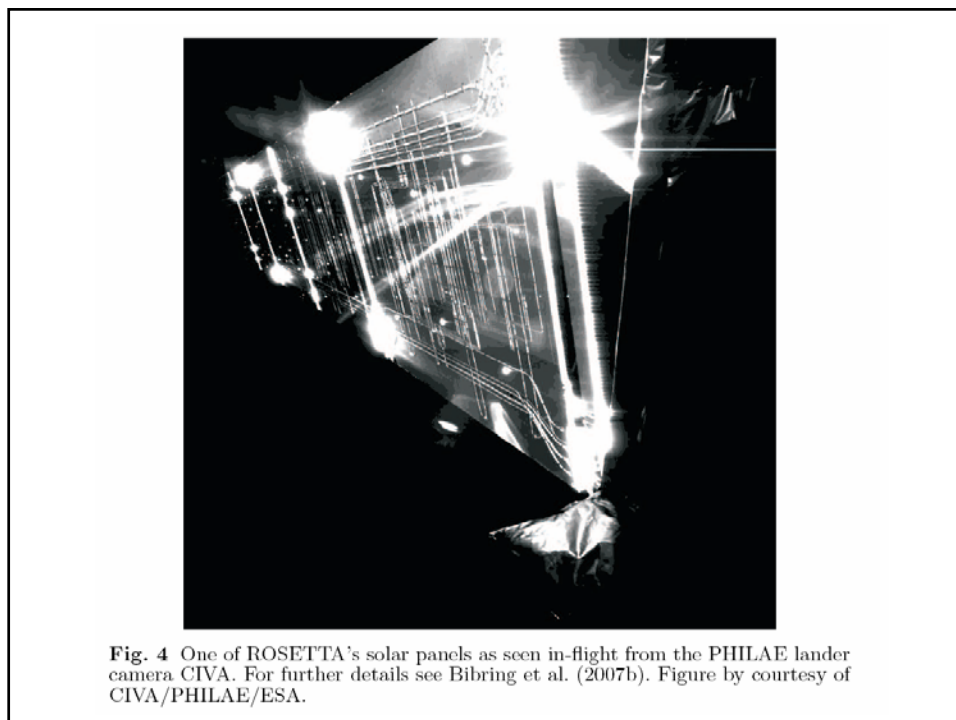
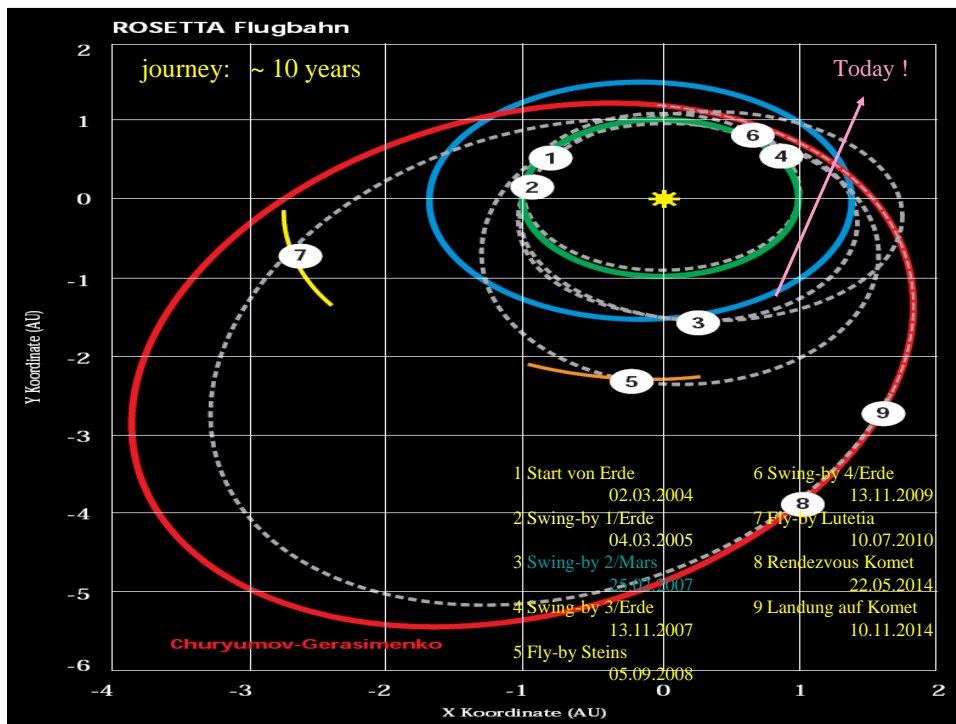
Launch of
Ariane V
2004



Science at comet
(2014/2015: 12 month)

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Rosetta
Osiris
observations
of Tempel 1
„Deep Impact“

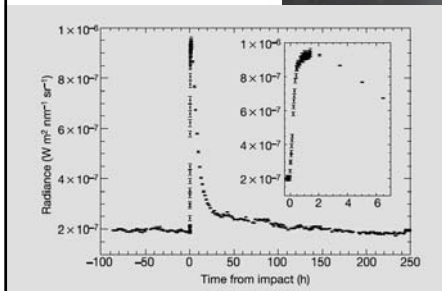


Figure 3 | Light curve of the cometary dust in the orange filter. The data

Küppers et al. 2006

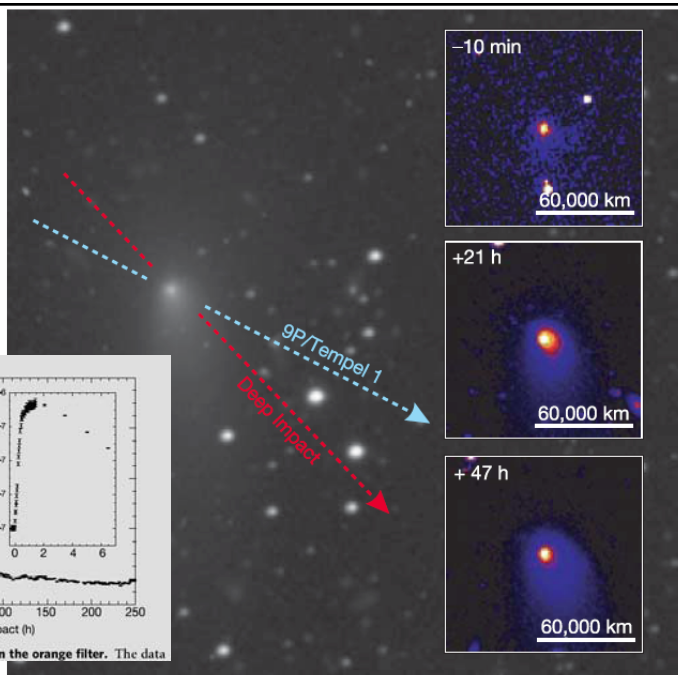


Figure 1 | The Comet 9P/Tempel 1 as seen by the Osiris NAC. The NAC

Outlook and milestones towards comet 67P/Churyumov-Gerasimenko

Table 1 Milestones of the ROSETTA Mission

Mission Event	Nominal Date
Launch	March 2, 2004
First Earth Gravity Assist	March 4, 2005
Mars Gravity Assist	February 25, 2007
Second Earth Gravity Assist	November 13, 2007
2867 Steins Flyby	September 5, 2008
Third Earth Gravity Assist	November 13, 2009
21 Lutetia Flyby	July 10, 2010
Rendezvous Manoeuvre 1 and Start of Hibernation	January 23, 2011
Exit Hibernation	January, 2014
Comet Rendezvous	Spring, 2014
Maneuver Between 4.5 and 4.0 AU	
Start of Near-Nucleus Operations at 3.25 AU	August 22, 2014
PHILAE Deployment	November, 2015
Perihelion Passage	August, 2015
End of Nominal Mission	December 31, 2015

